Chern-Simons AdS_5 supergravity in a Randall-Sundrum background

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Abstract

Chern-Simons AdS supergravity theories are gauge theories for the super-AdS group. These theories possess a fermionic symmetry which differs from standard supersymmetry. In this paper, we study five-dimensional Chern-Simons AdS supergravity in a Randall-Sundrum scenario with two Minkowski 3-branes. After making modifications to the D = 5 Chern-Simons AdS supergravity action and fermionic symmetry transformations, we obtain a \mathbb{Z}_2 -invariant total action $S = \tilde{S}_{bulk} + S_{brane}$ and fermionic transformations δ_{ϵ} . While $\delta_{\epsilon} \tilde{S}_{bulk} = 0$, the fermionic symmetry is broken by S_{brane} . Our total action reduces to the original Randall-Sundrum model when \tilde{S}_{bulk} is restricted to its gravitational sector. We solve the Killing spinor equations for a bosonic configuration with vanishing su(N) and u(1) gauge fields.

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1 Introduction

Chern-Simons AdS supergravity [1, 2, 3] theories can be constructed only in odd spacetime dimensions. As the name implies, they are gauge theories for supersymmetric extensions of the AdS group.¹ They have a fiber bundle structure and hence are potentially renormalizable [2]. The dynamical fields form a single adS superalgebra-valued connection and hence the supersymmetry algebra closes automatically off-shell without requiring auxiliary fields [4]. The Lagrangian in dimension D = 2n - 1 is a Chern-Simons (2n - 1)form for the super-adS connection and is a polynomial of order n in the corresponding curvature. Unlike standard supergravity theories, there can be a mismatch between the number of bosonic and fermionic degrees of freedom.² For this reason, the 'supersymmetry' of Chern-Simons AdS supergravity theories is perhaps better referred to as a fermionic symmetry.

D = 11, N = 1 Chern-Simons AdS supergravity may correspond to an off-shell supergravity limit of M-theory [2, 3]. It has expected features of M-theory which are not shared by D = 11 Cremmer-Julia-Scherk (CJS) supergravity [5]. These features include an osp(32|1) superalgebra [6] and higher powers of curvature [7]. Hořava-Witten theory [8] is obtained from CJS supergravity by compactifying on an S^1/\mathbb{Z}_2 orbifold and requiring gauge and gravitational anomalies to cancel. This theory gives the low energy, strongly coupled limit of the heterotic $E_8 \times E_8$ string theory. In light of the above discussion, it would be interesting to reformulate Hořava-Witten theory with D=11, N=1 Chern-Simons AdS supergravity.

Reformulating Hořava-Witten theory as described above may prove to be difficult. It is simpler to compactify the five-dimensional version of Chern-Simons AdS supergravity on an S^1/\mathbb{Z}_2 orbifold and ignore anomaly cancellation issues. Canonical sectors of D =5 Chern-Simons AdS supergravity have been investigated in locally AdS_5 backgrounds possessing a spatial boundary with topology $S^1 \times S^1 \times S^1$ located at infinity [9]. In this paper, as a preamble to reformulating Hořava-Witten theory, we will study D = 5Chern-Simons AdS supergravity in a Randall-Sundrum background with two Minkowski 3-branes [10]. We choose coordinates $x^{\mu} = (x^{\bar{\mu}}, x^5)$ to parameterize the five-dimensional spacetime manifold.³ In terms of these coordinates, the background metric takes the form

$$g_{\mu\nu}dx^{\mu}dx^{\nu} = \mathfrak{a}^{2}(x^{5})\eta^{(4)}_{\bar{\mu}\bar{\nu}}dx^{\bar{\mu}}dx^{\bar{\nu}} + (dx^{5})^{2}, \qquad (1.1)$$

where $\eta_{\mu\bar{\nu}}^{(4)} = \text{diag}(-1, 1, 1, 1)_{\mu\bar{\nu}}$, $\mathfrak{a}(x^5) \equiv \exp(-|x^5|/\ell)$ is the warp factor, and ℓ is the AdS_5 curvature radius. The coordinate x^5 parameterizes an S^1/\mathbb{Z}_2 orbifold, where the circle S^1 has radius ρ and \mathbb{Z}_2 acts as $x^5 \to -x^5$. We choose the range $-\pi\rho \leq x^5 \leq \pi\rho$ with the endpoints identified as $x^5 \simeq x^5 + 2\pi\rho$. The Minkowski 3-branes are located at

¹The AdS group in dimension $D \ge 2$ is SO(D-1,2). The corresponding super-AdS groups are given in [3]. For D=5 and D=11, the super-AdS groups are respectively SU(2,2|N) and OSp(32|N).

²For example, in D=5 Chern-Simons AdS supergravity [1], the number of bosonic degrees of freedom $(N^2 + 15)$ is equal to the number of fermionic degrees of freedom (8N) only for N=3 and N=5.

³We use indices $\mu, \nu, \ldots = 0, 1, 2, 3, 5$ for local spacetime and $a, b, \ldots = \dot{0}, \dot{1}, \dot{2}, \dot{3}, \dot{5}$ for tangent spacetime. The corresponding metrics, $g_{\mu\nu}$ and $\eta_{ab} = \text{diag}(-1, 1, 1, 1, 1)_{ab}$, are related by $g_{\mu\nu} = e_{\mu}{}^{a}e_{\nu}{}^{b}\eta_{ab}$, where $e_{\mu}{}^{a}$ is the fünfbein. Barred indices $\bar{\mu}, \bar{\nu}, \ldots = 0, 1, 2, 3$, and $\bar{a}, \bar{b}, \ldots = \dot{0}, \dot{1}, \dot{2}, \dot{3}$ denote the fourdimensional counterparts of μ, ν, \ldots and a, b, \ldots , respectively.

the \mathbb{Z}_2 fixed points $x^5 = 0$ and $x^5 = \pi \rho$. These 3-branes have corresponding tensions $\mathcal{T}^{(0)}$ and $\mathcal{T}^{(\pi\rho)}$ and may support (3+1)-dimensional field theories.

This paper is organized as follows: In Section 2, we construct a \mathbb{Z}_2 -invariant bulk theory. This bulk theory is obtained by making modifications to the D=5 Chern-Simons AdS supergravity action and fermionic symmetry transformations which allow consistent orbifold conditions to be imposed. The variation of the resulting bulk action S_{bulk} under the resulting fermionic transformations δ_{ϵ} vanishes everywhere except at the \mathbb{Z}_2 fixed points. We calculate $\delta_{\epsilon}S_{bulk}$ in Section 3. In Section 4, we modify S_{bulk} and δ_{ϵ} to obtain a modified \mathbb{Z}_2 -invariant bulk theory. The modified bulk action \tilde{S}_{bulk} is invariant under the modified fermionic transformations δ_{ϵ} . In Section 5, we complete our model by adding the brane action S_{brane} . We show in Section 6 that our total action

$$S = S_{bulk} + S_{brane} \tag{1.2}$$

reduces to the original Randall-Sundrum model [10] when \tilde{S}_{bulk} is restricted to its gravitational sector. In Section 7, we solve the Killing spinor equations for a purely bosonic configuration with vanishing su(N) and u(1) gauge fields. Our concluding remarks are given in Section 8. Finally, in the Appendix, we work out the fünfbein, spin connection, curvature 2-form components, Ricci tensor, and Ricci scalar for our metric (1.1).

2 \mathbb{Z}_2 -invariant bulk theory

In this section, we construct a \mathbb{Z}_2 -invariant bulk theory. The bulk theory is obtained by making modifications to the D = 5 Chern-Simons AdS supergravity [1] action and fermionic symmetry transformations which allow consistent orbifold conditions to be imposed.

The field content of D=5 Chern-Simons AdS supergravity is the fünfbein $e_{\mu}{}^{a}$, the spin connection $\omega_{\mu}{}^{ab}$, the su(N) gauge connection $A_{\mu} = A^{i}_{\mu}\tau_{i}$, the u(1) gauge connection B_{μ} , and N complex gravitini $\psi_{\mu r}$ which transform as Dirac spinors in a vector representation of su(N).⁴ These fields form a connection for the adS superalgebra su(2, 2|N). The action and fermionic symmetry transformations are given in [9] in terms of the AdS_5 curvature radius ℓ . The only free parameter in the action is a dimensionless constant k. To allow consistent \mathbb{Z}_2 orbifold conditions to be imposed, we make the following modifications:

1. Rescale the su(N) and u(1) gauge connections:

$$A \to g_A A, \quad B \to g_B B.$$

⁴We use indices $i, j, \ldots = 1, \ldots, N^2 - 1$ to label the $N \times N$ -dimensional su(N) generators τ_i . The indices $r, s, \ldots = 1, \ldots, N$ label vector representations of su(N). We will use the notation $A_s^r \equiv A^i(\tau_i)_s^r$. Spinor indices α, β, \ldots will sometimes be suppressed.

2. Replace g_A , g_B , ℓ^{-1} , and k by the \mathbb{Z}_2 -odd expressions⁵

$$G_A \equiv g_A \operatorname{sgn}(x^5), \quad G_B \equiv g_B \operatorname{sgn}(x^5), \quad L^{-1} \equiv \ell^{-1} \operatorname{sgn}(x^5),$$

 $K \equiv k \operatorname{sgn}(x^5).$

In this manner, we obtain the bulk action

$$S_{bulk} = S_{grav} + S_{su(N)} + S_{u(1)} + S_{ferm},$$
(2.1)

where

$$S_{grav} = \int \frac{1}{8} K \varepsilon_{abcde} \left(\frac{1}{L} R^{ab} R^{cd} e^{e} + \frac{2}{3L^{3}} R^{ab} e^{c} e^{d} e^{e} + \frac{1}{5L^{5}} e^{a} e^{b} e^{c} e^{d} e^{e} \right),$$

$$S_{su(N)} = \int iK \operatorname{str} \left(G_{A}^{3} A F^{2} - \frac{1}{2} G_{A}^{4} A^{3} F + \frac{1}{10} G_{A}^{5} A^{5} \right),$$

$$S_{u(1)} = \int K \left[- \left(\frac{1}{4^{2}} - \frac{1}{N^{2}} \right) G_{B}^{3} B(dB)^{2} + \frac{3}{4L^{2}} \left(T^{a} T_{a} - \frac{L^{2}}{2} R^{ab} R_{ab} - R^{ab} e_{a} e_{b} \right) G_{B} B - \frac{3}{N} G_{A}^{2} G_{B} F_{s}^{r} F_{r}^{s} B \right],$$

$$S_{ferm} = \int \frac{3}{2i} K \left(\bar{\psi}_{\alpha}^{r} \mathcal{R}_{\beta}^{\alpha} \nabla \psi_{r}^{\beta} + \bar{\psi}_{\alpha}^{s} \mathcal{F}_{s}^{r} \nabla \psi_{r}^{\alpha} \right) + c.c.,$$
(2.2)

and the transformations

$$\delta_{\epsilon}e^{a} = -\frac{1}{2}(\bar{\psi}^{r}\Gamma^{a}\epsilon_{r} - \bar{\epsilon}^{r}\Gamma^{a}\psi_{r}), \qquad \delta_{\epsilon}\omega^{ab} = \frac{1}{4}(\bar{\psi}^{r}\Gamma^{ab}\epsilon_{r} - \bar{\epsilon}^{r}\Gamma^{ab}\psi_{r}), \\ \delta_{\epsilon}\psi_{r} = -\nabla\epsilon_{r}, \qquad \delta_{\epsilon}\bar{\psi}^{r} = -\nabla\bar{\epsilon}^{r}, \\ \delta_{\epsilon}A^{r}_{s} = i\left(\bar{\psi}^{r}\epsilon_{s} - \bar{\epsilon}^{r}\psi_{s}\right), \qquad \delta_{\epsilon}B = i\left(\bar{\psi}^{r}\epsilon_{r} - \bar{\epsilon}^{r}\psi_{r}\right).$$
(2.3)

In these expressions, Γ^a are the Dirac matrices⁶, $\Gamma^{ab} \equiv \frac{1}{2} (\Gamma^a \Gamma^b - \Gamma^b \Gamma^a)$, $R^{ab} = d\omega^{ab} + \omega^{ac} \omega_c^{\ b}$ is the curvature 2-form, $T^a = de^a + \omega^a{}_b e^b$ is the torsion 2-form, $F = dA + G_A A^2 = F^i \tau_i$ is the su(N) curvature,

$$\mathcal{R}^{\alpha}_{\beta} \equiv \frac{1}{2L} T^a (\Gamma_a)^{\alpha}_{\beta} + \frac{1}{4} (R^{ab} + \frac{1}{L^2} e^a e^b) (\Gamma_{ab})^{\alpha}_{\beta} + \frac{i}{4} G_B \, dB \, \delta^{\alpha}_{\beta} - \frac{1}{2} \psi^{\alpha}_s \bar{\psi}^s_{\beta},$$

$$\mathcal{F}^r_s \equiv F^r_s + \frac{i}{N} G_B \, dB \, \delta^r_s - \frac{1}{2} \bar{\psi}^r_{\beta} \psi^{\beta}_s,$$
(2.4)

str is a symmetrized trace satisfying $\operatorname{str}(\tau_i \tau_j \tau_k) \equiv \frac{1}{2i} \operatorname{tr}(\{\tau_i, \tau_j\}\tau_k), \nabla$ is the $adS_5 \times su(N) \times u(1)$ covariant derivative, and

$$\nabla \psi_r \equiv \left(d + \frac{1}{4}\omega^{ab}\Gamma_{ab} + \frac{1}{2L}e^a\Gamma_a\right)\psi_r - G_A A_r^s\psi_s + i\left(\frac{1}{4} - \frac{1}{N}\right)G_B B\psi_r,$$

$$\nabla \epsilon_r \equiv \left(d + \frac{1}{4}\omega^{ab}\Gamma_{ab} + \frac{1}{2L}e^a\Gamma_a\right)\epsilon_r - G_A A_r^s\epsilon_s + i\left(\frac{1}{4} - \frac{1}{N}\right)G_B B\epsilon_r.$$
 (2.5)

Note that the results in the Appendix can be used to show that the torsion vanishes for our metric.

We impose the following orbifold conditions:

 $^6\mathrm{We}$ choose a chiral basis for the 4×4 Dirac matrices

$$\Gamma^{a} = \left(\Gamma^{\bar{a}}, \Gamma^{5}\right) = \left(\begin{bmatrix} \mathbf{0} & -i\sigma^{\bar{a}} \\ -i\bar{\sigma}^{\bar{a}} & \mathbf{0} \end{bmatrix}, \begin{bmatrix} -\mathbf{1} & \mathbf{0} \\ \mathbf{0} & \mathbf{1} \end{bmatrix} \right)$$

where $\sigma^{\bar{a}} = (\mathbf{1}, \vec{\sigma})$ and $\bar{\sigma}^{\bar{a}} = (\mathbf{1}, -\vec{\sigma})$. These matrices satisfy tr $(\Gamma_a \Gamma_b \Gamma_c \Gamma_d \Gamma_e) = -4i\varepsilon_{abcde}$, where ε_{abcde} is the Levi-Civita tensor and $\varepsilon^{0\dot{1}\dot{2}\dot{3}\dot{5}} = 1$.

The signum function $\operatorname{sgn}(x^5)$ is +1 for $0 < x^5 < \pi\rho$ and -1 for $-\pi\rho < x^5 < 0$. It obeys $\operatorname{sgn}^2(x^5) = 1$ and $\partial_5 \operatorname{sgn}(x^5) = 2[\delta(x^5) - \delta(x^5 - \pi\rho)]$.

1. Periodicity on S^1 . The fields and the fermionic parameters ϵ_r , denoted generically by ϕ , are required to be periodic on the circle S^1 . That is,

$$\phi(x^{\bar{\mu}}, x^5) = \phi(x^{\bar{\mu}}, x^5 + 2\pi\rho).$$
(2.6)

2. \mathbb{Z}_2 parity assignments. The bosonic field components

$$\Phi = e_{\mu}{}^{\bar{a}}, e_{5}{}^{\dot{5}}, A_{5}^{i}, B_{5}, \quad \Theta = e_{\bar{\mu}}{}^{\dot{5}}, e_{5}{}^{\bar{a}}, A_{\bar{\mu}}^{i}, B_{\bar{\mu}}$$

are chosen to satisfy

$$\Phi(x^{\mu}, x^{5}) = +\Phi(x^{\mu}, -x^{5}), \quad \Theta(x^{\mu}, x^{5}) = -\Theta(x^{\mu}, -x^{5}).$$
(2.7)

That is, the Φ components are \mathbb{Z}_2 -even and the Θ components are \mathbb{Z}_2 -odd. For the gravitini, we require

$$\Gamma^{5} \psi_{\bar{\mu}r}(x^{\bar{\mu}}, x^{5}) = + \psi_{\bar{\mu}r}(x^{\bar{\mu}}, -x^{5}),$$

$$\Gamma^{5} \psi_{5r}(x^{\bar{\mu}}, x^{5}) = - \psi_{5r}(x^{\bar{\mu}}, -x^{5}).$$
(2.8)

Finally, the fermionic parameters ϵ_r are required to satisfy

$$\Gamma^{5} \epsilon_{r}(x^{\bar{\mu}}, x^{5}) = + \epsilon_{r}(x^{\mu}, -x^{5}).$$
(2.9)

These conditions imply that the \mathbb{Z}_2 -odd quantities vanish at the orbifold fixed points. It is straightforward to check that S_{bulk} is \mathbb{Z}_2 -even and that the transformations (2.3) are consistent with the \mathbb{Z}_2 parity assignments.

3 Calculation of $\delta_{\epsilon}S_{bulk}$

The D = 5 Chern-Simons AdS supergravity action is invariant (up to a boundary term) under its fermionic symmetry transformations. In Section 2, we modified this action and its fermionic transformations to obtain a \mathbb{Z}_2 -invariant bulk theory. Due to the signum functions introduced by the modifications, $\delta_{\epsilon}S_{bulk}$ contains terms which have no counterpart in the unmodified theory. More specifically, the extra terms arise from ∂_5 acting on the signum functions to yield delta functions. Such 'delta function' contributions to $\delta_{\epsilon}S_{bulk}$ can potentially spoil the fermionic symmetry only at the \mathbb{Z}_2 fixed points. Thus, S_{bulk} is invariant under its fermionic transformations everywhere except perhaps at the \mathbb{Z}_2 fixed points. In this section, we will calculate $\delta_{\epsilon}S_{bulk}$. For our metric and \mathbb{Z}_2 parity assignments, the uncancelled variation $\delta_{\epsilon}S_{bulk}$ arises from the variation of the 4-Fermi terms. The 4-Fermi terms are

$$S_{\psi^{4}} = \frac{3i}{4} \int K \left(\bar{\psi}_{\alpha}^{r} \psi_{s}^{\alpha} \bar{\psi}_{\beta}^{s} \nabla \psi_{r}^{\beta} + \bar{\psi}_{\alpha}^{s} \bar{\psi}_{\beta}^{r} \nabla \psi_{r}^{\alpha} \right) + c.c.$$

$$= \frac{3i}{2} \int K \bar{\psi}_{\alpha}^{r} \psi_{s}^{\alpha} \bar{\psi}_{\beta}^{s} \nabla \psi_{r}^{\beta} + c.c.$$

$$= \frac{3i}{2} \int K \left(\bar{\psi}^{r} \psi_{s} \right) \left(\bar{\psi}^{s} \nabla \psi_{r} \right) + c.c.$$

$$= \frac{3i}{2} \frac{1}{5!} \int d^{5}x \, \varepsilon^{\mu\nu\rho\sigma\lambda} K \left(\bar{\psi}_{\mu}^{r} \psi_{\nu s} \right) \left(\bar{\psi}_{\rho}^{s} \nabla_{\sigma} \psi_{\lambda r} \right) + c.c.$$

$$= \frac{3i}{2} \frac{1}{4!} \int d^{5}x \, K \left[\varepsilon^{5\bar{\nu}\bar{\rho}\bar{\sigma}\bar{\lambda}} \left(\bar{\psi}_{5}^{r} \psi_{\bar{\nu}s} \right) \left(\bar{\psi}_{\bar{\rho}}^{s} \nabla_{\bar{\sigma}} \psi_{\bar{\lambda}r} \right) + \varepsilon^{\bar{\mu}\bar{\nu}\bar{\rho}\bar{\sigma}\bar{\lambda}} \left(\bar{\psi}_{\bar{\mu}}^{r} \psi_{5s} \right) \left(\bar{\psi}_{\bar{\rho}}^{s} \nabla_{\bar{\sigma}} \psi_{\bar{\lambda}r} \right) \right.$$

$$\left. + \varepsilon^{\bar{\mu}\bar{\nu}\bar{\rho}\bar{\sigma}\bar{5}} \left(\bar{\psi}_{\bar{\mu}}^{r} \psi_{\bar{\nu}s} \right) \left(\bar{\psi}_{\bar{\rho}}^{s} \nabla_{\bar{\sigma}} \psi_{\bar{\lambda}r} \right) + \varepsilon^{\bar{\mu}\bar{\nu}\bar{\rho}\bar{5}\bar{\lambda}} \left(\bar{\psi}_{\bar{\rho}}^{r} \psi_{\bar{\nu}s} \right) \left(\bar{\psi}_{\bar{\rho}}^{s} \nabla_{\bar{\sigma}} \psi_{\bar{\lambda}r} \right) \right.$$

$$\left. + \varepsilon^{\bar{\mu}\bar{\nu}\bar{\rho}\bar{\sigma}\bar{5}} \left(\bar{\psi}_{\bar{\mu}}^{r} \psi_{\bar{\nu}s} \right) \left(\bar{\psi}_{\bar{\rho}}^{s} \nabla_{\bar{\sigma}} \psi_{5r} \right) \right] + c.c.$$

$$(3.1)$$

Let us now compute $\delta_{\epsilon}S_{bulk}$ by applying δ_{ϵ} to (3.1) and dropping all terms which contribute no delta functions. For our metric and \mathbb{Z}_2 parity assignments, we can drop all but

- 1. The ∂_{μ} part of ∇_{μ} .
- 2. The $-\partial_{\mu}\epsilon_r$ part of $\delta_{\epsilon} = -\nabla_{\mu}\epsilon_r$.

The only contributing terms are thus contained in the expression

$$Q \equiv -\frac{3i}{2} \frac{1}{4!} \int d^5 x \, K \left\{ \varepsilon^{5\bar{\nu}\bar{\rho}\bar{\sigma}\bar{\lambda}} \left(\partial_5 \bar{\epsilon}^r \psi_{\bar{\nu}s} \right) \left(\bar{\psi}^s_{\bar{\rho}} \partial_{\bar{\sigma}} \psi_{\bar{\lambda}r} \right) \right. \\ \left. + \varepsilon^{\bar{\mu}5\bar{\rho}\bar{\sigma}\bar{\lambda}} \left(\bar{\psi}^r_{\bar{\mu}} \partial_5 \epsilon_s \right) \left(\bar{\psi}^s_{\bar{\rho}} \partial_{\bar{\sigma}} \psi_{\bar{\lambda}r} \right) + \varepsilon^{\bar{\mu}\bar{\nu}5\bar{\sigma}\bar{\lambda}} \left(\bar{\psi}^r_{\bar{\mu}} \psi_{\bar{\nu}s} \right) \left(\partial_5 \bar{\epsilon}^s \partial_{\bar{\sigma}} \psi_{\bar{\lambda}r} \right) \right. \\ \left. + \varepsilon^{\bar{\mu}\bar{\nu}\bar{\rho}5\bar{\lambda}} \left[\left(\partial_{\bar{\mu}} \bar{\epsilon}^r \psi_{\bar{\nu}s} \right) \left(\bar{\psi}^s_{\bar{\rho}} \partial_5 \psi_{\bar{\lambda}r} \right) + \left(\bar{\psi}^r_{\bar{\mu}} \partial_{\bar{\nu}} \epsilon_s \right) \left(\bar{\psi}^s_{\bar{\rho}} \partial_5 \psi_{\bar{\lambda}r} \right) \right. \\ \left. + \left(\bar{\psi}^r_{\bar{\mu}} \psi_{\bar{\nu}s} \right) \left(\partial_{\bar{\rho}} \bar{\epsilon}^s \partial_5 \psi_{\bar{\lambda}r} \right) + \left(\bar{\psi}^r_{\bar{\mu}} \psi_{\bar{\nu}s} \right) \left(\bar{\psi}^s_{\bar{\rho}} \partial_5 \partial_{\bar{\lambda}} \epsilon_r \right) \right] \\ \left. + \varepsilon^{\bar{\mu}\bar{\nu}\bar{\rho}\bar{\sigma}5} \left(\bar{\psi}^r_{\bar{\mu}} \psi_{\bar{\nu}s} \right) \left(\bar{\psi}^s_{\bar{\rho}} \partial_{\bar{\sigma}} \partial_5 \epsilon_r \right) \right\} + c.c. \tag{3.2}$$

More specifically, the delta function terms contained in Q are obtained by integrating by parts and keeping only the terms in which ∂_5 acts on K. Thus,

$$\delta_{\epsilon}S_{bulk} = \frac{3i}{2}\frac{1}{4!}\int d^{5}x\,\partial_{5}K\left\{\varepsilon^{5\bar{\nu}\bar{\rho}\bar{\sigma}\bar{\lambda}}\left(\bar{\epsilon}^{r}\psi_{\bar{\nu}s}\right)\left(\bar{\psi}_{\bar{\rho}}^{s}\partial_{\bar{\sigma}}\psi_{\bar{\lambda}r}\right)\right.\\ \left.+\varepsilon^{\bar{\mu}\bar{\nu}\bar{\rho}\bar{\sigma}\bar{\lambda}}\left(\bar{\psi}_{\bar{\mu}}^{r}\epsilon_{s}\right)\left(\bar{\psi}_{\bar{\rho}}^{s}\partial_{\bar{\sigma}}\psi_{\bar{\lambda}r}\right) + \varepsilon^{\bar{\mu}\bar{\nu}5\bar{\sigma}\bar{\lambda}}\left(\bar{\psi}_{\bar{\mu}}^{r}\psi_{\bar{\nu}s}\right)\left(\bar{\epsilon}^{s}\partial_{\bar{\sigma}}\psi_{\bar{\lambda}r}\right)\right.\\ \left.+\varepsilon^{\bar{\mu}\bar{\nu}\bar{\rho}\bar{5}\bar{\lambda}}\left[\left(\partial_{\bar{\mu}}\bar{\epsilon}^{r}\psi_{\bar{\nu}s}\right)\left(\bar{\psi}_{\bar{\rho}}^{s}\psi_{\bar{\lambda}r}\right) + \left(\bar{\psi}_{\bar{\mu}}^{r}\partial_{\bar{\nu}}\epsilon_{s}\right)\left(\bar{\psi}_{\bar{\rho}}^{s}\psi_{\bar{\lambda}r}\right)\right.\\ \left.+\left(\bar{\psi}_{\bar{\mu}}^{r}\psi_{\bar{\nu}s}\right)\left(\partial_{\bar{\rho}}\bar{\epsilon}^{s}\psi_{\bar{\lambda}r}\right) + \left(\bar{\psi}_{\bar{\mu}}^{r}\psi_{\bar{\nu}s}\right)\left(\bar{\psi}_{\bar{\rho}}^{s}\partial_{\bar{\lambda}}\epsilon_{r}\right)\right]\\ \left.+\varepsilon^{\bar{\mu}\bar{\nu}\bar{\rho}\bar{\sigma}\bar{5}}\left(\bar{\psi}_{\bar{\mu}}^{r}\psi_{\bar{\nu}s}\right)\left(\bar{\psi}_{\bar{\rho}}^{s}\partial_{\bar{\sigma}}\epsilon_{r}\right)\right\} + c.c., \tag{3.3}$$

where

$$\partial_5 K = 2k \left[\delta(x^5) - \delta(x^5 - \pi \rho) \right]. \tag{3.4}$$

4 Modified \mathbb{Z}_2 -invariant bulk theory

The result (3.3) for $\delta_{\epsilon}S_{bulk}$ demonstrates that S_{bulk} is not invariant under the fermionic transformations δ_{ϵ} . In this section, we will modify S_{bulk} and δ_{ϵ} by replacing the $adS_5 \times su(N) \times u(1)$ covariant derivative ∇ with $\widetilde{\nabla}$, where

$$\widetilde{\nabla}_{\sigma}\psi_{\lambda r} \equiv \nabla_{\sigma}\psi_{\lambda r} + 2\delta_{\sigma}^{5}\delta_{\lambda}^{\bar{\lambda}} \left[\delta(x^{5}) - \delta(x^{5} - \pi\rho)\right] \operatorname{sgn}(x^{5})\Gamma_{5}\psi_{\bar{\lambda}r},$$

$$\widetilde{\nabla}_{\sigma}\epsilon_{r} \equiv \nabla_{\sigma}\epsilon_{r} + 2\delta_{\sigma}^{5} \left[\delta(x^{5}) - \delta(x^{5} - \pi\rho)\right] \operatorname{sgn}(x^{5})\Gamma_{5}\epsilon_{r}.$$
(4.1)

We will show that the modified bulk action

$$\tilde{S}_{bulk} \equiv S_{bulk} \left(\nabla \to \widetilde{\nabla} \right) \equiv S_{bulk} + \Delta S_{bulk} \tag{4.2}$$

is invariant under the modified transformations

$$\tilde{\delta}_{\epsilon} \equiv \delta_{\epsilon} \left(\nabla \to \widetilde{\nabla} \right) \equiv \delta_{\epsilon} + \Delta \delta_{\epsilon}. \tag{4.3}$$

That is, we will show that

$$\tilde{\delta}_{\epsilon}\tilde{S}_{bulk} = \delta_{\epsilon}S_{bulk} + (\Delta\delta_{\epsilon})S_{bulk} + \tilde{\delta}_{\epsilon}(\Delta S_{bulk})$$
(4.4)

vanishes. It is straightforward to check that \tilde{S}_{bulk} is \mathbb{Z}_2 -invariant and the transformations $\tilde{\delta}_{\epsilon}$ are consistent with our \mathbb{Z}_2 parity assignments.

We begin by computing $(\Delta \delta_{\epsilon}) S_{bulk}$. For our metric and \mathbb{Z}_2 parity assignments, the only part of S_{bulk} which is not invariant under $\Delta \delta_{\epsilon}$ is S_{ψ^4} (given by (3.1)). Note that

$$(\Delta\delta_{\epsilon})\psi_{\lambda r} = -2\delta_{\lambda}^{5} \left[\delta(x^{5}) - \delta(x^{5} - \pi\rho)\right] \operatorname{sgn}(x^{5})\Gamma_{5}\epsilon_{r}.$$
(4.5)

Thus, after using $K \operatorname{sgn}(x^5) = k$, (2.9), and (3.4), we obtain

$$(\Delta\delta_{\epsilon}) S_{bulk} = -\frac{3i}{2} \frac{1}{4!} \int d^5 x \, \partial_5 K \left[\varepsilon^{5\bar{\nu}\bar{\rho}\bar{\sigma}\bar{\lambda}} \left(\bar{\epsilon}^r \psi_{\bar{\nu}s} \right) \left(\bar{\psi}^s_{\bar{\rho}} \partial_{\bar{\sigma}} \psi_{\bar{\lambda}r} \right) \right. \\ \left. + \varepsilon^{\bar{\mu}\bar{\nu}\bar{\rho}\bar{\sigma}\bar{\lambda}} \left(\bar{\psi}^r_{\bar{\mu}} \epsilon_s \right) \left(\bar{\psi}^s_{\bar{\rho}} \partial_{\bar{\sigma}} \psi_{\bar{\lambda}r} \right) + \varepsilon^{\bar{\mu}\bar{\nu}5\bar{\sigma}\bar{\lambda}} \left(\bar{\psi}^r_{\bar{\mu}} \psi_{\bar{\nu}s} \right) \left(\bar{\epsilon}^s \partial_{\bar{\sigma}} \psi_{\bar{\lambda}r} \right) \right. \\ \left. + \varepsilon^{\bar{\mu}\bar{\nu}\bar{\rho}\bar{\sigma}5} \left(\bar{\psi}^r_{\bar{\mu}} \psi_{\bar{\nu}s} \right) \left(\bar{\psi}^s_{\bar{\rho}} \partial_{\bar{\sigma}} \epsilon_r \right) \right] + c.c.$$

$$(4.6)$$

Now, let us compute $\tilde{\delta}_{\epsilon}(\Delta S_{bulk})$. For our metric and \mathbb{Z}_2 parity assignments, the only part of S_{bulk} which is changed by the replacement $\nabla \to \widetilde{\nabla}$ is S_{ψ^4} . After using $K \operatorname{sgn}(x^5) = k$, (2.8), and (3.4), we obtain

$$\Delta S_{bulk} = \frac{3i}{2} \frac{1}{4!} \int d^5 x \, \partial_5 K \, \varepsilon^{\bar{\mu}\bar{\nu}\bar{\rho}5\bar{\lambda}} \left(\bar{\psi}^r_{\bar{\mu}}\psi_{\bar{\nu}s} \right) \left(\bar{\psi}^s_{\bar{\rho}}\psi_{\bar{\lambda}r} \right) + c.c. \tag{4.7}$$

Applying $\tilde{\delta}_{\epsilon}$ to (4.7) yields

$$\tilde{\delta}_{\epsilon} \left(\Delta S_{bulk} \right) = -\frac{3i}{2} \frac{1}{4!} \int d^5 x \, \partial_5 K \, \varepsilon^{\bar{\mu}\bar{\nu}\bar{\rho}5\bar{\lambda}} \left[\left(\partial_{\bar{\mu}} \bar{\epsilon}^r \psi_{\bar{\nu}s} \right) \left(\bar{\psi}^s_{\bar{\rho}} \psi_{\bar{\lambda}r} \right) + \left(\bar{\psi}^r_{\bar{\mu}} \partial_{\bar{\nu}} \epsilon_s \right) \left(\bar{\psi}^s_{\bar{\rho}} \psi_{\bar{\lambda}r} \right) \\
+ \left(\bar{\psi}^r_{\bar{\mu}} \psi_{\bar{\nu}s} \right) \left(\partial_{\bar{\rho}} \bar{\epsilon}^s \psi_{\bar{\lambda}r} \right) + \left(\bar{\psi}^r_{\bar{\mu}} \psi_{\bar{\nu}s} \right) \left(\bar{\psi}^s_{\bar{\rho}} \partial_{\bar{\lambda}} \epsilon_r \right) \right] + c.c.$$
(4.8)

Using the results (3.3), (4.6), and (4.8) in (4.4) yields

$$\tilde{\delta}_{\epsilon}\tilde{S}_{bulk} = 0. \tag{4.9}$$

5 Brane action

To complete our model, we add the brane action S_{brane} . In the absence of particle excitations, the brane action consists of brane tensions. That is,

$$S_{brane} = -\int d^5 x \, e^{(4)} \left[\mathcal{T}^{(0)} \delta(x^5) + \mathcal{T}^{(\pi\rho)} \delta(x^5 - \pi\rho) \right] + \text{excitations}, \tag{5.1}$$

where $e^{(4)} \equiv \det(e_{\bar{\mu}}{}^{\bar{a}})$. As discussed in Section 2, \mathbb{Z}_2 -odd quantities vanish at the \mathbb{Z}_2 fixed points. Thus, it is clear that S_{brane} is \mathbb{Z}_2 -even. Further discussion of 3-brane actions can be found in [11].

6 Connection with original RS model

In this section, we will show that our total action $S = \tilde{S}_{bulk} + S_{brane}$ reduces to the original Randall-Sundrum model [10] when \tilde{S}_{bulk} is restricted to its gravitational sector.

The gravitational sector of \tilde{S}_{bulk} is S_{grav} , given by the first equation of (2.2). S_{grav} consists of three terms:

- 1. The 'Gauss-Bonnet' term $\int \frac{1}{8} K \varepsilon_{abcde} R^{ab} R^{cd} e^e / L$.
- 2. The 'Einstein-Hilbert' term $\int \frac{1}{8} \cdot \frac{2}{3} K \varepsilon_{abcde} R^{ab} e^c e^d e^e / L^3$.
- 3. The 'cosmological constant' term $\int \frac{1}{8} \cdot \frac{1}{5} K \varepsilon_{abcde} e^a e^b e^c e^d e^e / L^5$.

For our metric, the first term can be expressed as a linear combination of the other two. Summing the three terms yields an effective Einstein-Hilbert term and an effective cosmological constant term. To demonstrate this explicitly, let us evaluate S_{grav} for our metric. Using the results in the Appendix, we obtain

$$\varepsilon_{abcde} R^{ab} R^{cd} e^{e} = d^{5} x \ e \left(-\frac{120}{\ell^{4}} + \frac{192}{\ell^{3}} \left[\delta(x^{5}) - \delta(x^{5} - \pi\rho) \right] \right),$$

$$\varepsilon_{abcde} R^{ab} e^{c} e^{d} e^{e} = d^{5} x \ e \ (-6R) \ ,$$

$$\varepsilon_{abcde} e^{a} e^{b} e^{c} e^{d} e^{e} = d^{5} x \ e \ (-5!) \ ,$$
(6.1)

where $e \equiv \det(e_{\mu}{}^{a})$. Thus,

$$S_{grav} = \int d^5 x \ e \ \frac{1}{8} \left\{ \frac{k}{\ell} \left(-\frac{120}{\ell^4} + \frac{192}{\ell^3} \left[\delta(x^5) - \delta(x^5 - \pi\rho) \right] \right) + \frac{2k}{3\ell^3} \left(-6R \right) + \frac{k}{5\ell^5} \left(-5! \right) \right\}$$

$$= \int d^5 x \ e \ \frac{k}{\ell^3} \left\{ -\frac{15}{\ell^2} + \frac{24}{\ell} \left[\delta(x^5) - \delta(x^5 - \pi\rho) \right] - \frac{1}{2}R - \frac{3}{\ell^2} \right\}$$

$$= \int d^5 x \ e \ \frac{k}{\ell^3} \left\{ \frac{3}{2} \left(-\frac{20}{\ell^2} + \frac{16}{\ell} \left[\delta(x^5) - \delta(x^5 - \pi\rho) \right] \right) - \frac{1}{2}R + \frac{12}{\ell^2} \right\}$$

$$= \int d^5 x \ e \ \frac{k}{\ell^3} \left(R + \frac{12}{\ell^2} \right)$$

$$= \int d^5 x \ e \ (2M^3R - \Lambda) , \qquad (6.2)$$

where M is the five-dimensional gravitational mass scale⁷, Λ is the bulk cosmological constant, and

$$M^{3} = \frac{k}{2\ell^{3}}, \quad \Lambda = -\frac{24M^{3}}{\ell^{2}}.$$
 (6.3)

Combining the result (6.2) with (5.1), we obtain the action of the original Randall-Sundrum model. It is shown in [10] that the five-dimensional vacuum Einstein's equations for this system,

$$R_{\mu\nu} - \frac{1}{2}g_{\mu\nu}R = -\frac{1}{4M^3} \left\{ g_{\mu\nu}\Lambda + \frac{e^{(4)}}{e} \delta^{\bar{\mu}}_{\mu} \delta^{\bar{\nu}}_{\nu} g_{\bar{\mu}\bar{\nu}} \left[\mathcal{T}^{(0)}\delta(x^5) + \mathcal{T}^{(\pi\rho)}\delta(x^5 - \pi\rho) \right] \right\}, \quad (6.4)$$

are solved by our metric provided that the relations

$$\mathcal{T}^{(0)} = -\mathcal{T}^{(\pi\rho)} = 24M^3/\ell, \quad \Lambda = -24M^3/\ell^2$$
 (6.5)

are satisfied.

7 Killing spinors

In this section, we will solve the Killing spinor equations for a purely bosonic configuration with vanishing su(N) and u(1) gauge fields. In this case, the Killing spinor equations reduce to

$$0 = \delta_{\epsilon} \psi_{\bar{\mu}r} = -\partial_{\bar{\mu}} \epsilon_r - \frac{1}{2} \frac{\mathfrak{a}'}{\mathfrak{a}} \Gamma_{\bar{\mu}} \left(\Gamma_{5} - 1 \right) \epsilon_r,$$

$$0 = \delta_{\epsilon} \psi_{5r} = -\partial_{5} \epsilon_r + \frac{1}{2} \frac{\mathfrak{a}'}{\mathfrak{a}} \Gamma_{5} \epsilon_r - 2 \left[\delta(x^5) - \delta(x^5 - \pi\rho) \right] \operatorname{sgn}(x^5) \Gamma_{5} \epsilon_r.$$
(7.1)

 $^{^{7}}M$ is related to the four-dimensional gravitational mass scale $M_{(4)} = 2.43 \times 10^{18}$ GeV by $M_{(4)}^{2} = M^{3} \int_{-\pi\rho}^{+\pi\rho} dx^{5} \mathfrak{a}^{2}(x^{5}) = M^{3} \ell \left[1 - \exp(-2\pi\rho/\ell)\right]$. The effective mass scales on the 3-branes at $x^{5} = 0$ and $x^{5} = \pi\rho$ are respectively $M_{(4)}$ and $M_{(4)}e^{-\pi\rho/\ell}$. If the Standard Model fields live on the 3-brane at $x^{5} = \pi\rho$, then $M_{(4)}e^{-\pi\rho/\ell}$ can be associated with the electroweak scale.

To solve these equations, split ϵ_r into \mathbb{Z}_2 -even (ϵ_r^+) and \mathbb{Z}_2 -odd (ϵ_r^-) pieces:

$$\epsilon_r = \epsilon_r^+ + \epsilon_r^-,\tag{7.2}$$

where

$$\epsilon^{\pm} \equiv \frac{1}{2} \left(\epsilon_r \pm \Gamma_{5} \epsilon_r \right) = \pm \Gamma_{5} \epsilon_r^{\pm}.$$
(7.3)

We obtain the following system of equations:

$$\partial_{\bar{\mu}}\epsilon_{r}^{+} = -\left(\mathfrak{a}'/\mathfrak{a}\right)\Gamma_{\bar{\mu}}\Gamma_{5}\epsilon_{r}^{-},$$

$$\partial_{\bar{\mu}}\epsilon_{r}^{-} = 0,$$

$$\partial_{5}\epsilon_{r}^{+} = +\frac{1}{2}\left(\mathfrak{a}'/\mathfrak{a}\right)\epsilon_{r}^{+} - 2\left[\delta(x^{5}) - \delta(x^{5} - \pi\rho)\right]\operatorname{sgn}(x^{5})\epsilon_{r}^{+},$$

$$\partial_{5}\epsilon_{r}^{-} = -\frac{1}{2}\left(\mathfrak{a}'/\mathfrak{a}\right)\epsilon_{r}^{-} + 2\left[\delta(x^{5}) - \delta(x^{5} - \pi\rho)\right]\operatorname{sgn}(x^{5})\epsilon_{r}^{-}.$$
(7.4)

These equations are solved by

$$\begin{aligned} \epsilon_r^+ &= \mathfrak{a}^{1/2} \left[- \left(\mathfrak{a}'/\mathfrak{a}^2 \right) x^{\bar{\mu}} \Gamma_{\bar{\mu}} \Gamma_{5} \operatorname{sgn}(x^5) \chi_r^{-(0)} + \chi_r^{+(0)} \right] \\ &= \mathfrak{a}^{1/2} \left[\left(1/\ell \right) x^{\bar{\mu}} \delta_{\bar{\mu}}{}^{\bar{a}} \Gamma_{\bar{a}} \Gamma_{5} \chi_r^{-(0)} + \chi_r^{+(0)} \right], \\ \epsilon_r^- &= \mathfrak{a}^{-1/2} \operatorname{sgn}(x^5) \chi_r^{-(0)}, \end{aligned}$$
(7.5)

where $\chi_r^{+(0)}$ and $\chi_r^{-(0)}$ are constant (projected) spinors.⁸ Thus, our solution for the Killing spinors is

$$\epsilon_r = \mathfrak{a}^{1/2} \chi_r^{+(0)} + \mathfrak{a}^{-1/2} \operatorname{sgn}(x^5) \left(1 - \frac{\mathfrak{a}'}{\mathfrak{a}} x^{\bar{\mu}} \Gamma_{\bar{\mu}} \Gamma_{\bar{5}} \right) \chi_r^{-(0)}.$$
(7.6)

8 Conclusion

We have constructed a Randall-Sundrum scenario from D=5 Chern-Simons AdS supergravity. Our total action $S = \tilde{S}_{bulk} + S_{brane}$ is \mathbb{Z}_2 -invariant. \tilde{S}_{bulk} is invariant under the fermionic transformations $\tilde{\delta}_{\epsilon}$. However,

$$\tilde{\delta}_{\epsilon}S_{brane} = -\int d^5x \,\tilde{\delta}_{\epsilon}e^{(4)} \left[\mathcal{T}^{(0)}\delta(x^5) + \mathcal{T}^{(\pi\rho)}\delta(x^5 - \pi\rho) \right] + \cdots, \qquad (8.1)$$

where

$$\tilde{\delta}_{\epsilon} e^{(4)} = e^{(4)} \left[-\frac{1}{2} \left(\bar{\psi}^r_{\bar{\mu}} \Gamma^{\bar{\mu}} \epsilon_r - \bar{\epsilon}^r \Gamma^{\bar{\mu}} \psi_{\bar{\mu}r} \right) \right].$$
(8.2)

Thus, the fermionic symmetry is broken by S_{brane} . Nevertheless, the Killing spinors of Section 7 are globally defined.

Our model reduces to the original Randall-Sundrum model [10] when \tilde{S}_{bulk} is restricted to its gravitational sector. The original Randall-Sundrum model involves the fine-tuning relations

$$\mathcal{T}^{(0)} = -\mathcal{T}^{(\pi\rho)} = 24M^3/\ell, \quad \Lambda = -24M^3/\ell^2.$$

⁸It is straightforward to check that (7.5) satisfies the first, second, and fourth equations of (7.4). There is, however, a subtlety in checking that (7.5) satisfies the third equation of (7.4). Unlike ϵ_r^- , ϵ_r^+ is a smooth function of x^5 . Thus, the second term on the right side of the third equation of (7.4) contributes nothing.

Randall-Sundrum scenarios constructed from standard D=5 supergravity theories yield these relations (up to an overall normalization factor) as a consequence of local supersymmetry (some examples are given in [12]). In our case, the relation $\Lambda = -24M^3/\ell^2$ follows from our metric choice. We do not obtain the relations $\mathcal{T}^{(0)} = -\mathcal{T}^{(\pi\rho)} = 24M^3/\ell$ as a consequence of local fermionic symmetry. These are fine-tuning relations in our model.

A Appendix

In this Appendix, we work out the fünfbein, spin connection, curvature 2-form components, Ricci tensor, and Ricci scalar for our metric (1.1). For the fünfbein, we obtain

$$e_{\bar{\mu}}{}^{\bar{a}} = \mathfrak{a} \delta_{\bar{\mu}}{}^{\bar{a}}, \quad e_{\bar{\mu}\bar{a}} = e_{\bar{\mu}}{}^{\bar{b}} \eta_{\bar{b}\bar{a}}, \quad e^{\bar{\mu}\bar{a}} = g^{\bar{\mu}\bar{\nu}} e_{\bar{\nu}}{}^{\bar{a}},$$

$$e_{\bar{a}}{}^{\bar{\mu}} = \mathfrak{a}^{-1} \delta_{\bar{a}}{}^{\bar{\mu}}, \quad e_{\bar{a}\bar{\mu}} = e_{\bar{a}}{}^{\bar{\nu}} g_{\bar{\nu}\bar{\mu}}, \quad e^{\bar{a}\bar{\mu}} = \eta^{\bar{a}\bar{b}} e_{\bar{b}}{}^{\bar{\mu}},$$

$$e_{5}{}^{\dot{5}} = e_{55} = e^{55} = 1, \quad e_{5}{}^{5} = e_{55} = e^{55} = 1.$$
(A.1)

Our conventions for the spin connection, curvature 2-form components, Ricci tensor, and Ricci scalar are respectively

$$\omega_{\mu}{}^{ab} = \frac{1}{2} e^{\nu a} (\partial_{\mu} e_{\nu}{}^{b} - \partial_{\nu} e_{\mu}{}^{b}) - \frac{1}{2} e^{\nu b} (\partial_{\mu} e_{\nu}{}^{a} - \partial_{\nu} e_{\mu}{}^{a}) - \frac{1}{2} e^{\rho a} e^{\sigma b} (\partial_{\rho} e_{\sigma c} - \partial_{\sigma} e_{\rho c}) e^{c}{}_{\mu}, R_{\mu\nu}{}^{ab} = \partial_{\mu} \omega_{\nu}{}^{ab} - \partial_{\nu} \omega_{\mu}{}^{ab} + \omega_{\mu}{}^{ac} \omega_{\nu c}{}^{b} - \omega_{\nu}{}^{ac} \omega_{\mu c}{}^{b}, R_{\nu\sigma} = R_{\mu\nu}{}^{ab} e_{a}{}^{\mu} e_{b\sigma}, \quad R = e_{a}{}^{\mu} e_{b}{}^{\nu} R_{\mu\nu}{}^{ab}.$$

For the metric (1.1), the nonzero quantities here are

$$\omega_{\bar{\mu}}{}^{\bar{a}\dot{5}} = -\omega_{\bar{\mu}}{}^{\dot{5}\bar{a}} = \mathfrak{a}'\delta_{\bar{\mu}}{}^{\bar{a}} = -e_{\bar{\mu}}{}^{\bar{a}}/L, \qquad (A.2)$$

$$R_{\bar{\mu}\bar{\nu}}{}^{\bar{a}\bar{b}} = -\mathfrak{a}'^{2} (\delta_{\bar{\mu}}{}^{\bar{a}} \delta_{\bar{\nu}}{}^{\bar{b}} - \delta_{\bar{\mu}}{}^{\bar{b}} \delta_{\bar{\nu}}{}^{\bar{a}}) = -(e_{\bar{\mu}}{}^{\bar{a}} e_{\bar{\nu}}{}^{\bar{b}} - e_{\bar{\mu}}{}^{\bar{b}} e_{\bar{\nu}}{}^{\bar{a}})/\ell^{2},$$

$$R_{5\bar{\mu}}{}^{\bar{a}\dot{5}} = \mathfrak{a}'' \delta_{\bar{\mu}}{}^{\bar{a}} = e_{\bar{\mu}}{}^{\bar{a}} \left\{ 1/\ell^{2} - 2[\delta(x^{5}) - \delta(x^{5} - \pi\rho)]/\ell \right\},$$
(A.3)

$$R_{\bar{\mu}\bar{\nu}} = -(\mathfrak{a}\mathfrak{a}'' + 3\mathfrak{a}'^2)\eta_{\bar{\mu}\bar{\nu}} = -\left\{4/\ell^2 - 2[\delta(x^5) - \delta(x^5 - \pi\rho)]/\ell\right\}g_{\bar{\mu}\bar{\nu}},$$

$$R_{55} = -4\mathfrak{a}^{-1}\mathfrak{a}'' = -\left\{4/\ell^2 - 8[\delta(x^5) - \delta(x^5 - \pi\rho)]/\ell\right\},$$
(A.4)

$$R = -8\mathfrak{a}^{-1}\mathfrak{a}'' - 12\mathfrak{a}^{-2}\mathfrak{a}'^2 = -20/\ell^2 + 16[\delta(x^5) - \delta(x^5 - \pi\rho)]/\ell,$$
(A.5)

and those related to (A.3) by $R_{\mu\nu}{}^{ab} = -R_{\nu\mu}{}^{ab} = -R_{\mu\nu}{}^{ba}$. The prime symbol \prime denotes partial differentiation with respect to x^5 .

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