

Angra Neutrino Project: status and plans



J.C. Anjos¹, A.F. Barbosa¹, E. Casimiro¹, R.M.O. Galvão¹, H.P. Lima Jr.¹, J. Magnin¹,
H. da Motta¹, A. Schilhtz¹, R. Shellard¹, F.R.A. Simão¹, A. Berstein²,
N. Bowden³, W. Fulgione⁴, R.Z. Funchal⁵, M.M. Guzzo⁶, E. Kemp⁶,
O.L.G. Peres⁶, D. Reyna⁷, H. Nunokawa⁸

¹Centro Brasileiro de Pesquisas Físicas – CBPF - Brazil

²Lawrence Livermore National Laboratory – USA

³Sandia National laboratories – USA

⁴Istituto Nazionale di Astrofisica – Istituto di Fisica dello Spazio Interplanetario – Italy

⁵Instituto de Física – Universidade de São Paulo – USP -Brazil

⁶Instituto de Física Gleb Wataghin, Universidade Estadual de Campinas – UNICAMP – Brazil

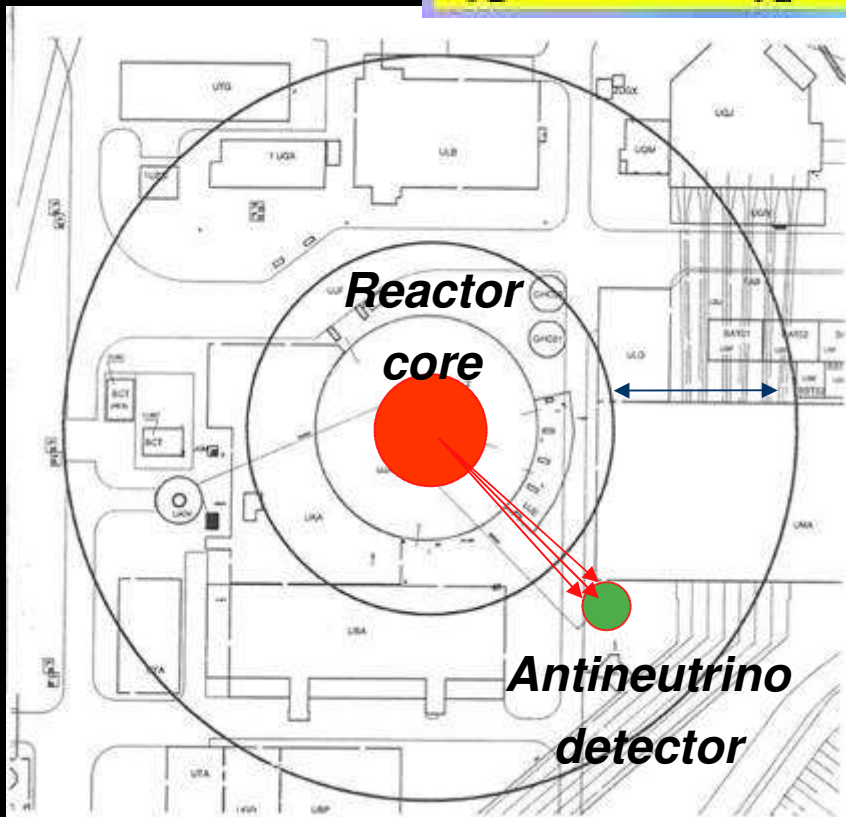
⁷Argonne National Laboratory – USA

⁸Departamento de Física, Pontifícia Universidade Católica do Rio de Janeiro – PUC-RJ - Brazil

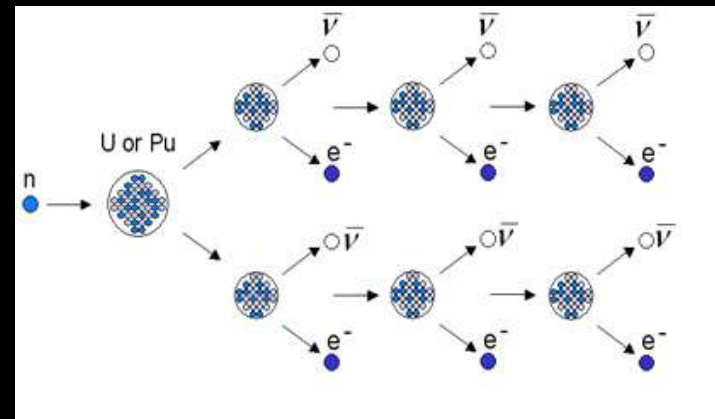
Why the interest in antineutrino detectors?

- *Search for new methods on safeguards verification*
- *Antineutrinos can not be shielded and are produced in very large amounts*
- *Antineutrinos produced in reactors can reveal fissile composition of nuclear fuel*
- *Non-intrusive: Remotely monitor real-time reactor state: thermal power and fissioning material*

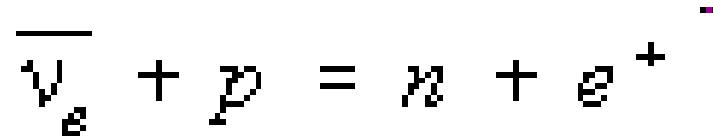
Non intrusive method to check reactor activity



Plutonium production chain

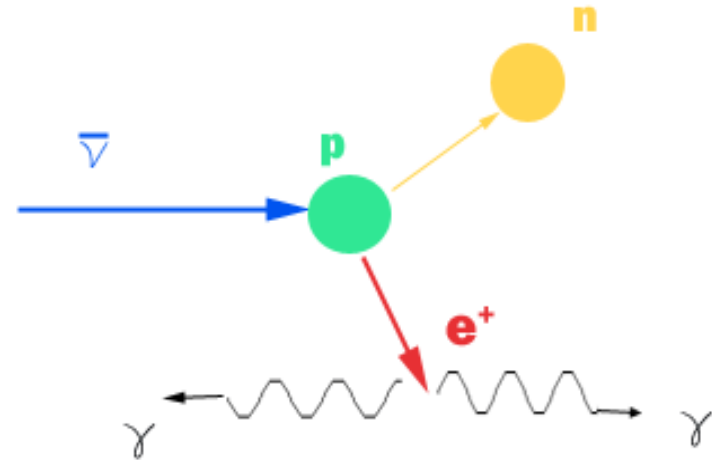


Detection of Antineutrinos



The antineutrino interacts with a proton producing...

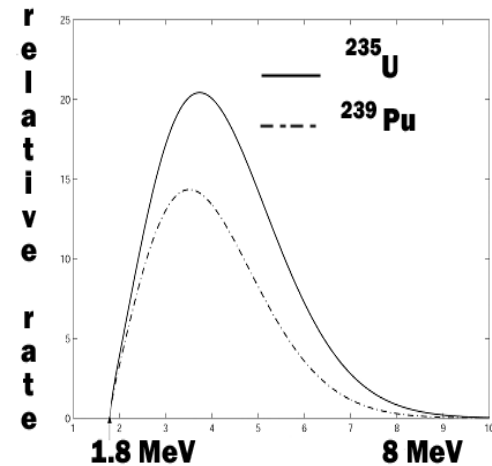
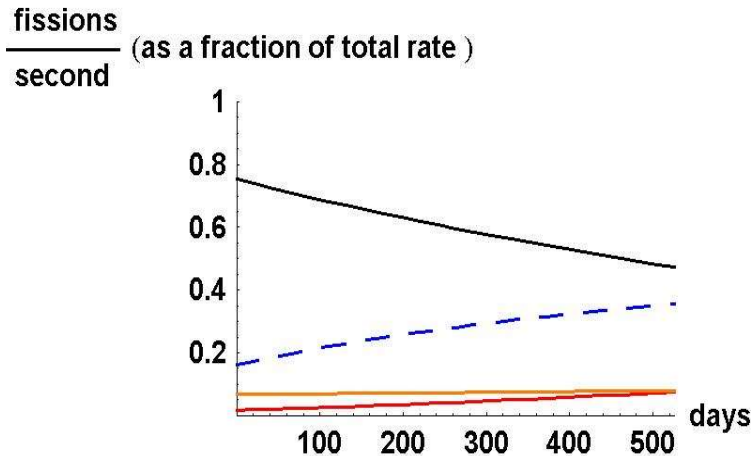
- A 1-7 MeV positron
- A few keV neutron
- mean time interval 28 μ sec



Both final state particles deposit energy in a scintillating detector over 10s or 100s of microsecond time intervals (depending on the medium)

Both energy depositions and the time interval are measured

The Basic Technical Idea



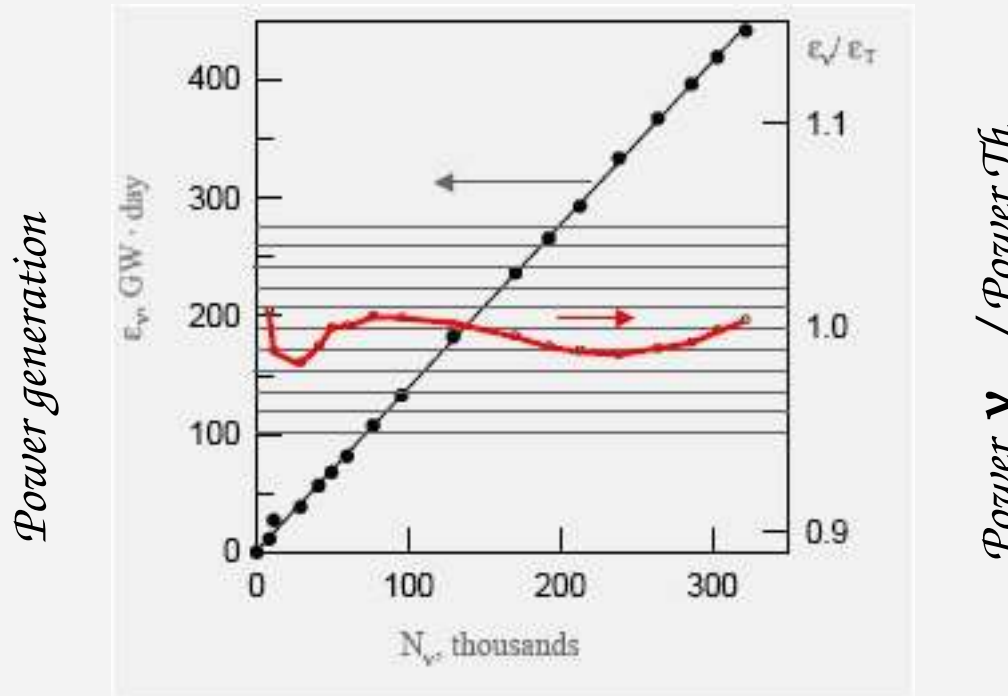
- *Relation between delivered thermal power and antineutrino flux*

$$N_{\nu} = \gamma \cdot (1 + k) \cdot P_{\text{th}}$$

*Dependence on
detector features*

*Dependence on
fuel composition*

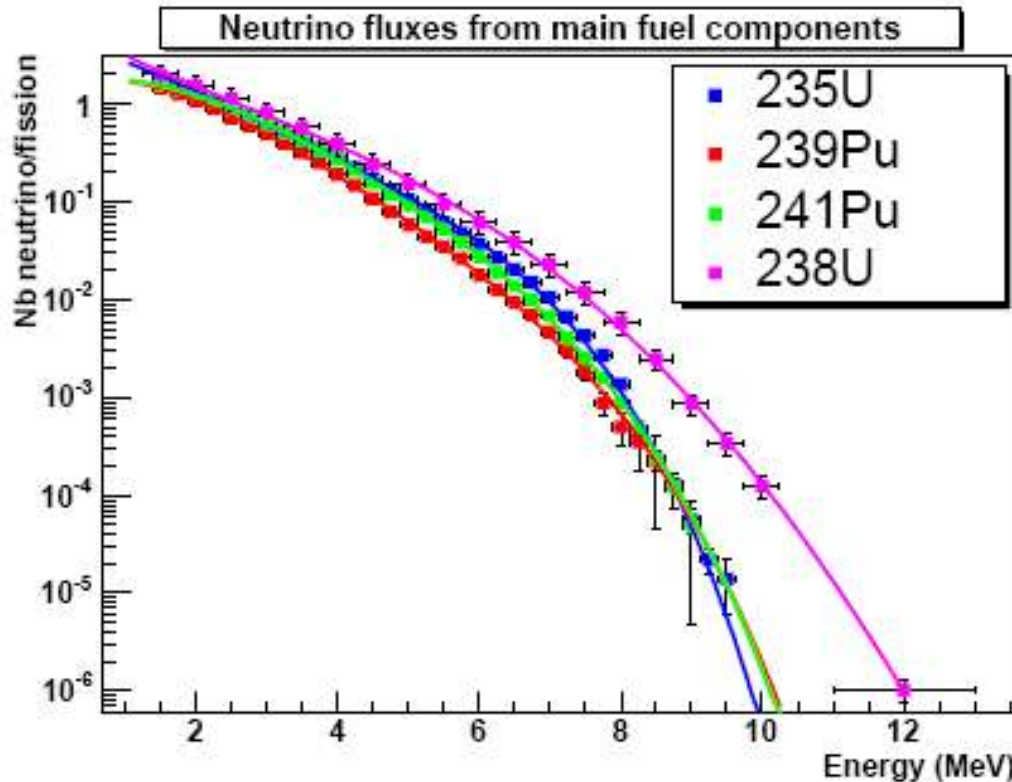
Reactor power x neutrino flux



Number of antineutrinos

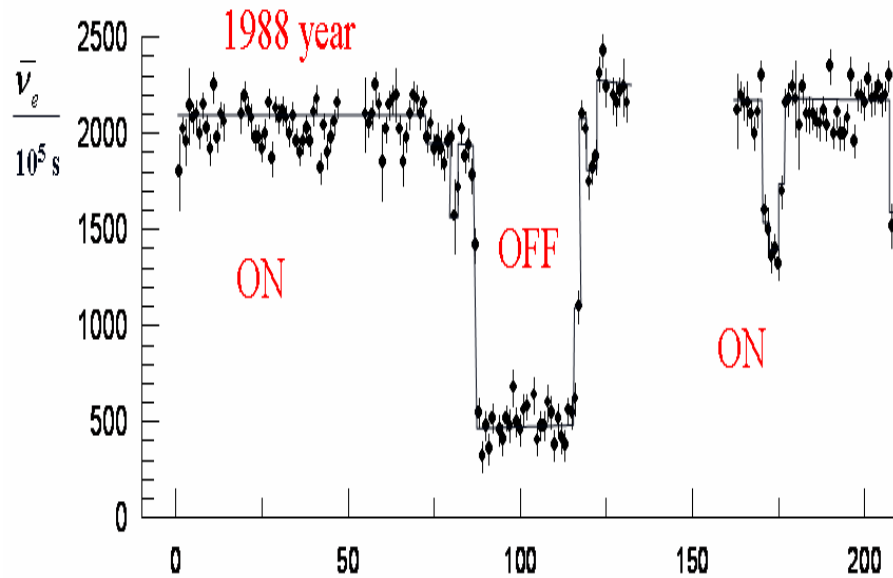
Antineutrino Emission Spectrum from Reactor Fuel

	Mean energy per fission (MeV)	Start of Cycle	End of Cycle
^{235}U	201.7 ± 0.6	60.5%	45.0%
^{238}U	205.0 ± 0.9	7.7%	8.3%
^{239}Pu	210.0 ± 0.9	27.2%	38.8%
^{241}Pu	212.4 ± 1.0	4.6%	7.9%

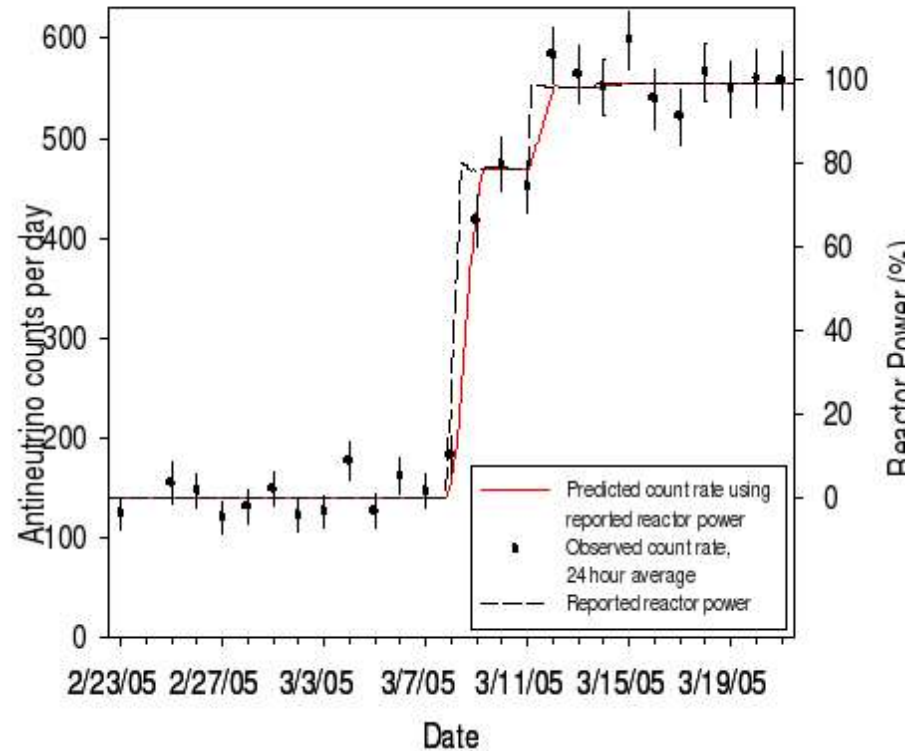


*Antineutrino spectra
measured by ILL group
1983-1989*

Checking reactor activity:



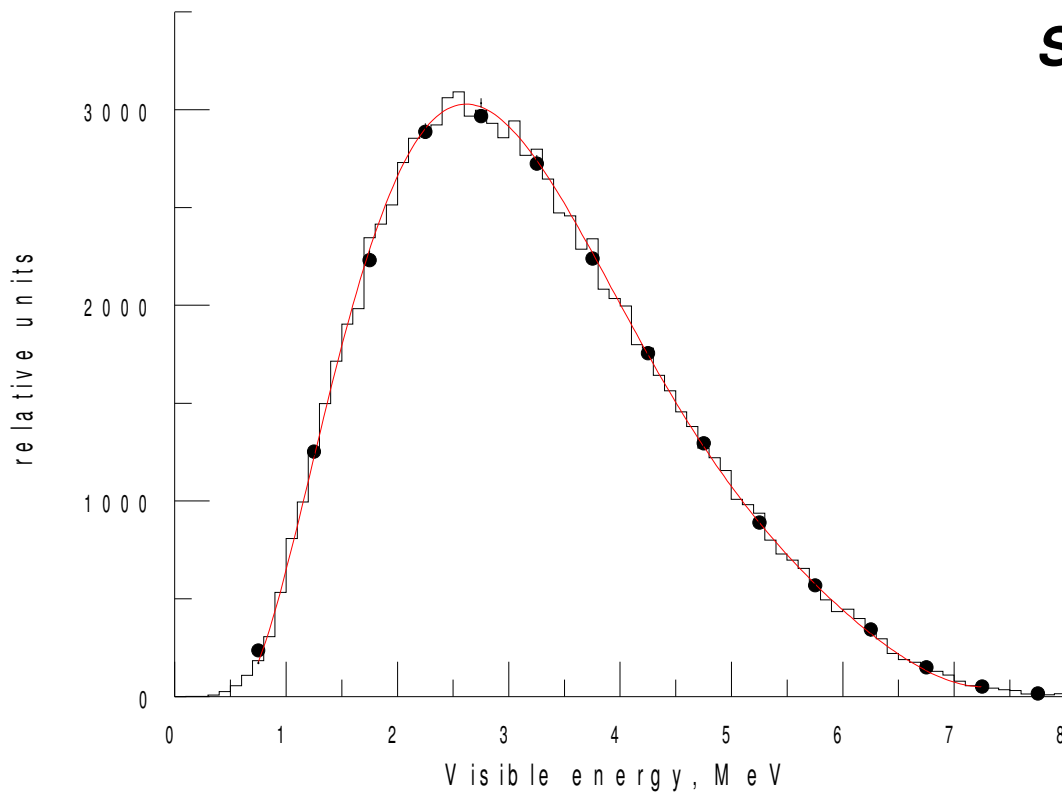
Rovno (Ukraine)



San Onofre (USA)

Fuel composition

Virtual experiment:



Simulated spectrum:

$^{235}\text{U} = 0.614,$

$^{239}\text{Pu} = 0.274,$

$^{238}\text{U} = 0.074,$

$^{241}\text{Pu} = 0.038,$

fitting spectrum:

$^{235}\text{U} = 0.631 \pm 10\%,$

$^{239}\text{Pu} = 0.260,$

$^{238}\text{U} = 0.074,$

$^{241}\text{Pu} = 0.035,$

Very Near Detector: Standard 3 volumes Design

A) Target ($R_1=0.5\text{m}$; $h_1=1.3\text{m}$)

- **Acrylic vessel + lqd scintillator(+Gd)**

B) Gamma-Catcher ($R_2=0.8\text{m}$ $h_2=1.9\text{m}$)

- **Acrylic vessel + lqd scintillator**

C) Buffer ($R_3=1.4\text{m}$; $h_3=3.10\text{m}$)

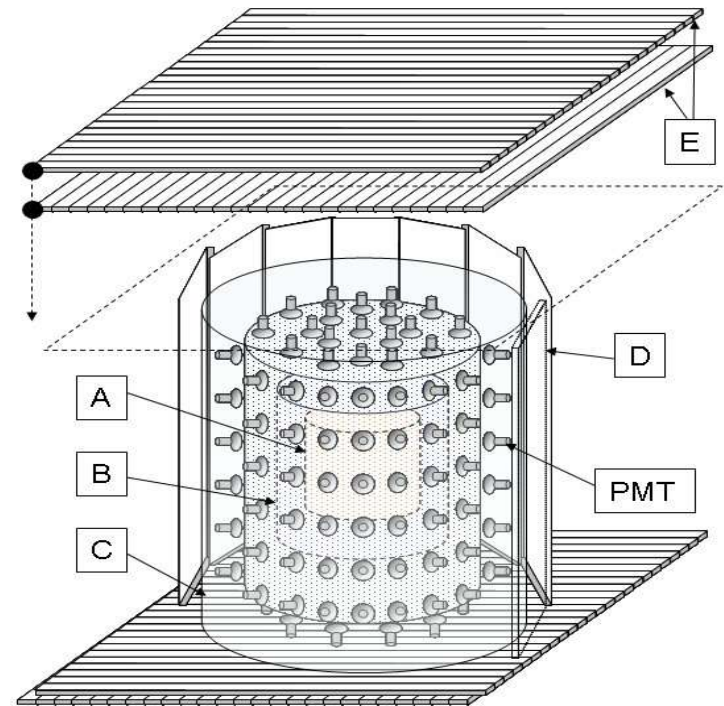
- **Steel vessel + mineral oil**

D) Vertical Tiles of Veto System

E) X-Y Horizontal Tiles of Veto System

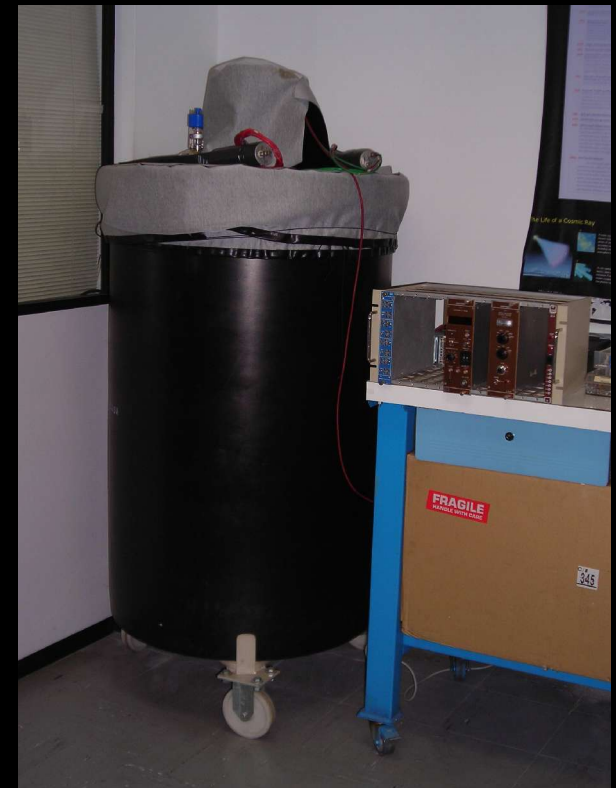
- **Plastic scintillator padles**

above and under the external steel cylinder:
muon tracking through the detector



Phase I: Setup infrastructure at the Angra site:

- **20' container near the reactor building**
- **Measurement of local muon flux:**
- **Cerenkov detector (Auger test tank)**
- **Muon telescope (4 Minos type scintillator planes)**
- **Measurement of radioactive background: (rocks and sand)**



Phase II: Deploy LVD tank

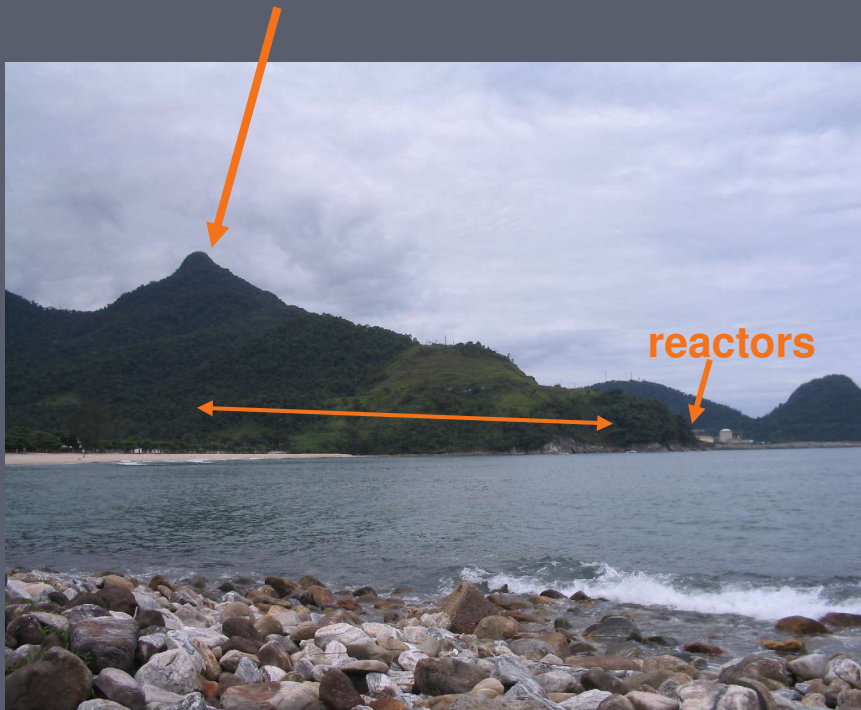
- **1 ton gadolinium doped liquid scintillator tank**
- **test signal+background**
- **Tests with Californium source**
- **Final site selection for underground laboratory**

Phase III:

- *Construction of the underground laboratory.*
- *Construction of three volume detector and muon veto.*
- *Deployment of detector parts, integration and commissioning.*

Phase IV (2013?): high precision measurement of θ_{13} ?

“Morro do Frade”



- **Near (reference) detector:**
 - 50 ton detector (7.2 m dia)
 - 300 m from core
 - 250 m.w.e.
- **Far (oscillation) detector:**
 - 500 tons (12.5 m dia)
 - 1500 m from core
 - 2000 m.w.e.
(under “Frade” peak)
- **Very Near detector:**
 - 1 ton prototype project
 - < 50m of reactor core
- **Detector Construction**
 - Standard 3 volume design

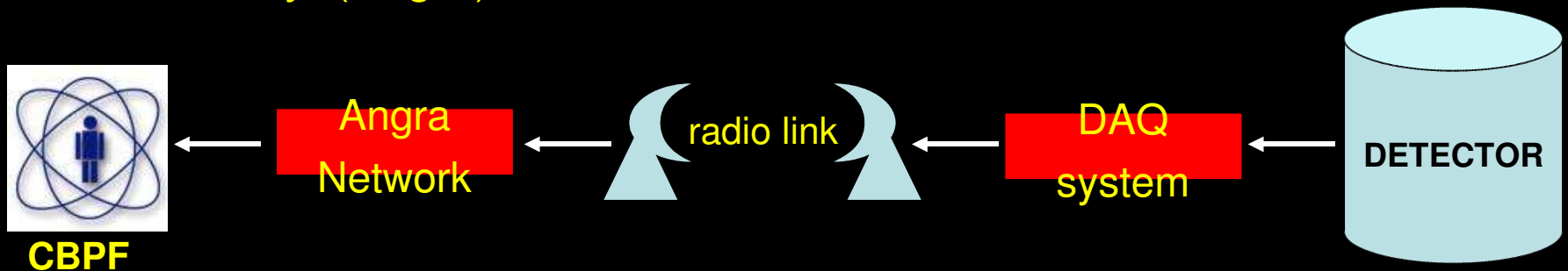
Detector Operation

Remote data acquisition

The detector continuously acquires data, recording events that may be classified as:

- Neutrino events
- VEM
- Other (ex: muons)

Data is locally (Angra) stored and sent to a central station at CBPF



Angra Project: Present Status

- **Meeting September 05, 2006 with Eletronuclear representatives to define next steps.**
- **Detailed project under way to be presented to the Minister of Science and Technology and to FAPESP.**
- **Start to test components at CBPF and UNICAMP:
(phototubes + VME electronics)**

Conclusions

- *Previous experiments demonstrate a good capability of using Antineutrinos for Nuclear reactor distant monitoring.*
- *High precision thermal power and fuel composition measurement can be achieved.*
- *Better accuracy for antineutrino spectra of U & Pu is needed.*
- *Good opportunity develop experimental neutrino physics in Brazil and to contribute to new safeguards techniques.*
- *Short baseline Neutrino Oscillations : collaboration with Double Chooz? High precision experiment around 2013?*