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MOST PROBABLE MASSES OF ISOTOPIC DISTRIBUTIONS
AND SYSTEMATICS OF PHOTOSPALLATION YIELDS

Ву

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## ABSTRACT

A Z-dependence of the most probable mass number  $A_{\mathrm{mp}}$  for isotopic distributions of spallation residuals is proposed. The modified five-parameter semiempirical formula thus deduced seems to reproduce fairly well most of the experimental data of photospallation yields.

Key-words: Photospallation; Spallation; Isotopic distributions;
Photo nuclear reactions.

Charge-Distribution-Mass-Distributions (isotopic distributions) of spallation residuals are described well enough by multiparametre semiempirical formulae ( $^{1-18}$ ) that, without any need of expensive and time consuming computer calculations ( $^{10,20}$ ), have been proved to be suitable in reproducing the cross sections (or yields) of spallation residuals as a function of the mass- and atomic number ( $A_t, Z_t$ ) of the target element and those of the product ( $A_i, Z_i$ ), and of the nominal nucleonand proton loss ( $\Delta A_i, \Delta Z_i$ ).

Almost all these formulae contain, as a <u>free parametre</u>, the most probable ( $\underline{mp}$ ) mass,  $A_{\underline{mp}}$ , of a given isotopic distribution on the mass-yield plane.

Due to the fact that  $A_{mp}(Z)$  appears in the argument of a gaussian (or almost-gaussian) function,  $A_{mp}$  becomes an extremely critical parametre.

The aim of the present paper was to search for an analytical form of  $A_{mp}(Z)$  that could give better results and cover much wider range of atomic numbers than do other formulae, especially in the case of photon-initiated spallation at intermediate energies (up to a few GeV).

Regardless of the experimental irradiation conditions,

Amp has always been shown to assume values which differ very

little from the A values of the most abundant naturally occurring

isotopes (with only a very few exceptions, within 0.5 to 1 A units)

for any peculiar distribution (at fixed Z).

More precisely,  $A_{mp}$  seems to be very close to the average A value  $(A_{sw})$  of the <u>n</u> existing stable masses for a given Z, <u>weighted</u> over the percent natural isotopic abundance w

$$(\sum_{j=1}^{n} w_{j} = 1)$$
.  
For each  $Z_{i}$ ,  $A_{mp}(Z_{i})$  is consequently written as

(1) 
$$A_{mp}(z_i) \approx A_{sw}(z_i) = \sum_{j=1}^{n} [w_j A_{sj}(z_i)]$$
,

s standing for "stable", even if one or more Asj are radioactive in nature.

Following the hypothesis analitically expressed by eq. (1),  $A_{sw}$  should represente <u>centre of gravity</u> of the <u>n</u>  $A_{sj}$  masses, and  $A_{mp}(Z_i)$  the top of the distribution of isotopes having atomic number  $Z_i$ .

A careful regression analysis, with some data rejection, based on eq. (1) gave, from  $Z=6\,$  up to Z=83, the following Z-dependence

(2) 
$$A_{sw(fit)} = \alpha Z^{\beta}$$
,

with

(3) 
$$\begin{cases} \alpha = 1.590 \\ \beta = 1.103 \end{cases}$$

and a coefficient of correlation:  $r^2 = 0.985$ .

To verify the goodness of the fitting procedure and the homogeneity of the two-sample data  $(A_{sw}, A_{sw}(fit))$  both parametric statistics were used.

Statistical Goodness of Fit (GOF) tests showed strong evidence as for homogeneity and correlation of the two samples were concerned (e.g.,  $R_{xy} \approx 1$ , Spearman's  $\rho = 0.999$ , and reduced  $\chi^2_{\nu=75} = 1.1$  as deduced from the Bartlett's criterion with a

confidence interval of 90%).

A Student t-test (with  $\nu$  = 74 degrees of freedom and a probability of 0.975) was employed to obtain the errors effecting  $\alpha$  and  $\beta$ 

(4) 
$$\alpha = 1.590 \pm 0.010$$

and

(5) 
$$\beta = 1.103 \pm 0.009$$

From eqs. (1-5) we can write

(6) 
$$A_{mp} = A_{fit} = (1.590 \pm 0.010) z^{(1.103\pm0.009)}$$

(to avoid excess of symbols Afit takes the place of Asw, (fit)).

Table I shows a comparison between  $A_{sw}$  and  $A_{fit}$  (\*  $A_{mp}$ ) for 13 values of Z (plus other two values out of systematics).

Plotted in Fig. 1 are the values of  $A_{\mbox{fit}}(Z)$  (filled squares, from eq. (2)) and those of all naturally occurring isotopes  $A_{\mbox{s}}$  (open squares) \*.

Following the hypotheses which eqs. (1) and (2) are based on, the filled squares of Fig. 1 should represent the bottom of the stability valley.

It is now very interesting to consider some of the  $A_{\rm mp}$  dependences on Z proposed in ( $^{1-18,21}$ ) and in the present paper (eq. 2) in order to see how well they match the most abundant  $A_{\rm S}(Z_{\rm i})$ .

There are, from  ${}^{12}_{6}$ C to  ${}^{209}_{33}$ Bi, 274 natural elements for 76 Z values. About 15% of such element are radioactive, nevertheless they were included in the course of the analysis and in Fig.1, for they contribute to some extent to  $A_{SW}$ .

In doing this, we chose the R test, that is

(7) 
$$R = \exp(\varepsilon)$$
,

with

(8) 
$$\varepsilon = \left\{ \frac{1}{m+n} \sum_{k=1}^{m} \sum_{i=1}^{n} \left[ \ln (A_{s,ik}/A_{mp,ik}) \right]^{2} \right\}^{1/2} =$$

$$= \left\{ \frac{1}{274} \sum_{k=1}^{m} \sum_{j=1}^{n} \left[ \ln (A_{s,jk}/A_{fit,jk}) \right]^{2} \right\}^{1/2}$$

for grouped sets of m Z values (see later in Table II).

The quantity R takes, of course, values greater than unity. The greater the R value is, the larger is the difference  $\ln A_{s,ik} - \ln A_{mp,ik}$  (or  $A_{s,ik} - A_{mp,ik}$ ) and R allows a suitable test to be made in order to compare different  $A_{mp}$  dependences with each other.

Table II lists the results of such a comparison. More-over, Fig. 2 shows the histogram of frequency distribution of  $A_{mp}-A_s \quad \text{for the whole set of 274 A}_s \quad \text{values (6 } \le \text{ Z } \le \text{ 83). A reduced}$   $x^2 = 1.2 \text{ was calculated.}$ 

At this point we are likely to analyse some expected changes of the values of parametres in our previous cross section formulae ( $^{4,5,11}$ ), due to the introduction of the present analytical form of  $A_{mo}$ .

As reference cross section formula for isotopic distribution we chose that reported as eq. (15) in Ref.  $(^{11})$ , which is written as

(9) 
$$\overline{\sigma}_{i} = a\overline{\sigma}_{N}K_{A} \exp \left[-b(A-A_{mp}(Z_{i}))^{2}\right]$$

with

$$A_{mp} = cZ_i - d$$

and

$$\begin{cases}
a = 3.30 \pm 0.08 \\
\overline{\sigma}_{N} = 260 \text{ } \mu \text{b} \\
K_{A} = (1.67 \pm 0.03) A_{t}^{-}(0.0580 \pm 0.0010) \\
b = 0.243 \pm 0.005 \\
c = 2.28 \pm 0.07 \\
d = 2.18 \pm 0.09
\end{cases}$$

Cross section formula (9) gives the mean cross section  $\frac{\text{per}}{\sigma_N}$  photon (energy range from about 0.2 GeV to about 1 GeV;  $\overline{\sigma}_N = 260~\mu \text{b}$  for 0.3 GeV - 1. GeV range) of a spallation residual  $(A_1, Z_1)$  from a  $(A_1, Z_1)$  target. It contains the five parameters:  $\underline{a}$  (a normalisation factor),  $K_A$  (slope of the yield-surface ridge  $(^{1}, ^{16-18})$  for isotopic distributions), b (which is related to the full-width-at-half-maximum  $\Gamma = 2(\ln 2/6)^{1/2}$ ), and c and d (wich define  $A_{mp}$ ). The  $\overline{\sigma}_N$  is the mean cross section  $\underline{\text{per}}$  photon of the elementary gamma-nucleon interaction  $(^{4}, ^{5})$ .

If one substitutes  $A_{mp}$  of eq. (2) in eq. (9), the following equation is obtained for the natural log of the normalised cross-section (4,5)  $\overline{\sigma}_i^*$ 

(11) 
$$\ln \overline{\sigma}_{i}^{*} = \ln [\overline{\sigma}_{i}/(K_{A}^{-\Delta Z_{i}} \times \overline{\sigma}_{N})] = \ln (\zeta a) - \eta b (A_{i} - 1.590Z_{i}^{1.103})^{2},$$

with  $\zeta$  and  $\eta$  factors accounting for the changes in the parametres  $\underline{a}$  and  $\underline{b}$  (no change in  $K_{\underline{A}}$  was espected, for  $K_{\underline{A}}$  is much less sensitive to changes in  $A_{\underline{mp}}$ ).

A new regression analysis was thus carried out with a

set of 112 measured mean cross section <u>per</u> photon for spallation residuals from V, Mn, Fe, and Co targets  $(^{22,23})$ , and for  $^{23}$ Na and  $^{24}$ Na photoproduction from various targets  $(^{4,5})$ .

The regression line gave

(12) 
$$\zeta a = 3.89 \pm 0.11$$

and

(13) 
$$\eta b = 0.247 \pm 0.007$$

(coefficient of determination  $r^2 = 0.92$ ; Spearman's  $\rho = -0.925$ ; reduced  $\chi^2 = 1.71$ ).

The calculated trend of  $\overline{\sigma}^*$  from eqs. (11), (12), and (13) is plotted in Fig. 3 (full line marked with  $\underline{a}$ ). Fig. 3 also reports  $\overline{\sigma}^*$  from (<sup>9,10</sup>) (full line  $\underline{b}$ ) and the experimentally determined  $\overline{\sigma}^*$  values.

Reduced  $\chi^2$  of 1.5 and 2.7 were obtained from statistical treatments of experimental data calculated ones from eq. (11) and eq. (9), respectively.

From eq. (11), eq. (12) and eq. (13),  $\overline{\sigma}_i$  is written as

(14) 
$$\overline{\sigma}_{i} = 3.89 \, \overline{\sigma}_{N} \, K_{A}^{-\Delta z_{i}} \exp \left[-0.247 \, (A_{i}-1.590 \, z_{i}^{1.103})^{2}\right]$$
.

We tested eq. (14) further on by a set of cross sections  $\underline{per}$  equivalent quantum (bremsstrahlung end-point energy  $E_0$ =1GeV) of gold isotopes from a  $^{209}$ Bi target ( $^{24,25}$ ). Shown in Fig. 4 is the trend of eq. (14), modified for bremsstrahlung irradiations, and the experimental yields (note that the best-fit curve reaches its maximum a A = A<sub>g</sub> = 197).

In concluding this note, we wish to put some stress on

### what follows:

- i) eqs. (2) and (6) reproduce fairly well the  $A_{sw}$  and  $A_{mp}$  values (see Table I and Figs. 1 and 2);
- priate one to give a Z-dependence of A<sub>mp</sub> (from Table II it is readily seen that eq. (2) gives the best result, also in consideration of the fact that it covers the whole range of Z; although very good results are also encountered for (12), especially for the ICSD-not-gaussian Rudstam's formula (12), the Z range is smaller than that covered by eq. (2));
- iii) eq. (14) seems to reproduce experimental spallation yields with a fair approximation (reduced  $x^2 = 1.5$ ; see also Figs. 3 and 4);
  - iv) all the gof tests furnish clear evidence for paired samples homogeneity and correlation.

The work is presently being carried on for isobaric distributions and charge-dispersion curves of photospallation residuals, also at photon energies above 1 GeV.

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## FIGURE CAPTIONS

- Fig. 1 Trend of N (number of neutrons = A-Z)  $\underline{vs}$  Z.

  Open squares:  $N_s(Z)$  for naturally occurring isotopes (mass  $A_s$ ).

  Filled squares:  $N_{mp}(Z)$  (i.e.  $A_{mp}-Z$ )  $\underline{vs}$  Z from eq. (2).

  For Z=43 ( $T_c$ ) and Z=61 ( $p_m$ ), and for 83 < Z < 90, calculated values have only been plotted (the same for Z > 92). The dashed lines at Z=83 (Bi) indicate the upper limit of validity of eq. (2).
- Fig. 2  $A_{mp}$ - $A_{s}$  frequency distribution (histogram) and best-fit gaussian frequency distribution (dashed curve). Histogram data: mode = 0.00 (dashed straight line), median = 0.42, mean = 0.56,  $\Gamma$  = 6, and area = 274 (a total of 274 elements having been considered). For the gaussian-shaped curve:  $\mu$  = 0.533 and  $\Gamma$  = 5.83.
- Fig. 3 Normalised cross section  $\overline{\sigma}^*$  vs  $A_s A_{mp}$ . Curve a represents  $\overline{\sigma}_i^* = (\overline{\sigma}_i/K_A^{-1} \overline{\sigma}_N a) = \zeta \exp[-0.247(A_{si} A_{mp,i})^2]$ , with  $A_{mp,i} = 1.590 z_i^{1.103}$  and  $\zeta = 1.18$  (see text). Curve b represents  $\overline{\sigma}_i^* = (\overline{\sigma}_i/K_A^{-1} \overline{\sigma}_w a) = \exp[-0.243(A_{si} A_{mp,i})^2]$ , with  $A_{mp,i}(z_i) = 2.28 z_i 2.18$ . Experimental points from Ref. (22) (open circles) and Ref. (23,4,5) (filled circles). As for the FWHM,  $\Gamma_a = 3.35$  and  $\Gamma_b = 3.38$  were obtained.
- Fig. 4 Yield of gold isotopes from  $^{209}$ Bi at E $_0$  = 1 GeV. Open circles: data of Ref. ( $^{24}$ ). Filled circles: data of Ref. ( $^{25}$ ). For  $^{195}$ Au (filled circle in parentheses)

some uncertainties in the decay scheme and difficulties in radioactivity measurments were met. The curve represent Eq. (14).

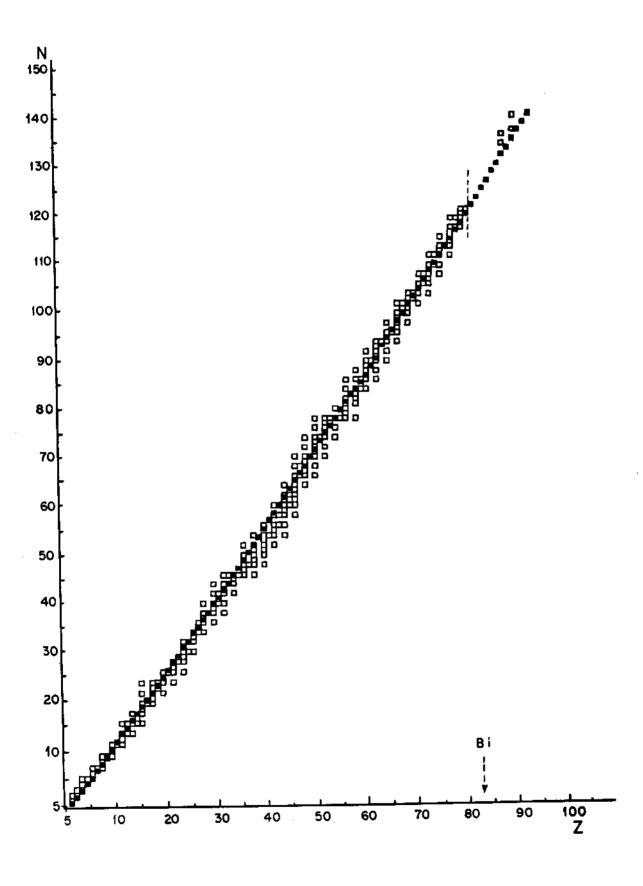


Fig. 1

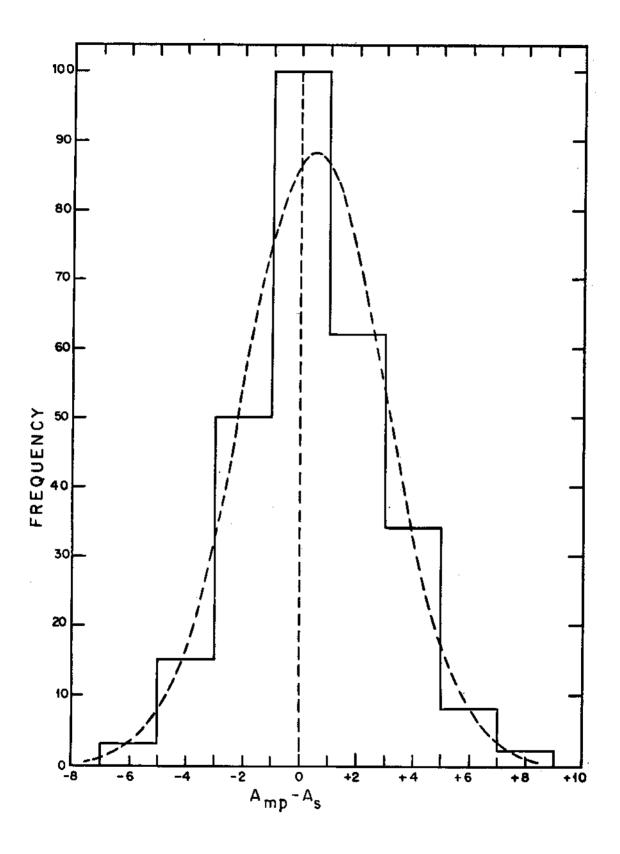


Fig. 2

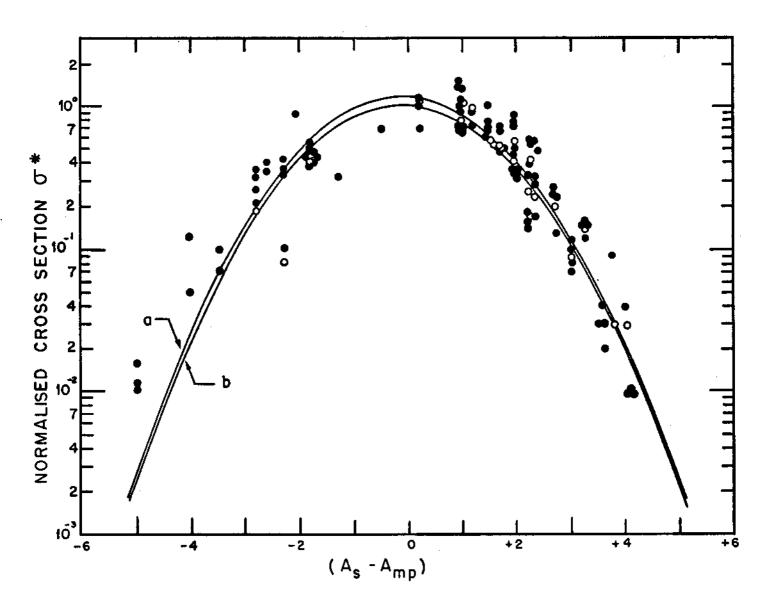


Fig. 3

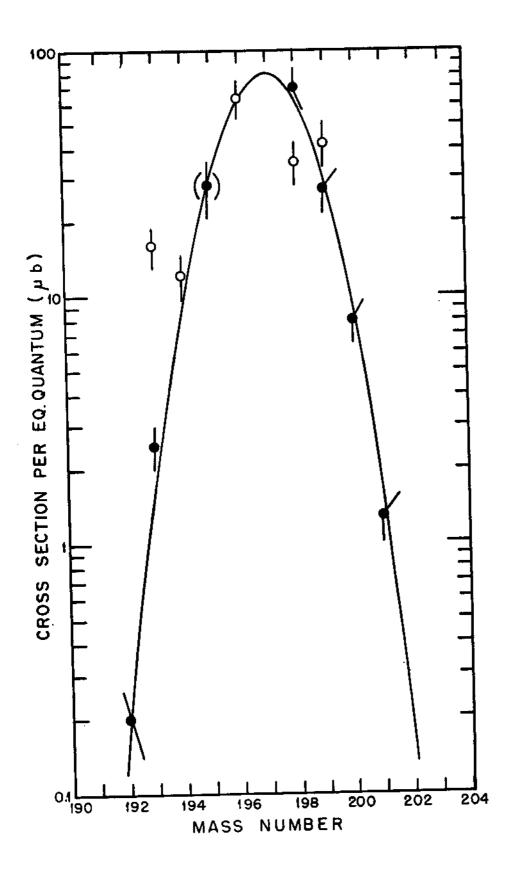


Fig. 4

TABLE I. - Comparison between  $A_{sw}$  and  $A_{fit}$ .

ELEMENT	Z	Asw	A <sub>fit</sub>	∇୫ (*)	Spread (eq.6)			
		Eq. (1)	Eq. (2)		A <sub>fit,min</sub>	A <sub>fit,max</sub>		
С	6	12.01	11.47	+4.5	11.26			
0	8	16.00	15.76	+1.5	15.43	16.09		
Ne	10	20.18	20.16	+0.1	19.71	20.61		
Al	13	27.00	26.92	+0.3	26.27	27.58		
Cl	17	35.49	36.19	-2.0	35.26	37.15		
Ar	18	39.99	38.54	+3.6	37.53	39.58		
Sc	21	45.00	45.69	-1.3	44.44	46.97		
V	23	51.00	50.51	+1.0	49.10	51.96		
As	33	75.00	75.22	-0.3	72.94	77 <b>.</b> 5 <b>7</b>		
I	53	127.00	126.85	+0.1	122.59	131.24		
Sm	62	150.35	150.80	-0.3	145.59	156.20		
Au	79	197.00	197.01	(-)0.0	189.87	204.40		
Bi	83	209.00	208.04	+0.5	200.43	215.92		
Th (**)	90	232.00	227.47	+2.0	219.03	236.23		
บ(**)	92	237.97	233.05	+2.1	224.37	242.06		

 $<sup>(*)^{\</sup>Lambda } = [(A_{sw} - A_{fit})/A_{sw}] \times 100$ 

<sup>(\*\*)</sup>Out of systematics.

TABLE II. - R-values for different Amp formulae.

Cumulative Z ranges	83 62-92 (*) 6-40 6-60 6-83 6-92 (**)	013 1.013 1.038 1.032 1.030 1.029 21, 4, 5	1.034 1.047 4, 5	- 1.038 1.045 9,5	1.035 1.045 4,	1.038 1.048 12	1.038 1.047 12	1.045 1.047 12	1.033 1.037 12	((C) and) there are a coop a coop of the c
Partial 2 ranges	83	1.013	1		1	ı	ı	1	1	1.006
	41-60	1.019	1.005	1.057	1.060	1.063	1.061	1.049	1.044	1.017
	21-40	1.023	1.028	1.025	1.026	1.037	1.034	1.032	1.031	1.023
	6-20	1.052	1.041	1.051	1.044	1.040	1.044	1.059	1.035	1.038

(\*) Same Z values as in column 4, plus Z=90 and Z=92 (out of systematics). (\*\*) Same Z values as in column 8, plus Z=90 and Z=92 (out of systematics).

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