

NOTAS DE FISICA

VOLUME II

Nº 4

SHELL EFFECT ON PHOTONUCLEAR REACTIONS

by

J. GOLDEMBERG and J. LEITE LOPES

CENTRO BRASILEIRO DE PESQUISAS FISICAS

Av. Wenceslau Braz, 71

RIO DE JANEIRO

1955

SHELL EFFECT ON PHOTONUCLEAR REACTIONS\*§

J. Goldemberg  
Faculdade de Filosofia, Ciências e Letras  
Universidade de São Paulo  
São Paulo

and

J. Leite Lopes  
Centro Brasileiro de Pesquisas Físicas  
Rio de Janeiro, D.F.

(April 15, 1955)

A previous analysis of some experimental data on photo-nuclear reactions led the authors<sup>1</sup> to point out that the mean-square displacement of nucleons in the nuclear ground state presents evidence in favor of an alpha-particle structure of nuclei. Although the theoretical basis for this conclusion is not free of criticism—since correlations were neglected in the formula for the harmonic mean energy<sup>2</sup> of photon absorption, which we used—this evidence might be investigated experimentally in a more direct way at higher energies.

\* Publication sponsored by the Conselho Nacional de Pesquisas.

§ This paper will be published in Physical Review.

On the other hand, the shell structure of nuclei has been revealed, in the energy region of the photonuclear effect, by Nathans and Halpern<sup>3</sup>. Features of different nuclear models are expected to be presented by nuclear matter. The predominance of one special model over the others depends on the energy and mass number regions and probably on the manner of excitation. It is, therefore, desirable to look for information extracted from experiment and independent, whenever possible, from "a priori" theoretical views and disputable assumptions.

Nathans and Halpern found that the width of the giant resonance, as given by the total neutron yield, is smaller for magic nuclei. This criterion seems to us not very satisfactory since the width might depend on the relative importance of  $(\gamma, 2n)$  and  $(\gamma, np)$  reactions, and also on the threshold of these reactions. A way to avoid this ambiguity is to plot the energy distance from the threshold for  $(\gamma, n)$  reactions to the first point where the cross section is half maximum, against  $A$ . This quantity  $\Delta$  gives a measure of the steepness of rise of the  $(\gamma, n)$  cross-section curve as a function of the energy, as shown in Fig. 1. For magic nuclei and neighbors, one expects  $\Delta$  to be larger than for the others. This is indeed the case, as indicated in Fig. 2 and in Table I, where  $\Delta$  is plotted against  $A$ . The shell structure is clearly exhibited in this curve and the effect is neater for higher mass numbers.

For very light nuclei, there are marked deviations from the expected curve; this might be accounted for by the fact that

for such nuclei the  $(\gamma, n)$  process is not dominant. Thus, where as carbon, which has a complete subshell, has a low  $\Delta$ , nitrogen has an anomalously high  $\Delta$ . Magic nuclei with  $N = 2, 8, 20, 28, 50, 82,$  and  $126$  show their shell structure very clearly in the graph. At  $N = 50$ , we find an anomaly: Nb ( $N = 52$ ), which is not a magic nucleus, shows a value  $\Delta = 5.8$ ; although Rb ( $N = 50$ ), which is magic, has a  $\Delta = 5.4$ , and Mo ( $N = 50$ ), also a magic nucleus, has  $\Delta = 2.1$ . This anomaly suggests that new measurements should be made with these elements.

One of the authors (J.L.L.) wishes to express his appreciation to his colleague Professor M.D.S. Santos for the hospitality extended him last January at the Betatron Laboratory of the University of São Paulo.

1. J. Goldemberg and J. Leite Lopes, Nuovo Cimento 12, 817 (1954).
2. J. S. Lvinger and H. A. Bethe, Phys. Rev. 78, 115 (1950).
3. R. Nathan and J. Halporn, Phys. Rev. 93, 437 (1954).

TABLE I. Values of  $\Delta$  (the difference between the energy at which the  $(\gamma, n)$  cross section is half-maximum and the threshold) or several nuclei.

Element	Z	N	A	Threshold	$E(\frac{1}{2} \sigma_{\max})$	$\Delta$ (Mev)	Reference
C	6	6	12	18.70	20.6	1.9	a, b
N	7	7	14	10.54	19.3	8.8	c
O	18	18	16	15.60	21.8	6.2	c
Na	11	12	23	12.10	15.4	3.3	a
Mg	12	12	24	16.0	18.2	2.2	d
Al	13	14	27	14.0	17.5	3.5	a
Si	14	14	28	16.9	18.8	1.9	e
P	15	16	31	12.4	16.2	3.8	f
S	16	16	32	15.0	18.0	3.0	g
Cl	17	18	35	13.0	16.8	3.8	h
A	18	20	38	10.5	15.0	4.5	i
K	19	20	39	12.5	17.5	5.0	h
Ca	20	20	40	15.8	18.0	2.2	e
Cr	24	26	50	13.4	16.8	3.4	j
V	23	28	51	10.8	15.5	4.7	j
Mn	25	30	55	10.0	14.5	4.5	a
Fe	26	28	54	11.7	17.7	6.0	k
Co	27	32	59	10.2	15.0	4.8	a
Ni	28	30	58	12.0	15.5	3.5	l
Cu	29	34	63	10.6	14.5	3.9	m
Cu	29	36	65	10.2	14.1	3.9	m
Zn	30	34	64	11.6	14.5	2.9	n
As	33	42	75	10.1	13.0	2.9	a
Br	35	46	81	10.0	14.0	4.0	o
Rb	37	50	87	9.3	14.7	5.4	p
Nb	41	52	93	8.7	14.5	5.8	a

TABLE I (Cont.)

Element	Z	N	A	Threshold	$E(-\frac{1}{2}v_{max})$	$\Delta$ (Mev)	Reference
Mo	42	50	92	13.1	15.1	2.1	p
Ag	47	62	109	9.3	13.2	3.9	z
In	49	66	115	9.0	12.5	3.5	a
Sb	51	70	121	9.3	12.4	3.1	m
I	53	74	127	9.1	13.2	4.1	a
La	57	82	139	8.0	12.5	4.5	j
Ta	73	108	181	8.8	11.5	2.7	m
Au	79	118	197	7.9	11.4	4.1	a
Bi	83	126	209	7.4	11.6	4.2	a
U	92	146	238	6.0	9.5	3.5	r

- a Montalbetti, Katz and Goldemberg, Phys. Rev. 91, 659 (1953).
- b Haslam, Johns, and Horsley, Phys. Rev. 82, 270 (1951).
- c Johns, Horslye, Haslam, and Quinton, Phys. Rev. 82, 856 (1951).
- d L. Katz and A. G. W. Cameron, Phys. Rev. 84, 1115 (1951).
- e Summers-Gill, Haslam, and Katz, Can. J. Phys. 31, 70 (1953).
- f L. Katz and A. S. Penfold, Phys. Rev. 81, 815 (1951).
- g Haslam, Summers-Gill, and Crosby, Can. J. Phys. 30, 257 (1952).
- h O. Borello, Ann. acad. Brasil. scienc. (to be published).
- i McPherson, Pederson, and Katz, Can. J. Phys. 32, 593 (1954).
- j J. Goldemberg and L. Katz, Can. J. Phys. 32, 49 (1953).
- k E. Silva, Ann. acad. brasil. scienc. (to be published).
- l Katz, Johns, Baker, Haslam, and Douglas, Phys. Rev. 82, 271 (1951).
- m Johns, Katz, Douglas and Haslam, Phys. Rev. 80, 1062 (1950).
- n S. S. Vilaga, Ann. acad. brasil. scienc. (to be published).
- o Katz, Pease, and Moody, Can. J. Phys. 30, 476 (1952).
- p Katz, Baker and Montalbetti, Can. J. Phys. 31, 250 (1953).
- q B. C. Diven and G. Almy, Phys. Rev. 80, 407 (1950).
- r R. Nathans and J. Halpern, Phys. Rev. 93, 437 (1954).

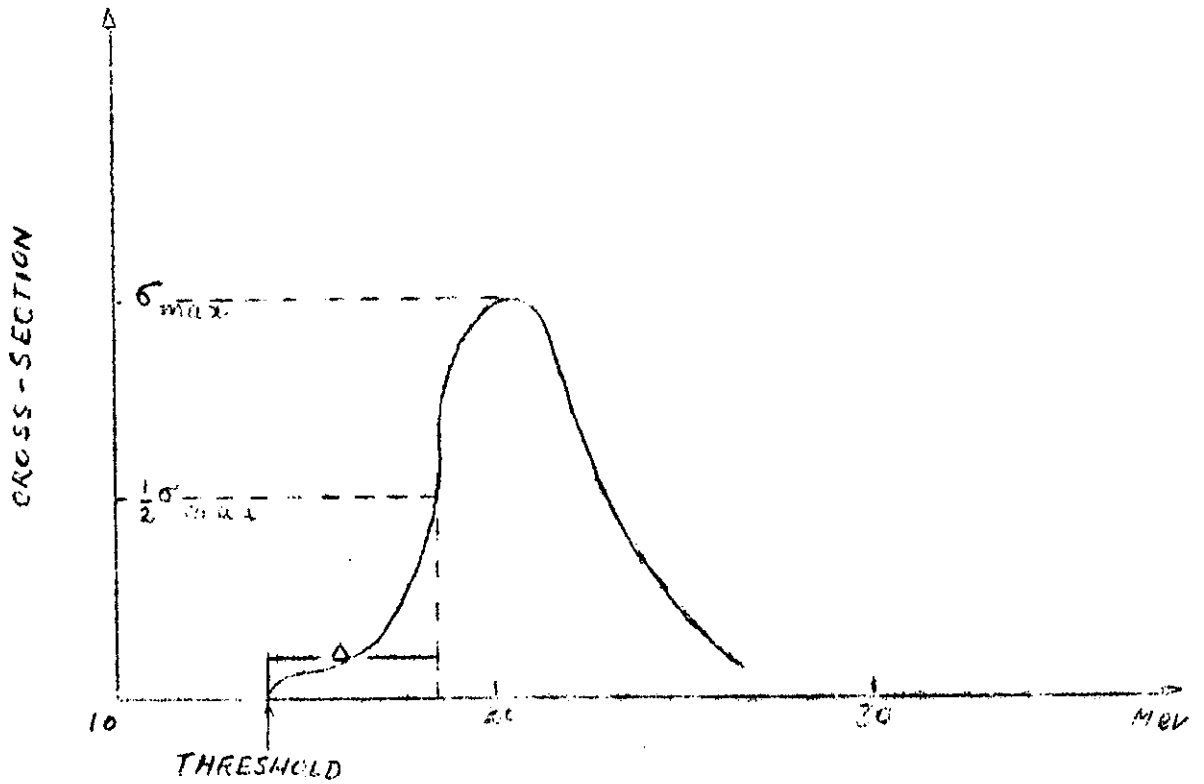
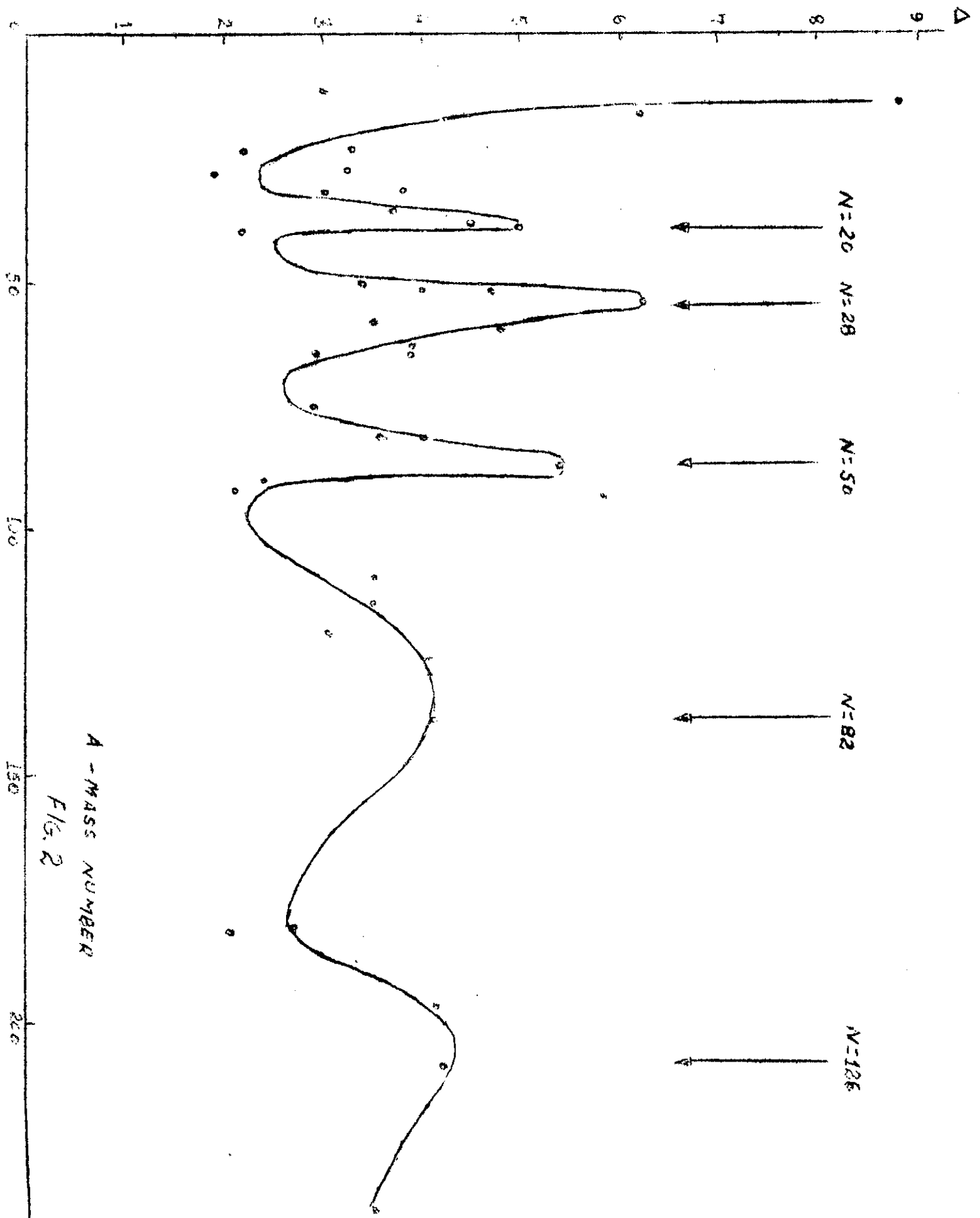


FIG 1



A - MASS NUMBER  
FIG. 2