# Chern-Simons $A d S_{5}$ supergravity in a Randall-Sundrum background 

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#### Abstract

Chern-Simons AdS supergravity theories are gauge theories for the super-AdS group. These theories possess a fermionic symmetry which differs from standard supersymmetry. In this paper, we study five-dimensional Chern-Simons AdS supergravity in a Randall-Sundrum scenario with two Minkowski 3-branes. After making modifications to the $D=5$ Chern-Simons AdS supergravity action and fermionic symmetry transformations, we obtain a $\mathbb{Z}_{2}$-invariant total action $S=\tilde{S}_{\text {bulk }}+S_{\text {brane }}$ and fermionic transformations $\tilde{\delta}_{\epsilon}$. While $\tilde{\delta}_{\epsilon} \tilde{S}_{\text {bulk }}=0$, the fermionic symmetry is broken by $S_{\text {brane }}$. Our total action reduces to the original Randall-Sundrum model when $\tilde{S}_{\text {bulk }}$ is restricted to its gravitational sector. We solve the Killing spinor equations for a bosonic configuration with vanishing $s u(N)$ and $u(1)$ gauge fields.


[^0]
## 1 Introduction

Chern-Simons AdS supergravity [1, 2, 3] theories can be constructed only in odd spacetime dimensions. As the name implies, they are gauge theories for supersymmetric extensions of the AdS group. ${ }^{1}$ They have a fiber bundle structure and hence are potentially renormalizable [2]. The dynamical fields form a single adS superalgebra-valued connection and hence the supersymmetry algebra closes automatically off-shell without requiring auxiliary fields [4]. The Lagrangian in dimension $D=2 n-1$ is a Chern-Simons ( $2 n-1$ )form for the super-adS connection and is a polynomial of order $n$ in the corresponding curvature. Unlike standard supergravity theories, there can be a mismatch between the number of bosonic and fermionic degrees of freedom. ${ }^{2}$ For this reason, the 'supersymmetry' of Chern-Simons AdS supergravity theories is perhaps better referred to as a fermionic symmetry.
$D=11, N=1$ Chern-Simons $A d S$ supergravity may correspond to an off-shell supergravity limit of M-theory [2,3]. It has expected features of M-theory which are not shared by $D=11$ Cremmer-Julia-Scherk (CJS) supergravity [5]. These features include an $\operatorname{osp}(32 \mid 1)$ superalgebra [6] and higher powers of curvature [7]. Hořava-Witten theory [8] is obtained from CJS supergravity by compactifying on an $S^{1} / \mathbb{Z}_{2}$ orbifold and requiring gauge and gravitational anomalies to cancel. This theory gives the low energy, strongly coupled limit of the heterotic $E_{8} \times E_{8}$ string theory. In light of the above discussion, it would be interesting to reformulate Hořava-Witten theory with $D=11, N=1$ Chern-Simons AdS supergravity.

Reformulating Hořava-Witten theory as described above may prove to be difficult. It is simpler to compactify the five-dimensional version of Chern-Simons AdS supergravity on an $S^{1} / \mathbb{Z}_{2}$ orbifold and ignore anomaly cancellation issues. Canonical sectors of $D=$ 5 Chern-Simons AdS supergravity have been investigated in locally $\operatorname{AdS} S_{5}$ backgrounds possessing a spatial boundary with topology $S^{1} \times S^{1} \times S^{1}$ located at infinity [9]. In this paper, as a preamble to reformulating Hořava-Witten theory, we will study $D=5$ Chern-Simons AdS supergravity in a Randall-Sundrum background with two Minkowski 3 -branes [10]. We choose coordinates $x^{\mu}=\left(x^{\bar{\mu}}, x^{5}\right)$ to parameterize the five-dimensional spacetime manifold. ${ }^{3}$ In terms of these coordinates, the background metric takes the form

$$
\begin{equation*}
g_{\mu \nu} d x^{\mu} d x^{\nu}=\mathfrak{a}^{2}\left(x^{5}\right) \eta_{\bar{\mu} \bar{\nu}}^{(4)} d x^{\bar{\mu}} d x^{\bar{\nu}}+\left(d x^{5}\right)^{2} \tag{1.1}
\end{equation*}
$$

where $\eta_{\bar{\mu} \bar{\nu}}^{(4)}=\operatorname{diag}(-1,1,1,1)_{\bar{\mu} \bar{\nu}}, \mathfrak{a}\left(x^{5}\right) \equiv \exp \left(-\left|x^{5}\right| / \ell\right)$ is the warp factor, and $\ell$ is the $A d S_{5}$ curvature radius. The coordinate $x^{5}$ parameterizes an $S^{1} / \mathbb{Z}_{2}$ orbifold, where the circle $S^{1}$ has radius $\rho$ and $\mathbb{Z}_{2}$ acts as $x^{5} \rightarrow-x^{5}$. We choose the range $-\pi \rho \leq x^{5} \leq \pi \rho$ with the endpoints identified as $x^{5} \simeq x^{5}+2 \pi \rho$. The Minkowski 3-branes are located at

[^1]the $\mathbb{Z}_{2}$ fixed points $x^{5}=0$ and $x^{5}=\pi \rho$. These 3-branes have corresponding tensions $\mathcal{T}^{(0)}$ and $\mathcal{T}^{(\pi \rho)}$ and may support $(3+1)$-dimensional field theories.

This paper is organized as follows: In Section 2, we construct a $\mathbb{Z}_{2}$-invariant bulk theory. This bulk theory is obtained by making modifications to the $D=5$ Chern-Simons AdS supergravity action and fermionic symmetry transformations which allow consistent orbifold conditions to be imposed. The variation of the resulting bulk action $S_{\text {bulk }}$ under the resulting fermionic transformations $\delta_{\epsilon}$ vanishes everywhere except at the $\mathbb{Z}_{2}$ fixed points. We calculate $\delta_{\epsilon} S_{b u l k}$ in Section 3. In Section 4, we modify $S_{b u l k}$ and $\delta_{\epsilon}$ to obtain a modified $\mathbb{Z}_{2}$-invariant bulk theory. The modified bulk action $\tilde{S}_{\text {bulk }}$ is invariant under the modified fermionic transformations $\tilde{\delta}_{\epsilon}$. In Section 5 , we complete our model by adding the brane action $S_{\text {brane }}$. We show in Section 6 that our total action

$$
\begin{equation*}
S=\tilde{S}_{\text {bulk }}+S_{\text {brane }} \tag{1.2}
\end{equation*}
$$

reduces to the original Randall-Sundrum model [10] when $\tilde{S}_{\text {bulk }}$ is restricted to its gravitational sector. In Section 7, we solve the Killing spinor equations for a purely bosonic configuration with vanishing $s u(N)$ and $u(1)$ gauge fields. Our concluding remarks are given in Section 8. Finally, in the Appendix, we work out the fünfbein, spin connection, curvature 2-form components, Ricci tensor, and Ricci scalar for our metric (1.1).

## $2 \mathbb{Z}_{2}$-invariant bulk theory

In this section, we construct a $\mathbb{Z}_{2}$-invariant bulk theory. The bulk theory is obtained by making modifications to the $D=5$ Chern-Simons AdS supergravity [1] action and fermionic symmetry transformations which allow consistent orbifold conditions to be imposed.

The field content of $D=5$ Chern-Simons AdS supergravity is the fünfbein $e_{\mu}{ }^{a}$, the spin connection $\omega_{\mu}^{a b}$, the $s u(N)$ gauge connection $A_{\mu}=A_{\mu}^{i} \tau_{i}$, the $u(1)$ gauge connection $B_{\mu}$, and $N$ complex gravitini $\psi_{\mu r}$ which transform as Dirac spinors in a vector representation of $\operatorname{su}(N) .{ }^{4}$ These fields form a connection for the adS superalgebra $s u(2,2 \mid N)$. The action and fermionic symmetry transformations are given in [9] in terms of the $A d S_{5}$ curvature radius $\ell$. The only free parameter in the action is a dimensionless constant $k$. To allow consistent $\mathbb{Z}_{2}$ orbifold conditions to be imposed, we make the following modifications:

1. Rescale the $s u(N)$ and $u(1)$ gauge connections:

$$
A \rightarrow g_{A} A, \quad B \rightarrow g_{B} B
$$

[^2]2. Replace $g_{A}, g_{B}, \ell^{-1}$, and $k$ by the $\mathbb{Z}_{2}$-odd expressions ${ }^{5}$
\[

$$
\begin{aligned}
& G_{A} \equiv g_{A} \operatorname{sgn}\left(x^{5}\right), \quad G_{B} \equiv g_{B} \operatorname{sgn}\left(x^{5}\right), \quad L^{-1} \equiv \ell^{-1} \operatorname{sgn}\left(x^{5}\right) \\
& K \equiv k \operatorname{sgn}\left(x^{5}\right)
\end{aligned}
$$
\]

In this manner, we obtain the bulk action

$$
\begin{equation*}
S_{\text {bulk }}=S_{\text {grav }}+S_{s u(N)}+S_{u(1)}+S_{\text {ferm }} \tag{2.1}
\end{equation*}
$$

where

$$
\begin{align*}
S_{\text {grav }}= & \int \frac{1}{8} K \varepsilon_{a b c d e}\left(\frac{1}{L} R^{a b} R^{c d} e^{e}+\frac{2}{3 L^{3}} R^{a b} e^{c} e^{d} e^{e}+\frac{1}{5 L^{5}} e^{a} e^{b} e^{c} e^{d} e^{e}\right), \\
S_{s u(N)}= & \int i K \operatorname{str}\left(G_{A}^{3} A F^{2}-\frac{1}{2} G_{A}^{4} A^{3} F+\frac{1}{10} G_{A}^{5} A^{5}\right), \\
S_{u(1)}= & \int K\left[-\left(\frac{1}{4^{2}}-\frac{1}{N^{2}}\right) G_{B}^{3} B(d B)^{2}+\frac{3}{4 L^{2}}\left(T^{a} T_{a}-\frac{L^{2}}{2} R^{a b} R_{a b}\right.\right. \\
& \left.\left.-R^{a b} e_{a} e_{b}\right) G_{B} B-\frac{3}{N} G_{A}^{2} G_{B} F_{s}^{r} F_{r}^{s} B\right], \\
S_{\text {ferm }}= & \int \frac{3}{2 i} K\left(\bar{\psi}_{\alpha}^{r} \mathcal{R}_{\beta}^{\alpha} \nabla \psi_{r}^{\beta}+\bar{\psi}_{\alpha}^{s} \mathcal{F}_{s}^{r} \nabla \psi_{r}^{\alpha}\right)+c . c ., \tag{2.2}
\end{align*}
$$

and the transformations

$$
\begin{align*}
\delta_{\epsilon} e^{a} & =-\frac{1}{2}\left(\bar{\psi}^{r} \Gamma^{a} \epsilon_{r}-\bar{\epsilon}^{r} \Gamma^{a} \psi_{r}\right), & \delta_{\epsilon} \omega^{a b} & =\frac{1}{4}\left(\bar{\psi}^{r} \Gamma^{a b} \epsilon_{r}-\bar{\epsilon}^{r} \Gamma^{a b} \psi_{r}\right), \\
\delta_{\epsilon} \psi_{r} & =-\nabla \epsilon_{r}, & \delta_{\epsilon} \bar{\psi}^{r} & =-\nabla \bar{\epsilon}^{r}, \\
\delta_{\epsilon} A_{s}^{r} & =i\left(\bar{\psi}^{r} \epsilon_{s}-\bar{\epsilon}^{r} \psi_{s}\right), & \delta_{\epsilon} B & =i\left(\bar{\psi}^{r} \epsilon_{r}-\bar{\epsilon}^{r} \psi_{r}\right) \tag{2.3}
\end{align*}
$$

In these expressions, $\Gamma^{a}$ are the Dirac matrices ${ }^{6}, \Gamma^{a b} \equiv \frac{1}{2}\left(\Gamma^{a} \Gamma^{b}-\Gamma^{b} \Gamma^{a}\right), R^{a b}=d \omega^{a b}+$ $\omega^{a c} \omega_{c}{ }^{b}$ is the curvature 2-form, $T^{a}=d e^{a}+\omega^{a}{ }_{b} e^{b}$ is the torsion 2-form, $F=d A+G_{A} A^{2}=$ $F^{i} \tau_{i}$ is the $s u(N)$ curvature,

$$
\begin{align*}
\mathcal{R}_{\beta}^{\alpha} & \equiv \frac{1}{2 L} T^{a}\left(\Gamma_{a}\right)_{\beta}^{\alpha}+\frac{1}{4}\left(R^{a b}+\frac{1}{L^{2}} e^{a} e^{b}\right)\left(\Gamma_{a b}\right)_{\beta}^{\alpha}+\frac{i}{4} G_{B} d B \delta_{\beta}^{\alpha}-\frac{1}{2} \psi_{s}^{\alpha} \bar{\psi}_{\beta}^{s} \\
\mathcal{F}_{s}^{r} & \equiv F_{s}^{r}+\frac{i}{N} G_{B} d B \delta_{s}^{r}-\frac{1}{2} \bar{\psi}_{\beta}^{r} \psi_{s}^{\beta} \tag{2.4}
\end{align*}
$$

str is a symmetrized trace satisfying $\operatorname{str}\left(\tau_{i} \tau_{j} \tau_{k}\right) \equiv \frac{1}{2 i} \operatorname{tr}\left(\left\{\tau_{i}, \tau_{j}\right\} \tau_{k}\right), \nabla$ is the $a d S_{5} \times s u(N) \times$ $u(1)$ covariant derivative, and

$$
\begin{align*}
\nabla \psi_{r} & \equiv\left(d+\frac{1}{4} \omega^{a b} \Gamma_{a b}+\frac{1}{2 L} e^{a} \Gamma_{a}\right) \psi_{r}-G_{A} A_{r}^{s} \psi_{s}+i\left(\frac{1}{4}-\frac{1}{N}\right) G_{B} B \psi_{r}, \\
\nabla \epsilon_{r} & \equiv\left(d+\frac{1}{4} \omega^{a b} \Gamma_{a b}+\frac{1}{2 L} e^{a} \Gamma_{a}\right) \epsilon_{r}-G_{A} A_{r}^{s} \epsilon_{s}+i\left(\frac{1}{4}-\frac{1}{N}\right) G_{B} B \epsilon_{r} . \tag{2.5}
\end{align*}
$$

Note that the results in the Appendix can be used to show that the torsion vanishes for our metric.

We impose the following orbifold conditions:

[^3]1. Periodicity on $S^{1}$. The fields and the fermionic parameters $\epsilon_{r}$, denoted generically by $\phi$, are required to be periodic on the circle $S^{1}$. That is,

$$
\begin{equation*}
\phi\left(x^{\bar{\mu}}, x^{5}\right)=\phi\left(x^{\bar{\mu}}, x^{5}+2 \pi \rho\right) . \tag{2.6}
\end{equation*}
$$

2. $\mathbb{Z}_{2}$ parity assignments. The bosonic field components

$$
\Phi=e_{\mu}^{\bar{a}}, e_{5}^{{ }^{\dot{5}}}, A_{5}^{i}, B_{5}, \quad \Theta=e_{\bar{\mu}}^{\dot{5}}, e_{5}^{\bar{a}}, A_{\bar{\mu}}^{i}, B_{\bar{\mu}}
$$

are chosen to satisfy

$$
\begin{equation*}
\Phi\left(x^{\mu}, x^{5}\right)=+\Phi\left(x^{\mu},-x^{5}\right), \quad \Theta\left(x^{\mu}, x^{5}\right)=-\Theta\left(x^{\mu},-x^{5}\right) \tag{2.7}
\end{equation*}
$$

That is, the $\Phi$ components are $\mathbb{Z}_{2}$-even and the $\Theta$ components are $\mathbb{Z}_{2}$-odd. For the gravitini, we require

$$
\begin{align*}
& \Gamma^{\dot{5}} \psi_{\bar{\mu} r}\left(x^{\bar{\mu}}, x^{5}\right)=+\psi_{\bar{\mu} r}\left(x^{\bar{\mu}},-x^{5}\right) \\
& \Gamma^{\dot{5}} \psi_{5 r}\left(x^{\bar{\mu}}, x^{5}\right)=-\psi_{5 r}\left(x^{\bar{\mu}},-x^{5}\right) \tag{2.8}
\end{align*}
$$

Finally, the fermionic parameters $\epsilon_{r}$ are required to satisfy

$$
\begin{equation*}
\Gamma^{\dot{5}} \epsilon_{r}\left(x^{\bar{\mu}}, x^{5}\right)=+\epsilon_{r}\left(x^{\mu},-x^{5}\right) \tag{2.9}
\end{equation*}
$$

These conditions imply that the $\mathbb{Z}_{2}$-odd quantities vanish at the orbifold fixed points. It is straightforward to check that $S_{\text {bulk }}$ is $\mathbb{Z}_{2}$-even and that the transformations (2.3) are consistent with the $\mathbb{Z}_{2}$ parity assignments.

## 3 Calculation of $\delta_{\epsilon} S_{\text {bulk }}$

The $D=5$ Chern-Simons AdS supergravity action is invariant (up to a boundary term) under its fermionic symmetry transformations. In Section 2, we modified this action and its fermionic transformations to obtain a $\mathbb{Z}_{2}$-invariant bulk theory. Due to the signum functions introduced by the modifications, $\delta_{\epsilon} S_{\text {bulk }}$ contains terms which have no counterpart in the unmodified theory. More specifically, the extra terms arise from $\partial_{5}$ acting on the signum functions to yield delta functions. Such 'delta function' contributions to $\delta_{\epsilon} S_{\text {bulk }}$ can potentially spoil the fermionic symmetry only at the $\mathbb{Z}_{2}$ fixed points. Thus, $S_{\text {bulk }}$ is invariant under its fermionic transformations everywhere except perhaps at the $\mathbb{Z}_{2}$ fixed points. In this section, we will calculate $\delta_{\epsilon} S_{\text {bulk }}$.

For our metric and $\mathbb{Z}_{2}$ parity assignments, the uncancelled variation $\delta_{\epsilon} S_{\text {bulk }}$ arises from the variation of the 4 -Fermi terms. The 4 -Fermi terms are

$$
\begin{align*}
S_{\psi^{4}}= & \frac{3 i}{4} \int K\left(\bar{\psi}_{\alpha}^{r} \psi_{s}^{\alpha} \bar{\psi}_{\beta}^{s} \nabla \psi_{r}^{\beta}+\bar{\psi}_{\alpha}^{s} \bar{\psi}_{\beta}^{r} \psi_{s}^{\beta} \nabla \psi_{r}^{\alpha}\right)+c . c . \\
= & \frac{3 i}{2} \int K \bar{\psi}_{\alpha}^{r} \psi_{s}^{\alpha} \bar{\psi}_{\beta}^{s} \nabla \psi_{r}^{\beta}+c . c . \\
= & \frac{3 i}{2} \int K\left(\bar{\psi}^{r} \psi_{s}\right)\left(\bar{\psi}^{s} \nabla \psi_{r}\right)+c . c . \\
= & \frac{3 i}{2} \frac{1}{5!} \int d^{5} x \varepsilon^{\mu \nu \rho \sigma \lambda} K\left(\bar{\psi}_{\mu}^{r} \psi_{\nu s}\right)\left(\bar{\psi}_{\rho}^{s} \nabla_{\sigma} \psi_{\lambda r}\right)+c . c . \\
= & \frac{3 i}{2} \frac{1}{4!} \int d^{5} x K\left[\varepsilon^{5 \bar{\nu} \bar{\sigma} \bar{\lambda}}\left(\bar{\psi}_{5}^{r} \psi_{\bar{\nu} s}\right)\left(\bar{\psi}_{\bar{\rho}}^{s} \nabla_{\bar{\sigma}} \psi_{\bar{\lambda} r}\right)+\varepsilon^{\bar{\mu} 5 \bar{\rho} \bar{\sigma} \bar{\lambda}}\left(\bar{\psi}_{\bar{\mu}}^{r} \psi_{5 s}\right)\left(\bar{\psi}_{\bar{\rho}}^{s} \nabla_{\bar{\sigma}} \psi_{\bar{\lambda} r}\right)\right. \\
& +\varepsilon^{\bar{\mu} \bar{\nu} \bar{\sigma} \bar{\lambda}}\left(\bar{\psi}_{\bar{\mu}}^{r} \psi_{\bar{\nu} s}\right)\left(\bar{\psi}_{5}^{s} \nabla_{\bar{\sigma}} \psi_{\bar{\lambda} r}\right)+\varepsilon^{\bar{\mu} \bar{\rho} \bar{\rho} \bar{\lambda}}\left(\bar{\psi}_{\bar{\mu}}^{r} \psi_{\bar{\nu} s}\right)\left(\bar{\psi}_{\bar{\rho}}^{s} \nabla_{5} \psi_{\bar{\lambda} r}\right) \\
& \left.+\varepsilon^{\bar{\mu} \bar{\rho} \bar{\rho} 5}\left(\bar{\psi}_{\bar{\mu}}^{r} \psi_{\bar{\nu} s}\right)\left(\bar{\psi}_{\bar{\rho}}^{s} \nabla_{\bar{\sigma}} \psi_{5 r}\right)\right]+c . c . \tag{3.1}
\end{align*}
$$

Let us now compute $\delta_{\epsilon} S_{\text {bulk }}$ by applying $\delta_{\epsilon}$ to (3.1) and dropping all terms which contribute no delta functions. For our metric and $\mathbb{Z}_{2}$ parity assignments, we can drop all but

1. The $\partial_{\mu}$ part of $\nabla_{\mu}$.
2. The $-\partial_{\mu} \epsilon_{r}$ part of $\delta_{\epsilon}=-\nabla_{\mu} \epsilon_{r}$.

The only contributing terms are thus contained in the expression

$$
\begin{align*}
Q \equiv & -\frac{3 i}{2} \frac{1}{4!} \int d^{5} x K\left\{\varepsilon^{5 \bar{\nu} \bar{\rho} \bar{\sigma}}\left(\partial_{5}{ }^{r} \psi_{\bar{\nu} s}\right)\left(\bar{\psi}_{\bar{\rho}}^{s} \partial_{\bar{\sigma}} \psi_{\bar{\lambda} r}\right)\right. \\
& +\varepsilon^{\bar{\mu} 5 \bar{\rho} \bar{\sigma} \bar{\lambda}}\left(\bar{\psi}_{\bar{\mu}}^{r} \partial_{5} \epsilon_{s}\right)\left(\bar{\psi}_{\bar{\rho}}^{s} \partial_{\bar{\sigma}} \psi_{\bar{\lambda} r}\right)+\varepsilon^{\bar{\mu} \bar{\nu} 5 \bar{\sigma} \bar{\lambda}}\left(\bar{\psi}_{\bar{\mu}}^{r} \psi_{\bar{\nu} s}\right)\left(\partial_{5} \bar{\epsilon}^{s} \partial_{\bar{\sigma}} \psi_{\bar{\lambda} r}\right) \\
& +\varepsilon^{\bar{\mu} \bar{\nu} \bar{\rho} \bar{\lambda}}\left[\left(\partial_{\bar{\mu}} \bar{\epsilon}^{r} \psi_{\bar{\nu} s}\right)\left(\bar{\psi}_{\bar{\rho}}^{s} \partial_{5} \psi_{\bar{\lambda} r}\right)+\left(\bar{\psi}_{\bar{\mu}}^{r} \partial_{\bar{\nu}} \epsilon_{s}\right)\left(\bar{\psi}_{\bar{\rho}}^{s} \partial_{5} \psi_{\bar{\lambda} r}\right)\right. \\
& \left.\quad+\left(\bar{\psi}_{\bar{\mu}}^{r} \psi_{\bar{\nu} s}\right)\left(\partial_{\bar{\rho}} \bar{\epsilon}^{s} \partial_{5} \psi_{\bar{\lambda} r}\right)+\left(\bar{\psi}_{\bar{\mu}}^{r} \psi_{\bar{\nu} s}\right)\left(\bar{\psi}_{\bar{\rho}}^{s} \partial_{5} \partial_{\bar{\lambda}} \epsilon_{r}\right)\right] \\
& \left.+\varepsilon^{\bar{\mu} \bar{\nu} \bar{\rho} 5}\left(\bar{\psi}_{\bar{\mu}}^{r} \psi_{\bar{\nu} s}\right)\left(\bar{\psi}_{\bar{\rho}}^{s} \partial_{\bar{\sigma}} \partial_{5} \epsilon_{r}\right)\right\}+ \text { c.c. } \tag{3.2}
\end{align*}
$$

More specifically, the delta function terms contained in $Q$ are obtained by integrating by parts and keeping only the terms in which $\partial_{5}$ acts on $K$. Thus,

$$
\begin{align*}
\delta_{\epsilon} S_{b u l k}= & \frac{3 i}{2} \frac{1}{4!} \int d^{5} x \partial_{5} K\left\{\varepsilon^{5 \bar{\nu} \bar{\rho} \bar{\lambda}}\left(\bar{\epsilon}^{r} \psi_{\bar{\nu} s}\right)\left(\bar{\psi}_{\bar{\rho}}^{s} \partial_{\bar{\sigma}} \psi_{\bar{\lambda} r}\right)\right. \\
& +\varepsilon^{\bar{\mu} 5 \bar{\rho} \bar{\sigma} \bar{\lambda}}\left(\bar{\psi}_{\bar{\mu}}^{r} \epsilon_{s}\right)\left(\bar{\psi}_{\bar{\rho}}^{s} \partial_{\bar{\sigma}} \psi_{\bar{\lambda} r}\right)+\varepsilon^{\bar{\mu} \bar{\nu} \bar{\sigma} \bar{\lambda}}\left(\bar{\psi}_{\bar{\mu}}^{r} \psi_{\bar{\nu} s}\right)\left(\bar{\epsilon}^{s} \partial_{\bar{\sigma}} \psi_{\bar{\lambda} r}\right) \\
& +\varepsilon^{\bar{\mu} \bar{\rho} \bar{\rho} \bar{\lambda}}\left[\left(\partial_{\bar{\mu}} \bar{\epsilon}^{r} \psi_{\bar{\nu} s}\right)\left(\bar{\psi}_{\bar{\rho}}^{s} \psi_{\bar{\lambda} r}\right)+\left(\bar{\psi}_{\bar{\mu}}^{r} \partial_{\bar{\nu}} \epsilon_{s}\right)\left(\bar{\psi}_{\bar{\rho}}^{s} \psi_{\bar{\lambda} r}\right)\right. \\
& \left.\quad+\left(\bar{\psi}_{\bar{\mu}}^{r} \psi_{\bar{\nu} s}\right)\left(\partial_{\bar{\rho}} \bar{\epsilon}^{s} \psi_{\bar{\lambda} r}\right)+\left(\bar{\psi}_{\bar{\mu}}^{r} \psi_{\bar{\nu} s}\right)\left(\bar{\psi}_{\bar{\rho}}^{s} \partial_{\bar{\lambda}} \epsilon_{r}\right)\right] \\
& \left.+\varepsilon^{\bar{\mu} \bar{\rho} \bar{\sigma} 5}\left(\bar{\psi}_{\bar{\mu}}^{r} \psi_{\bar{\nu} s}\right)\left(\bar{\psi}_{\bar{\rho}}^{s} \partial_{\bar{\sigma}} \epsilon_{r}\right)\right\}+c . c ., \tag{3.3}
\end{align*}
$$

where

$$
\begin{equation*}
\partial_{5} K=2 k\left[\delta\left(x^{5}\right)-\delta\left(x^{5}-\pi \rho\right)\right] \tag{3.4}
\end{equation*}
$$

## 4 Modified $\mathbb{Z}_{2}$-invariant bulk theory

The result (3.3) for $\delta_{\epsilon} S_{\text {bulk }}$ demonstrates that $S_{\text {bulk }}$ is not invariant under the fermionic transformations $\delta_{\epsilon}$. In this section, we will modify $S_{b u l k}$ and $\delta_{\epsilon}$ by replacing the $a d S_{5} \times$ $s u(N) \times u(1)$ covariant derivative $\nabla$ with $\widetilde{\nabla}$, where

$$
\begin{align*}
\widetilde{\nabla}_{\sigma} \psi_{\lambda r} & \equiv \nabla_{\sigma} \psi_{\lambda r}+2 \delta_{\sigma}^{5} \delta_{\lambda}^{\bar{\lambda}}\left[\delta\left(x^{5}\right)-\delta\left(x^{5}-\pi \rho\right)\right] \operatorname{sgn}\left(x^{5}\right) \Gamma_{\dot{5}} \psi_{\bar{\lambda} r}, \\
\widetilde{\nabla}_{\sigma} \epsilon_{r} & \equiv \nabla_{\sigma} \epsilon_{r}+2 \delta_{\sigma}^{5}\left[\delta\left(x^{5}\right)-\delta\left(x^{5}-\pi \rho\right)\right] \operatorname{sgn}\left(x^{5}\right) \Gamma_{\dot{5}} \epsilon_{r} \tag{4.1}
\end{align*}
$$

We will show that the modified bulk action

$$
\begin{equation*}
\tilde{S}_{b u l k} \equiv S_{b u l k}(\nabla \rightarrow \widetilde{\nabla}) \equiv S_{b u l k}+\Delta S_{b u l k} \tag{4.2}
\end{equation*}
$$

is invariant under the modified transformations

$$
\begin{equation*}
\tilde{\delta}_{\epsilon} \equiv \delta_{\epsilon}(\nabla \rightarrow \widetilde{\nabla}) \equiv \delta_{\epsilon}+\Delta \delta_{\epsilon} \tag{4.3}
\end{equation*}
$$

That is, we will show that

$$
\begin{equation*}
\tilde{\delta}_{\epsilon} \tilde{S}_{b u l k}=\delta_{\epsilon} S_{b u l k}+\left(\Delta \delta_{\epsilon}\right) S_{b u l k}+\tilde{\delta}_{\epsilon}\left(\Delta S_{b u l k}\right) \tag{4.4}
\end{equation*}
$$

vanishes. It is straightforward to check that $\tilde{S}_{\text {bulk }}$ is $\mathbb{Z}_{2}$-invariant and the transformations $\tilde{\delta}_{\epsilon}$ are consistent with our $\mathbb{Z}_{2}$ parity assignments.

We begin by computing $\left(\Delta \delta_{\epsilon}\right) S_{\text {bulk }}$. For our metric and $\mathbb{Z}_{2}$ parity assignments, the only part of $S_{\text {bulk }}$ which is not invariant under $\Delta \delta_{\epsilon}$ is $S_{\psi^{4}}$ (given by (3.1)). Note that

$$
\begin{equation*}
\left(\Delta \delta_{\epsilon}\right) \psi_{\lambda r}=-2 \delta_{\lambda}^{5}\left[\delta\left(x^{5}\right)-\delta\left(x^{5}-\pi \rho\right)\right] \operatorname{sgn}\left(x^{5}\right) \Gamma_{\dot{5}} \epsilon_{r} \tag{4.5}
\end{equation*}
$$

Thus, after using $K \operatorname{sgn}\left(x^{5}\right)=k$, (2.9), and (3.4), we obtain

$$
\begin{align*}
\left(\Delta \delta_{\epsilon}\right) S_{b u l k}= & -\frac{3 i}{2} \frac{1}{4!} \int d^{5} x \partial_{5} K\left[\varepsilon^{5 \bar{\nu} \bar{\rho} \bar{\sigma} \bar{\lambda}}\left(\bar{\epsilon}^{r} \psi_{\bar{\nu} s}\right)\left(\bar{\psi}_{\bar{\rho}}^{s} \partial_{\bar{\sigma}} \psi_{\bar{\lambda} r}\right)\right. \\
& +\varepsilon^{\bar{\mu} 5 \bar{\rho} \bar{\sigma} \bar{\lambda}}\left(\bar{\psi}_{\bar{\mu}}^{r} \epsilon_{s}\right)\left(\bar{\psi}_{\bar{\rho}}^{s} \partial_{\bar{\sigma}} \psi_{\bar{\lambda} r}\right)+\varepsilon^{\bar{\mu} \bar{\nu} 5 \bar{\sigma} \bar{\lambda}}\left(\bar{\psi}_{\bar{\mu}}^{r} \psi_{\bar{\nu} s}\right)\left(\bar{\epsilon}^{s} \partial_{\bar{\sigma}} \psi_{\bar{\lambda} r}\right) \\
& \left.+\varepsilon^{\bar{\mu} \bar{\rho} \bar{\rho} 5}\left(\bar{\psi}_{\bar{\mu}}^{r} \psi_{\bar{\nu} s}\right)\left(\bar{\psi}_{\bar{\rho}}^{s} \partial_{\bar{\sigma}} \epsilon_{r}\right)\right]+c . c . \tag{4.6}
\end{align*}
$$

Now, let us compute $\tilde{\delta}_{\epsilon}\left(\Delta S_{b u l k}\right)$. For our metric and $\mathbb{Z}_{2}$ parity assignments, the only part of $S_{b u l k}$ which is changed by the replacement $\nabla \rightarrow \widetilde{\nabla}$ is $S_{\psi^{4}}$. After using $K \operatorname{sgn}\left(x^{5}\right)=k$, (2.8), and (3.4), we obtain

$$
\begin{equation*}
\Delta S_{\text {bulk }}=\frac{3 i}{2} \frac{1}{4!} \int d^{5} x \partial_{5} K \varepsilon^{\bar{\mu} \bar{\nu} \bar{\rho} 5 \bar{\lambda}}\left(\bar{\psi}_{\bar{\mu}}^{r} \psi_{\bar{\nu} s}\right)\left(\bar{\psi}_{\bar{\rho}}^{s} \psi_{\bar{\lambda} r}\right)+\text { c.c. } \tag{4.7}
\end{equation*}
$$

Applying $\tilde{\delta}_{\epsilon}$ to (4.7) yields

$$
\begin{align*}
\tilde{\delta}_{\epsilon}\left(\Delta S_{b u l k}\right)= & -\frac{3 i}{2} \frac{1}{4!} \int d^{5} x \partial_{5} K \varepsilon^{\bar{\mu} \bar{\nu} \bar{\rho} \bar{\lambda}}\left[\left(\partial_{\bar{\mu}} \bar{\epsilon}^{r} \psi_{\bar{\nu} s}\right)\left(\bar{\psi}_{\bar{\rho}}^{s} \psi_{\bar{\lambda} r}\right)+\left(\bar{\psi}_{\bar{\mu}}^{r} \partial_{\bar{\nu}} \epsilon_{s}\right)\left(\bar{\psi}_{\bar{\rho}}^{s} \psi_{\bar{\lambda} r}\right)\right. \\
& \left.+\left(\bar{\psi}_{\bar{\mu}}^{r} \psi_{\bar{\nu} s}\right)\left(\partial_{\bar{\rho}} \bar{\epsilon}^{s} \psi_{\bar{\lambda} r}\right)+\left(\bar{\psi}_{\bar{\mu}}^{r} \psi_{\bar{\nu} s}\right)\left(\bar{\psi}_{\bar{\rho}}^{s} \partial_{\bar{\lambda}} \epsilon_{r}\right)\right]+c . c . \tag{4.8}
\end{align*}
$$

Using the results (3.3), (4.6), and (4.8) in (4.4) yields

$$
\begin{equation*}
\tilde{\delta}_{\epsilon} \tilde{S}_{b u l k}=0 \tag{4.9}
\end{equation*}
$$

## 5 Brane action

To complete our model, we add the brane action $S_{\text {brane }}$. In the absence of particle excitations, the brane action consists of brane tensions. That is,

$$
\begin{equation*}
S_{\text {brane }}=-\int d^{5} x e^{(4)}\left[\mathcal{T}^{(0)} \delta\left(x^{5}\right)+\mathcal{T}^{(\pi \rho)} \delta\left(x^{5}-\pi \rho\right)\right]+\text { excitations } \tag{5.1}
\end{equation*}
$$

where $e^{(4)} \equiv \operatorname{det}\left(e_{\bar{\mu}}{ }^{\bar{a}}\right)$. As discussed in Section $2, \mathbb{Z}_{2}$-odd quantities vanish at the $\mathbb{Z}_{2}$ fixed points. Thus, it is clear that $S_{\text {brane }}$ is $\mathbb{Z}_{2}$-even. Further discussion of 3-brane actions can be found in [11].

## 6 Connection with original RS model

In this section, we will show that our total action $S=\tilde{S}_{\text {bulk }}+S_{\text {brane }}$ reduces to the original Randall-Sundrum model [10] when $\tilde{S}_{\text {bulk }}$ is restricted to its gravitational sector.

The gravitational sector of $\tilde{S}_{b u l k}$ is $S_{\text {grav }}$, given by the first equation of (2.2). $S_{\text {grav }}$ consists of three terms:

1. The 'Gauss-Bonnet' term $\int \frac{1}{8} K \varepsilon_{a b c d e} R^{a b} R^{c d} e^{e} / L$.
2. The 'Einstein-Hilbert' term $\int \frac{1}{8} \cdot \frac{2}{3} K \varepsilon_{a b c d e} R^{a b} e^{c} e^{d} e^{e} / L^{3}$.
3. The 'cosmological constant' term $\int \frac{1}{8} \cdot \frac{1}{5} K \varepsilon_{a b c d e} e^{a} e^{b} e^{c} e^{d} e^{e} / L^{5}$.

For our metric, the first term can be expressed as a linear combination of the other two. Summing the three terms yields an effective Einstein-Hilbert term and an effective cosmological constant term. To demonstrate this explicitly, let us evaluate $S_{\text {grav }}$ for our metric. Using the results in the Appendix, we obtain

$$
\begin{align*}
& \varepsilon_{a b c d e} R^{a b} R^{c d} e^{e}=d^{5} x e\left(-\frac{120}{\ell^{4}}+\frac{192}{\ell^{3}}\left[\delta\left(x^{5}\right)-\delta\left(x^{5}-\pi \rho\right)\right]\right), \\
& \varepsilon_{a b c d e} R^{a b} e^{c} e^{d} e^{e}=d^{5} x e(-6 R) \\
& \varepsilon_{a b c d e} e^{a} e^{b} e^{c} e^{d} e^{e}=d^{5} x e(-5!) \tag{6.1}
\end{align*}
$$

where $e \equiv \operatorname{det}\left(e_{\mu}{ }^{a}\right)$. Thus,

$$
\begin{align*}
S_{\text {grav }}= & \int d^{5} x e \frac{1}{8}\left\{\frac{k}{\ell}\left(-\frac{120}{\ell^{4}}+\frac{192}{\ell^{3}}\left[\delta\left(x^{5}\right)-\delta\left(x^{5}-\pi \rho\right)\right]\right)\right. \\
& \left.+\frac{2 k}{3 \ell^{3}}(-6 R)+\frac{k}{5 \ell^{5}}(-5!)\right\} \\
= & \int d^{5} x e \frac{k}{\ell^{3}}\left\{-\frac{15}{\ell^{2}}+\frac{24}{\ell}\left[\delta\left(x^{5}\right)-\delta\left(x^{5}-\pi \rho\right)\right]-\frac{1}{2} R-\frac{3}{\ell^{2}}\right\} \\
= & \int d^{5} x e \frac{k}{\ell^{3}}\left\{\frac{3}{2}\left(-\frac{20}{\ell^{2}}+\frac{16}{\ell}\left[\delta\left(x^{5}\right)-\delta\left(x^{5}-\pi \rho\right)\right]\right)-\frac{1}{2} R+\frac{12}{\ell^{2}}\right\} \\
= & \int d^{5} x e \frac{k}{\ell^{3}}\left(R+\frac{12}{\ell^{2}}\right) \\
= & \int d^{5} x e\left(2 M^{3} R-\Lambda\right), \tag{6.2}
\end{align*}
$$

where $M$ is the five-dimensional gravitational mass scale ${ }^{7}, \Lambda$ is the bulk cosmological constant, and

$$
\begin{equation*}
M^{3}=\frac{k}{2 \ell^{3}}, \quad \Lambda=-\frac{24 M^{3}}{\ell^{2}} \tag{6.3}
\end{equation*}
$$

Combining the result (6.2) with (5.1), we obtain the action of the original RandallSundrum model. It is shown in [10] that the five-dimensional vacuum Einstein's equations for this system,

$$
\begin{equation*}
R_{\mu \nu}-\frac{1}{2} g_{\mu \nu} R=-\frac{1}{4 M^{3}}\left\{g_{\mu \nu} \Lambda+\frac{e^{(4)}}{e} \delta_{\mu}^{\bar{\mu}} \delta_{\nu}^{\bar{\nu}} g_{\bar{\mu} \bar{\nu}}\left[\mathcal{T}^{(0)} \delta\left(x^{5}\right)+\mathcal{T}^{(\pi \rho)} \delta\left(x^{5}-\pi \rho\right)\right]\right\} \tag{6.4}
\end{equation*}
$$

are solved by our metric provided that the relations

$$
\begin{equation*}
\mathcal{T}^{(0)}=-\mathcal{T}^{(\pi \rho)}=24 M^{3} / \ell, \quad \Lambda=-24 M^{3} / \ell^{2} \tag{6.5}
\end{equation*}
$$

are satisfied.

## 7 Killing spinors

In this section, we will solve the Killing spinor equations for a purely bosonic configuration with vanishing $s u(N)$ and $u(1)$ gauge fields. In this case, the Killing spinor equations reduce to

$$
\begin{align*}
& 0=\delta_{\epsilon} \psi_{\bar{\mu} r}=-\partial_{\bar{\mu}} \epsilon_{r}-\frac{1}{2} \frac{\mathfrak{a}^{\prime}}{\mathfrak{a}} \Gamma_{\bar{\mu}}\left(\Gamma_{\dot{5}}-1\right) \epsilon_{r}, \\
& 0=\delta_{\epsilon} \psi_{5 r}=-\partial_{5} \epsilon_{r}+\frac{1}{2} \frac{\mathfrak{a}^{\prime}}{\mathfrak{a}} \Gamma_{\dot{5}} \epsilon_{r}-2\left[\delta\left(x^{5}\right)-\delta\left(x^{5}-\pi \rho\right)\right] \operatorname{sgn}\left(x^{5}\right) \Gamma_{\dot{5}} \epsilon_{r} \tag{7.1}
\end{align*}
$$

[^4]To solve these equations, split $\epsilon_{r}$ into $\mathbb{Z}_{2}$-even $\left(\epsilon_{r}^{+}\right)$and $\mathbb{Z}_{2}$-odd $\left(\epsilon_{r}^{-}\right)$pieces:

$$
\begin{equation*}
\epsilon_{r}=\epsilon_{r}^{+}+\epsilon_{r}^{-}, \tag{7.2}
\end{equation*}
$$

where

$$
\begin{equation*}
\epsilon^{ \pm} \equiv \frac{1}{2}\left(\epsilon_{r} \pm \Gamma_{\dot{5}} \epsilon_{r}\right)= \pm \Gamma_{\dot{5}} \epsilon_{r}^{ \pm} . \tag{7.3}
\end{equation*}
$$

We obtain the following system of equations:

$$
\begin{align*}
& \partial_{\bar{\mu}} \epsilon_{r}^{+}=-\left(\mathfrak{a}^{\prime} / \mathfrak{a}\right) \Gamma_{\bar{\mu}} \Gamma_{\dot{5}} \epsilon_{r}^{-}, \\
& \partial_{\bar{\mu}} \epsilon_{r}^{-}=0, \\
& \partial_{5} \epsilon_{r}^{+}=+\frac{1}{2}\left(\mathfrak{a}^{\prime} / \mathfrak{a}\right) \epsilon_{r}^{+}-2\left[\delta\left(x^{5}\right)-\delta\left(x^{5}-\pi \rho\right)\right] \operatorname{sgn}\left(x^{5}\right) \epsilon_{r}^{+}, \\
& \partial_{5} \epsilon_{r}^{-}=-\frac{1}{2}\left(\mathfrak{a}^{\prime} / \mathfrak{a}\right) \epsilon_{r}^{-}+2\left[\delta\left(x^{5}\right)-\delta\left(x^{5}-\pi \rho\right)\right] \operatorname{sgn}\left(x^{5}\right) \epsilon_{r}^{-} . \tag{7.4}
\end{align*}
$$

These equations are solved by

$$
\begin{align*}
\epsilon_{r}^{+} & =\mathfrak{a}^{1 / 2}\left[-\left(\mathfrak{a}^{\prime} / \mathfrak{a}^{2}\right) x^{\bar{\mu}} \Gamma_{\bar{\mu}} \Gamma_{\dot{5}} \operatorname{sgn}\left(x^{5}\right) \chi_{r}^{-(0)}+\chi_{r}^{+(0)}\right] \\
& =\mathfrak{a}^{1 / 2}\left[(1 / \ell) x^{\bar{\mu}} \delta_{\bar{\mu}}{ }^{\bar{a}} \Gamma_{\bar{a}} \Gamma_{\dot{5}} \chi_{r}^{-(0)}+\chi_{r}^{+(0)}\right], \\
\epsilon_{r}^{-} & =\mathfrak{a}^{-1 / 2} \operatorname{sgn}\left(x^{5}\right) \chi_{r}^{-(0)}, \tag{7.5}
\end{align*}
$$

where $\chi_{r}^{+(0)}$ and $\chi_{r}^{-(0)}$ are constant (projected) spinors. ${ }^{8}$ Thus, our solution for the Killing spinors is

$$
\begin{equation*}
\epsilon_{r}=\mathfrak{a}^{1 / 2} \chi_{r}^{+(0)}+\mathfrak{a}^{-1 / 2} \operatorname{sgn}\left(x^{5}\right)\left(1-\frac{\mathfrak{a}^{\prime}}{\mathfrak{a}} x^{\bar{\mu}} \Gamma_{\bar{\mu}} \Gamma_{\dot{5}}\right) \chi_{r}^{-(0)} . \tag{7.6}
\end{equation*}
$$

## 8 Conclusion

We have constructed a Randall-Sundrum scenario from $D=5$ Chern-Simons AdS supergravity. Our total action $S_{\tilde{\delta}}=\tilde{S}_{\text {bulk }}+S_{\text {brane }}$ is $\mathbb{Z}_{2}$-invariant. $\tilde{S}_{\text {bulk }}$ is invariant under the fermionic transformations $\tilde{\delta}_{\epsilon}$. However,

$$
\begin{equation*}
\tilde{\delta}_{\epsilon} S_{\text {brane }}=-\int d^{5} x \tilde{\delta}_{\epsilon} e^{(4)}\left[\mathcal{T}^{(0)} \delta\left(x^{5}\right)+\mathcal{T}^{(\pi \rho)} \delta\left(x^{5}-\pi \rho\right)\right]+\cdots, \tag{8.1}
\end{equation*}
$$

where

$$
\begin{equation*}
\tilde{\delta}_{\epsilon} e^{(4)}=e^{(4)}\left[-\frac{1}{2}\left(\bar{\psi}_{\bar{\mu}}^{r} \Gamma^{\bar{\mu}} \epsilon_{r}-\bar{\epsilon}^{r} \Gamma^{\bar{\mu}} \psi_{\bar{\mu} r}\right)\right] . \tag{8.2}
\end{equation*}
$$

Thus, the fermionic symmetry is broken by $S_{\text {brane }}$. Nevertheless, the Killing spinors of Section 7 are globally defined.

Our model reduces to the original Randall-Sundrum model [10] when $\tilde{S}_{\text {bulk }}$ is restricted to its gravitational sector. The original Randall-Sundrum model involves the fine-tuning relations

$$
\mathcal{T}^{(0)}=-\mathcal{T}^{(\pi \rho)}=24 M^{3} / \ell, \quad \Lambda=-24 M^{3} / \ell^{2}
$$

[^5]Randall-Sundrum scenarios constructed from standard $D=5$ supergravity theories yield these relations (up to an overall normalization factor) as a consequence of local supersymmetry (some examples are given in [12]). In our case, the relation $\Lambda=-24 M^{3} / \ell^{2}$ follows from our metric choice. We do not obtain the relations $\mathcal{T}^{(0)}=-\mathcal{T}^{(\pi \rho)}=24 M^{3} / \ell$ as a consequence of local fermionic symmetry. These are fine-tuning relations in our model.

## A Appendix

In this Appendix, we work out the fünfbein, spin connection, curvature 2-form components, Ricci tensor, and Ricci scalar for our metric (1.1). For the fünfbein, we obtain

$$
\begin{align*}
& e_{\bar{\mu}}^{\bar{a}}=\mathfrak{a} \delta_{\bar{\mu}}^{\bar{a}}, \quad e_{\bar{\mu} \bar{a}}=e_{\bar{\mu}}^{\bar{b}} \eta_{\bar{b} \bar{a}}, \quad e^{\bar{\mu} \bar{a}}=g^{\bar{\mu} \bar{\nu}} e_{\bar{\nu}}^{\bar{a}}, \\
& e_{\bar{a}}^{\bar{\mu}}=\mathfrak{a}^{-1} \delta_{\bar{a}}^{\bar{\mu}}, \quad e_{\bar{a} \bar{\mu}}=e_{\bar{a}}^{\bar{\nu}} g_{\bar{\nu} \bar{\mu}}, \quad e^{\bar{a} \bar{\mu}}=\eta^{\bar{a} \bar{b}} e_{\bar{b}}^{\bar{\mu}}, \\
& e_{5}^{\dot{5}}=e_{5 \dot{5}}=e^{55}=1, \quad e_{\dot{5}}{ }^{5}=e_{\dot{5} 5}=e^{\dot{55}}=1 . \tag{A.1}
\end{align*}
$$

Our conventions for the spin connection, curvature 2-form components, Ricci tensor, and Ricci scalar are respectively

$$
\begin{aligned}
\omega_{\mu}{ }^{a b}= & \frac{1}{2} e^{\nu a}\left(\partial_{\mu} e_{\nu}{ }^{b}-\partial_{\nu} e_{\mu}{ }^{b}\right)-\frac{1}{2} e^{\nu b}\left(\partial_{\mu} e_{\nu}{ }^{a}-\partial_{\nu} e_{\mu}{ }^{a}\right) \\
& -\frac{1}{2} e^{\rho a} e^{\sigma b}\left(\partial_{\rho} e_{\sigma c}-\partial_{\sigma} e_{\rho c}\right) e^{c}{ }_{\mu}, \\
R_{\mu \nu}{ }^{a b}= & \partial_{\mu} \omega_{\nu}^{a b}-\partial_{\nu} \omega_{\mu}{ }^{a b}+\omega_{\mu}{ }^{a c} \omega_{\nu c}{ }^{b}-\omega_{\nu}{ }^{a c} \omega_{\mu c}{ }^{b}, \\
R_{\nu \sigma}= & R_{\mu \nu}{ }^{a b} e_{a}{ }^{\mu} e_{b \sigma}, \quad R=e_{a}{ }^{\mu} e_{b}{ }^{\nu} R_{\mu \nu}{ }^{a b} .
\end{aligned}
$$

For the metric (1.1), the nonzero quantities here are

$$
\begin{gather*}
\omega_{\bar{\mu}}^{\bar{a} \dot{5}}=-\omega_{\bar{\mu}}^{\dot{5} \bar{a}}=\mathfrak{a}^{\prime} \delta_{\bar{\mu}}^{\bar{a}}=-e_{\bar{\mu}}^{\bar{a}} / L,  \tag{A.2}\\
R_{\bar{\mu} \bar{\nu}}^{\overline{\bar{b}}}=-\mathfrak{a}^{\prime 2}\left(\delta_{\bar{\mu}}{ }^{\bar{a}} \delta_{\bar{\nu}}^{\bar{b}}-\delta_{\bar{\mu}}^{\bar{b}} \delta_{\bar{\nu}}{ }^{\bar{a}}\right)=-\left(e_{\bar{\mu}}{ }^{\bar{a}} e_{\bar{\nu}}^{\bar{b}}-e_{\bar{\mu}}^{\bar{b}} e_{\bar{\nu}}^{\bar{a}}\right) / \ell^{2}, \\
R_{5 \bar{\mu}}{ }^{\bar{a} \dot{5}}=\mathfrak{a}^{\prime \prime} \delta_{\bar{\mu}}^{\bar{a}}=e_{\bar{\mu}}^{\bar{a}}\left\{1 / \ell^{2}-2\left[\delta\left(x^{5}\right)-\delta\left(x^{5}-\pi \rho\right)\right] / \ell\right\},  \tag{A.3}\\
R_{\bar{\mu} \bar{\nu}}=-\left(\mathfrak{a} \mathfrak{a}^{\prime \prime}+3 \mathfrak{a}^{\prime 2}\right) \eta_{\bar{\mu} \bar{\nu}}=-\left\{4 / \ell^{2}-2\left[\delta\left(x^{5}\right)-\delta\left(x^{5}-\pi \rho\right)\right] / \ell\right\} g_{\bar{\mu} \bar{\nu}}, \\
R_{55}=-4 \mathfrak{a}^{-1} \mathfrak{a}^{\prime \prime}=-\left\{4 / \ell^{2}-8\left[\delta\left(x^{5}\right)-\delta\left(x^{5}-\pi \rho\right)\right] / \ell\right\},  \tag{A.4}\\
R=-8 \mathfrak{a}^{-1} \mathfrak{a}^{\prime \prime}-12 \mathfrak{a}^{-2} \mathfrak{a}^{\prime 2}=-20 / \ell^{2}+16\left[\delta\left(x^{5}\right)-\delta\left(x^{5}-\pi \rho\right)\right] / \ell, \tag{A.5}
\end{gather*}
$$

and those related to (A.3) by $R_{\mu \nu}{ }^{a b}=-R_{\nu \mu}{ }^{a b}=-R_{\mu \nu}{ }^{b a}$. The prime symbol $/$ denotes partial differentiation with respect to $x^{5}$.

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[^1]:    ${ }^{1}$ The AdS group in dimension $D \geq 2$ is $S O(D-1,2)$. The corresponding super-AdS groups are given in [3]. For $D=5$ and $D=11$, the super-AdS groups are respectively $S U(2,2 \mid N)$ and $\operatorname{OSp}(32 \mid N)$.
    ${ }^{2}$ For example, in $D=5$ Chern-Simons AdS supergravity [1], the number of bosonic degrees of freedom $\left(N^{2}+15\right)$ is equal to the number of fermionic degrees of freedom ( $8 N$ ) only for $N=3$ and $N=5$.
    ${ }^{3}$ We use indices $\mu, \nu, \ldots=0,1,2,3,5$ for local spacetime and $a, b, \ldots=\dot{0}, \dot{1}, \dot{2}, \dot{3}, \dot{5}$ for tangent spacetime. The corresponding metrics, $g_{\mu \nu}$ and $\eta_{a b}=\operatorname{diag}(-1,1,1,1,1)_{a b}$, are related by $g_{\mu \nu}=e_{\mu}{ }^{a} e_{\nu}{ }^{b} \eta_{a b}$, where $e_{\mu}{ }^{a}$ is the fünfbein. Barred indices $\bar{\mu}, \bar{\nu}, \ldots=0,1,2,3$, and $\bar{a}, \bar{b}, \ldots=\dot{0}, \dot{1}, \dot{2}, \dot{3}$ denote the fourdimensional counterparts of $\mu, \nu, \ldots$ and $a, b, \ldots$, respectively.

[^2]:    ${ }^{4}$ We use indices $i, j, \ldots=1, \ldots, N^{2}-1$ to label the $N \times N$-dimensional $s u(N)$ generators $\tau_{i}$. The indices $r, s, \ldots=1, \ldots, N$ label vector representations of $s u(N)$. We will use the notation $A_{s}^{r} \equiv A^{i}\left(\tau_{i}\right)_{s}^{r}$. Spinor indices $\alpha, \beta, \ldots$ will sometimes be suppressed.

[^3]:    ${ }^{5}$ The signum function $\operatorname{sgn}\left(x^{5}\right)$ is +1 for $0<x^{5}<\pi \rho$ and -1 for $-\pi \rho<x^{5}<0$. It obeys $\operatorname{sgn}^{2}\left(x^{5}\right)=1$ and $\partial_{5} \operatorname{sgn}\left(x^{5}\right)=2\left[\delta\left(x^{5}\right)-\delta\left(x^{5}-\pi \rho\right)\right]$.
    ${ }^{6}$ We choose a chiral basis for the $4 \times 4$ Dirac matrices

    $$
    \Gamma^{a}=\left(\Gamma^{\bar{a}}, \Gamma^{\dot{5}}\right)=\left(\left[\begin{array}{cc}
    \mathbf{0} & -i \sigma^{\bar{a}} \\
    -i \bar{\sigma}^{\bar{a}} & \mathbf{0}
    \end{array}\right],\left[\begin{array}{cc}
    \mathbf{- 1} & \mathbf{0} \\
    \mathbf{0} & \mathbf{1}
    \end{array}\right]\right)
    $$

    where $\sigma^{\bar{a}}=(\mathbf{1}, \vec{\sigma})$ and $\bar{\sigma}^{\bar{a}}=(\mathbf{1},-\vec{\sigma})$. These matrices satisfy $\operatorname{tr}\left(\Gamma_{a} \Gamma_{b} \Gamma_{c} \Gamma_{d} \Gamma_{e}\right)=-4 i \varepsilon_{a b c d e}$, where $\varepsilon_{a b c d e}$ is the Levi-Civita tensor and $\varepsilon^{\dot{0} \dot{2} \dot{3} \dot{5}}=1$.

[^4]:    ${ }^{7} M$ is related to the four-dimensional gravitational mass scale $M_{(4)}=2.43 \times 10^{18} \mathrm{GeV}$ by $M_{(4)}^{2}=$ $M^{3} \int_{-\pi \rho}^{+\pi \rho} d x^{5} \mathfrak{a}^{2}\left(x^{5}\right)=M^{3} \ell[1-\exp (-2 \pi \rho / \ell)]$. The effective mass scales on the 3 -branes at $x^{5}=0$ and $x^{5}=\pi \rho$ are respectively $M_{(4)}$ and $M_{(4)} e^{-\pi \rho / \ell}$. If the Standard Model fields live on the 3-brane at $x^{5}=\pi \rho$, then $M_{(4)} e^{-\pi \rho / \ell}$ can be associated with the electroweak scale.

[^5]:    ${ }^{8}$ It is straightforward to check that (7.5) satisfies the first, second, and fourth equations of (7.4). There is, however, a subtlety in checking that (7.5) satisfies the third equation of (7.4). Unlike $\epsilon_{r}^{-}, \epsilon_{r}^{+}$is a smooth function of $x^{5}$. Thus, the second term on the right side of the third equation of (7.4) contributes nothing.

