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FIELD OBSERVABLES IN THE RADIATION GAUGE FOR THE MAXWELL AND
SPIN 2 FIELDS

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ABSTRACT

It is shown that given the potentials for spin 1 zero rest mass fields, the process of producing field potentials which commute with the generator of gauge transformations is equivalent to modify the classical commutator (the Poisson bracket), which holds for these initial potentials, to the Dirac bracket which describes the commutation algebra for the gauge invariant field potentials in the radiation gauge. This result is extended for spin 2 zero rest mass fields without self-interaction. This last case is taken as the weak field approximation of the full non-linear gravitational field equations of general relativity.

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INTRODUCTION

For the Maxwell field is well known that a process for obtaining field potentials which are gauge invariant, in the Hamiltonian formalism, is obtained by introducing the transverse field potentials, as those which are divergence free. Since a spin 1 gauge transformation adds to the potentials the gradient of a scalar function, and thus a longitudinal vector, it is clear that any divergenceless potential will be invariant under such transformations, as long as the transverse character is conserved after the transformation. This may also be seen from the expression for the generator of gauge transformations,

$$C(x^0) = - \int \Lambda(\vec{x}, x^0) p_{r,r}(\vec{x}, x^0) d_3x$$

In this formula, Λ is the gauge function and p_r is the canonical momentum density for the spin 1 field. Latin indices indicate degrees of freedom going from 1 to 3. Greek indices go from 1 to 4. The term commutator used in this paper refer to the classical commutator, that is, the Poisson bracket. The metric tensor is the Minkowski tensor $\eta_{\mu\nu}$ with signature +2, thus all three-dimensional operations are done for the metric $\eta_{rs} = \delta_{rs}$, and no distinction is done between contravariant and covariant indices in three-dimensions.

Given arbitrary (gauge variant) potentials A_i , we have

$$\left[A_i, C \right] = \Lambda_{,i}$$

But the divergenceless field functional

$$A_i^T = A_i - \partial_i \frac{1}{\nabla^2} \partial_j A_j$$

has a null commutator with the gauge generator C . In this paper we show that the commutation algebra of the transverse potentials is identical to the algebra

of the Dirac bracket. It is known that the algebra of the Dirac bracket for the case of the Hamiltonian formulation of general relativity is realized by the usual Poisson bracket algebra of the so called "starred field functionals" ¹, which describe the field observables of the theory. Thus, we have shown that in the radiation gauge this conclusion may be extended for the spin 1 massless field in flat spacetime. Similarly, it is also proved that the same result applies to spin 2 massless fields without self interactions. This field is taken as the weak field approximation for the gravitational field equations of general relativity.

This result may also be recasted as a proof of equivalence between the A-D-M method for quantization ² and the B-K method ³, in the radiation gauge and within the approximations presently considered.

With regard to notation, we indicate the partial derivatives by any one of the symbols, $\partial_i \phi$, $\frac{\partial \phi}{\partial x^i}$ or $\phi_{,i}$, for any quantity ϕ . Covariant derivatives are not used due to our approximation of a linearized gravitational field.

1. GAUGE INVARIANT CANONICAL VARIABLES FOR THE MAXWELL FIELD IN FLAT SPACETIME.

The Hamiltonian theory for the Maxwell field contains one relation of constraint connecting the three components of the canonical momentum p_r .

$$B \equiv p_{r,r} \approx 0 \quad (1)$$

The symbol \approx means equal to zero in the weak sense in Dirac's notation ⁴. We want to prove that for this case the construction of the "starred field components" of B-K formalism is equivalent to the construction of the T-type poten

tials of the A-D-M formalism. As was stated in the introduction, the P.B. relations among the starred field components is the same as the Dirac bracket among the potentials themselves. Therefore, as long as we show that the T-type potentials have the same commutation algebra as the starred potentials, we have proved that the commutation algebra of the T-type potentials is the same as the Dirac bracket algebra among the potentials. This last algebra will be the commutation algebra for the field observables in the Hamiltonian formulation, now written entirely in terms of gauge invariant degrees of freedom.

For introducing the concept of the "starred field potentials" we have to introduce a gauge condition in terms of the dynamical components for the field. We choose the radiation gauge condition,

$$D \equiv A_{r,r} \approx 0 \quad (2)$$

The two constraints (1) and (2) form a set of two second-class constraints⁴, since the P.B. of D with B is,

$$[D, B'] = -\nabla^2 \delta(\vec{x}-\vec{x}') \quad (3)$$

Therefore, in presence of second class constraints two alternatives are possible, or we use the Dirac bracket directly instead of the P.B., or equivalently we still retain the P.B. but modify each component of the dynamical variables by adding to them a linear combination of the second-class constraints, such that it commutes with all second-class constraints. This last alternative defines the so called "starred dynamical variable". We use this process, by defining

$$A_i^*(\vec{x}, x^0) = A_i(\vec{x}, x^0) + \int \mu_i(\vec{x}, \vec{x}') P'_{r,r} d_3x' + \int \alpha_i(\vec{x}, \vec{x}') A'_{r,r} d_3x' \quad (4)$$

where A_i is the vector potential, the configuration type variable in the the

Hamiltonian formalism for electrodynamics. Similarly, in place of the canonical momentum p_i we write

$$p_i^*(x, x^0) = p_i(x, x^0) + \int \beta_i(x, x') p'_{r,r} d_3 x' + \int \gamma_i(x, x') A'_{r,r} d_3 x' \quad (5)$$

where the coefficients μ_i , α_i , β_i and γ_i are determined by the conditions

$$\left[A_i^*, B' \right] = \left[A_i^*, D' \right] = 0 \quad (6)$$

$$\left[p_i^*, B' \right] = \left[p_i^*, D' \right] = 0 \quad (7)$$

From the first two conditions (6) and (7) we see that the starred dynamical variables are gauge invariant, since the generator of gauge transformations is

$$C(x^0) = - \int \Lambda(\vec{x}, x^0) B d_3 x$$

and all starred dynamical functions commute with C , even if the gauge function Λ is also a function of the dynamical variables A_i and p_i (we have put $p_{r,r}$ equal to zero after computing all commutators). For the case of spin 1 massless fields, the imposition of gauge invariance for the canonical momentum is not really necessary since p_i is gauge invariant by definition. However, for spin 2 massless fields this imposition will be necessary, and since the method of definition of the "starred canonical variables" is general, we have maintained this condition here.

The conditions (6) and (7) imply in

$$\nabla'^2 \alpha_i(\vec{x}, \vec{x}') = \delta_{,i}(\vec{x} - \vec{x}') \quad (8)$$

$$\nabla'^2 \mu_i(\vec{x}, \vec{x}') = 0 \quad (9)$$

$$\nabla'^2 \beta_i(\vec{x}, \vec{x}') = \delta_{,i}(\vec{x} - \vec{x}') \quad (10)$$

$$\nabla'^2 \gamma_i(\vec{x}, \vec{x}') = 0 \quad (11)$$

The solution of (8) and (10) is given by,

$$\alpha_i = \beta_i = \frac{1}{4\pi} \left(\frac{1}{|\vec{x}-\vec{x}'|} \right)_{,i}$$

So that the A_i^* and p_i^* have the form

$$A_i^* = A_i + \frac{1}{4\pi} \int \left(\frac{1}{|\vec{x}-\vec{x}'|} \right)_{,i} A'_{r,r} d_3x' + \int \mu_i(\vec{x},\vec{x}') p'_{r,r} d_3x' \quad (12)$$

$$p_i^* = p_i + \frac{1}{4\pi} \int \left(\frac{1}{|\vec{x}-\vec{x}'|} \right)_{,i} p'_{r,r} d_3x' + \int \gamma_i(\vec{x},\vec{x}') A'_{r,r} d_3x' \quad (13)$$

with μ_i and γ_i solutions of the Laplace equation. These formulas may be written as

$$A_i^* = A_i - \partial_i \frac{1}{\nabla^2} \partial_r A_r + \int \mu_i(\vec{x},\vec{x}') p'_{r,r} d_3x'$$

$$p_i^* = p_i - \partial_i \frac{1}{\nabla^2} \partial_r p_r + \int \gamma_i(\vec{x},\vec{x}') A'_{r,r} d_3x'$$

Thus, up to the terms containing μ_i and γ_i , the A_i^* and p_i^* are just the transverse field variables of the A-D-M theory. It may be shown that the terms in μ_i and γ_i do not contribute to the commutation relations of the A_i^* and p_i^* (use (9) or (11)),

$$\left[A_i^*, p_k^* \right] = \delta_{ik} \delta(\vec{x}-\vec{x}') + \frac{1}{4\pi} \left(\frac{1}{|\vec{x}-\vec{x}'|} \right)_{,ik} \equiv \delta_{ik}(\vec{x}-\vec{x}') \quad (14)$$

where $\delta_{ik}(\vec{x}-\vec{x}')$ is the transverse delta function. Since all fields of interest, similarly to the case of the A-D-M theory have to be free of singularities and have to vanish at spatial infinity, we can take as the solution of the Laplace equation

$$\mu_i = \gamma_i = 0$$

In this case the A_i^* will commute with A_k^* (the same for the momentum p_i^*), and the identification of the "starred field potentials" with the T-type functionals of the A-D-M theory is completed.

$$A_i^* = A_i^T, p_i^* = p_i^T$$

From the relations,

$$\left[A_i^T, p_j^T \right] = \left[A_i^*, p_j^* \right] = \left[A_i, p_j \right]^*$$

(by $[f,g]^*$ we indicate the Dirac bracket of any given quantities f and g), we see that the commutation relations among the T-type functional are just the commutations arising from the Dirac bracket of the initial gauge variant A_i and the original momentum p_i .

2. THE CANONICAL VARIABLES h_{ij}^* AND p_{ij}^* FOR SPIN 2 IN THE LINEAR APPROXIMATION

Here we extend our previous conclusions for the weak field approximation of the general relativistic field equations of the gravitational field. The field obtained is a spin 2 massless field without self interactions. The corresponding Hamiltonian version contains four relations of constraint

$$\mathcal{H}_r \equiv 2 p_{rs,s} \approx 0 \quad (15)$$

$$\mathcal{H}_L \equiv h_{rs,rs} - h_{rr,ss} \approx 0 \quad (16)$$

which correspond to the unique constraint (1) for electrodynamics. Here we have four constraints due to the fact that a spin 2 gauge transformation involves four arbitrary gauge functions instead of just one gauge function

as was the case for the spin 1 massless field. In the Lagrangian formalism, which is a four-dimensional formalism, for the symmetric second rank Lorentz tensor $h_{\mu\nu}$, related to the metric $g_{\mu\nu}$ and to the Minkowski tensor $\eta_{\mu\nu}$ by

$$g_{\mu\nu} = \eta_{\mu\nu} + h_{\mu\nu}$$

we have the gauge transformation

$$h'_{\mu\nu}(x) = h_{\mu\nu}(x) - \Lambda_{\mu,\nu}(x) - \Lambda_{\nu,\mu}(x)$$

involving the four components of the gauge function Λ_{μ} . In the Hamiltonian theory we obtain a similar structure, but now divided into the gauge transformation of the configuration field variables, the h_{ij} , and the gauge transformation of the momentum variables p_{ij} . Before writing the expressions for the generators of these transformations, we give some formulas which will be necessary. The weak field approximation (from now on it will be called by W.F.A.) is given by taking the previous metric $g_{\mu\nu} = \eta_{\mu\nu} + h_{\mu\nu}$, where only $\eta_{\mu\nu}$ acts as the metric. The field equations are the spin 2 wave equation plus the Lorentz covariant gauge condition

$$\partial^{\nu} \gamma_{\mu\nu} = 0, \quad \partial^{\nu} = \eta^{\nu\alpha} \partial_{\alpha}, \quad \gamma_{\mu\nu} = h_{\mu\nu} - \frac{1}{2} \eta_{\mu\nu} \eta^{\alpha\beta} h_{\alpha\beta} \quad (17)$$

We call attention to the fact that in the W.F.A. the momentum p_{ik} is a first order quantity. The explicit expression for p_{ik} being obtained by linearization of the exact formula derived from the Dirac Lagrangian density for general relativity.

$$p_{ik} = - \left(h_{ao,a} - \frac{1}{2} h_{aa,o} \right) \delta_{ik} + \frac{1}{2} \left(h_{io,k} + h_{ko,i} - h_{ik,o} \right) \quad (18)$$

Under a gauge transformation on the potentials $h_{\mu\nu}$, of the form written before, the p_{ik} change according to

$$P'_{ik} = P_{ik} - \delta_{ik} \nabla^2 \Lambda^0 + \Lambda^0_{,ik} \quad (19)$$

The generator for this type of gauge transformation, in the Hamiltonian theory, is (recall that $\eta^{00} = -1$)

$$G(x^0) = - \int \Lambda^0(\vec{x}, x^0) \mathcal{H}_L(\vec{x}, x^0) d_3 x \quad (20)$$

Since,

$$[P_{ik}, G] = \Lambda^0_{,ik} - \delta_{ik} \nabla^2 \Lambda^0$$

The remaining part of the gauge transformations, that is, the part which acts on the configuration potentials h_{ik} is generated by

$$J(x^0) = \int \Lambda_S(\vec{x}, x^0) \mathcal{H}_S(\vec{x}, x^0) d_3 x \quad (21)$$

Thus, similarly as before, the constraints are basically the generators for the invariance function group of the theory.

The Lorentz covariant gauge condition (17) is separated into the conditions giving the radiation gauge for spin 2,

$$h_{00} = h_{0s} = 0 \quad (22-1)$$

$$A \equiv h_{ss} \approx 0, \quad B_r \equiv h_{rs,s} \approx 0 \quad (22-2)$$

The set of eight constraints given by (15), (16) and (22-2) is of second class since,

$$[\mathcal{H}_L, \mathcal{H}'_r] = 0, \quad [\mathcal{H}_L, A'] = 0, \quad [\mathcal{H}_L, B'_s] = 0, \quad [A, B'_s] = 0$$

$$[\mathcal{H}_r, A'] = -2 \delta_{,r}(\vec{x}-\vec{x}'), \quad [\mathcal{H}_r, B'_r] = (\delta_{rs} \nabla^2 + \partial_{rs}^2) \delta(\vec{x}-\vec{x}')$$

Thus, similarly as before, we define the quantities,

$$\begin{aligned}
h_{ij}^* &= h_{ij} + \int \mu_{ijr}(\vec{x}, \vec{x}') \mathcal{H}_r(\vec{x}, x^0) d_3 x' + \int \alpha_{ij}(\vec{x}, \vec{x}') \mathcal{H}_L(\vec{x}', x^0) d_3 x' \\
&+ \int \beta_{ij}(\vec{x}, \vec{x}') A(\vec{x}', x^0) d_3 x' + \int \gamma_{ijs}(\vec{x}, \vec{x}') B_s(\vec{x}', x^0) d_3 x' \quad (23)
\end{aligned}$$

$$\begin{aligned}
P_{ij}^* &= P_{ij} + \int \lambda_{ijr}(\vec{x}, \vec{x}') \mathcal{H}_r(\vec{x}', x^0) d_3 x' + \int \phi_{ij}(\vec{x}, \vec{x}') \mathcal{H}_L(\vec{x}', x^0) d_3 x' \\
&+ \int \psi_{ij}(\vec{x}, \vec{x}') A(\vec{x}', x^0) d_3 x' + \int \tau_{ijs}(\vec{x}, \vec{x}') B_s(\vec{x}', x^0) d_3 x' \quad (24)
\end{aligned}$$

The coefficients being determined by the conditions

$$[h_{ij}^*, \mathcal{Q}] = 0, \quad [P_{ij}^*, \mathcal{Q}] = 0$$

where by \mathcal{Q} we indicate the set of all eight constraints,

$$\mathcal{Q} = \{ A, B_s, \mathcal{H}_L, \mathcal{H}_S \}$$

We determine first the coefficients standing on Eq. (23). From the condition that h_{ij}^* commutes with \mathcal{H}_r , A and B_r we get the equations.

$$\begin{aligned}
\delta_{ir} \delta_{,j}(\vec{x}-\vec{x}') + \delta_{jr} \delta_{,i}(\vec{x}-\vec{x}') &= -2 \int \beta_{ij}(\vec{x}, \vec{x}'') \delta_{,r}(\vec{x}''-\vec{x}') d_3 x'' \\
- \int \gamma_{ijs}(\vec{x}, \vec{x}'') (\delta_{sr} \nabla''^2 + \partial_{rs}'') \delta(\vec{x}''-\vec{x}') d_3 x'' & \quad (25)
\end{aligned}$$

$$\frac{\partial}{\partial x'^r} \mu_{ijr}(\vec{x}, \vec{x}') = 0 \quad (26)$$

$$\nabla'^2 \mu_{ijr}(\vec{x}, \vec{x}') + \frac{\partial^2 \mu_{ijk}(\vec{x}, \vec{x}')}{\partial x'^k \partial x'^r} = 0 \quad (27)$$

For the commutator of h_{ij}^* with \mathcal{H}_L we get no information since h_{ij}^* automatically commutes with \mathcal{H}_L within the gauge conditions presently used.

We can write (25) in the form

$$\frac{\partial}{\partial x'^s} \left(-2 \delta_{sr} \beta_{ij} + \frac{\partial \gamma_{ijr}}{\partial x'^s} + \frac{\partial \gamma_{ijs}}{\partial x'^r} \right) = \frac{\partial}{\partial x'^s} (\delta_{ir} \delta_{sj} + \delta_{jr} \delta_{is}) \delta(\vec{x} - \vec{x}')$$

From this equation we can write

$$-2 \delta_{sr} \beta_{ij} + \gamma_{ijr,s'} + \gamma_{ijs,r'} = (\delta_{ir} \delta_{sj} + \delta_{jr} \delta_{is}) \delta(\vec{x} - \vec{x}') \quad (28)$$

For obtaining (28) we have neglected a divergenceless term $\phi_s(\vec{x}')$ (such term may be added to the left, or to the right, side of (28), but as we will see, we can obtain the desired solution without using this new term). Solving (28) for the β_{ij} ,

$$\beta_{ij} = -\frac{1}{3} (\delta_{ij} \delta(\vec{x} - \vec{x}') - \gamma_{ijr,r'}(\vec{x}, \vec{x}')) \quad (29)$$

Now, from (26) and (27) we have

$$\nabla'^2 \mu_{ijr} = 0$$

and since we look for fields which are free of singularities and which tend to zero at spatial infinity, we may take

$$\mu_{ijr} = 0 \quad (30)$$

The Eqs. (29) and (30) allow us to put the h^*_{ij} of (23) as

$$\begin{aligned} h^*_{ij} &= \tilde{h}_{ij} + \int \alpha_{ij}(\vec{x}, \vec{x}') \left(\frac{\partial B'_s}{\partial x'^s} - \nabla'^2 A' \right) d_3 x' \\ &+ \frac{1}{3} \int \frac{\partial \gamma_{ijr}(\vec{x}, \vec{x}')}{\partial x'^r} A' d_3 x' + \int \gamma_{ijs}(\vec{x}, \vec{x}') B'_s d_3 x' \end{aligned} \quad (31)$$

where \tilde{h}_{ij} is the trace free combination

$$\tilde{h}_{ij} = h_{ij} - \frac{1}{3} \delta_{ij} A$$

In order that the h_{ij}^* be of the type TT of the A-D-M theory, it is first of all necessary that the trace of h_{ij}^* vanishes. In the formula (23), or equivalently in (31), the coefficients α_{ij} and γ_{ijr} are symmetric over i, j (the other coefficients have the same symmetry). Taking trace in (31), it is simple to verify that h_{ii}^* is zero only if

$$\alpha_{ss}(\vec{x}, \vec{x}') = \gamma_{ssr}(\vec{x}, \vec{x}') = 0$$

Therefore, the two-point functions α_{ij} , γ_{ijr} have to be symmetric over i, j and for all \vec{x} and \vec{x}' have to be trace free. Besides this, they cannot depend on the dynamical variables since this would generate higher order terms which are neglected in the W.F.A. Since in (31) we have no further information on the explicit form for α_{ij} and γ_{ijr} , we can make use of this arbitrariness in the form of these functions for rewriting (31) with coefficients

$$\begin{aligned}\tilde{\alpha}_{ij} &= \alpha_{ij} - \frac{1}{3} \delta_{ij} \alpha_{ss} \\ \tilde{\gamma}_{ijr} &= \gamma_{ijr} - \frac{1}{3} \delta_{ij} \gamma_{ssr}\end{aligned}$$

in place of the α_{ij} and γ_{ijr} . These last two-point functions are symmetric over i, j (if the original two-point functions α_{ij} and γ_{ijr} are) and are trace free. This recalibration in (31) implies that h_{ij}^* is trace free, but if we compute its divergence on x^j we find,

$$\begin{aligned}h_{ij,j}^* &= h_{ij,j} - \frac{1}{3} h_{,i} + \int \tilde{\alpha}_{ij}(\vec{x}, \vec{x}') (B'_{m,m} - \nabla'^2 A') d_3 x' \\ &+ \frac{1}{3} \int \frac{\partial^2 \tilde{\gamma}_{ijr}(\vec{x}, \vec{x}')}{\partial x^j \partial x'^r} A' d_3 x' + \int \tilde{\gamma}_{ijs,j}(\vec{x}, \vec{x}') B'_s d_3 x'\end{aligned}$$

Imposing that the divergence of h_{ij}^* vanishes, we get a relation on the $\tilde{\alpha}_{ij}$ and $\tilde{\gamma}_{ijr}$. This relation, after partial integration may be presented in the form

$$\begin{aligned}
& \int \left(\frac{1}{3} \partial'_i A' - B'_i \right) \delta(\vec{x} - \vec{x}') d_3 x' = - \int \tilde{\alpha}_{ij, jm'}(\vec{x}, \vec{x}') B'_m d_3 x' + \\
& + \int \tilde{\alpha}_{ij, jm'}(\vec{x}, \vec{x}') \partial'_m A' d_3 x' - \frac{1}{3} \int \tilde{\gamma}_{ijr, j}(\vec{x}, \vec{x}') \partial'_r A' d_3 x' \\
& + \int \tilde{\gamma}_{ijs, j}(\vec{x}, \vec{x}') B'_s d_3 x'
\end{aligned}$$

We separate this equation into two relations, one containing only $\partial'_m A'$, the other involving only B'_m (this is possible since A and B_s represent independent combinations),

$$\begin{aligned}
\frac{1}{3} \int \delta_{im} \partial'_m A' \cdot \delta(\vec{x} - \vec{x}') d_3 x' &= \int \tilde{\alpha}_{ij, jm'}(\vec{x}, \vec{x}') \partial'_m A' d_3 x' \\
& - \frac{1}{3} \int \tilde{\gamma}_{ijm, j}(\vec{x}, \vec{x}') \partial'_m A' d_3 x' \\
- \int \delta_{is} B'_s \delta(\vec{x} - \vec{x}') d_3 x' &= - \int \tilde{\alpha}_{ij, js'}(\vec{x}, \vec{x}') B'_s d_3 x' \\
& + \int \tilde{\gamma}_{ijs, j}(\vec{x}, \vec{x}') B'_s d_3 x'
\end{aligned}$$

Consequently,

$$\begin{aligned}
\frac{1}{3} \delta_{im} \delta(\vec{x} - \vec{x}') &= \tilde{\alpha}_{ij, jm'}(\vec{x}, \vec{x}') - \frac{1}{3} \tilde{\gamma}_{ijm, j}(\vec{x}, \vec{x}') \\
- \delta_{is} \delta(\vec{x} - \vec{x}') &= - \tilde{\alpha}_{ij, js'}(\vec{x}, \vec{x}') + \tilde{\gamma}_{ijs, j}(\vec{x}, \vec{x}')
\end{aligned}$$

These equations are compatible for $\tilde{\alpha}_{ij} = 0$, since then we get just one independent equation.

$$\tilde{\gamma}_{ijs, j}(\vec{x}, \vec{x}') = - \delta_{is} \delta(\vec{x} - \vec{x}') \quad (32)$$

which is a condition fixing the value for $\tilde{\gamma}_{ijs}$. Its solution is,

$$\begin{aligned} \tilde{\gamma}_{ijs}(\vec{x}, \vec{x}') = & -\delta_{is} D_{,j}(\vec{x}-\vec{x}') - \delta_{js} D_{,i}(\vec{x}-\vec{x}') + \frac{2}{3} \delta_{ij} D_{,s}(\vec{x}-\vec{x}') \\ & + \frac{1}{2} \tilde{\Delta}_{ij,s}(\vec{x}-\vec{x}') \end{aligned} \quad (33)$$

where $D(\vec{x}-\vec{x}')$ is the Green function of the Poisson equation, and

$$\tilde{\Delta}_{ij}(\vec{x}-\vec{x}') = \frac{1}{\nabla^2} D_{,ij}(\vec{x}-\vec{x}') - \frac{1}{3} \delta_{ij} D(\vec{x}-\vec{x}') \quad (34)$$

The relation (33) represents a c-number two-point function symmetric over the first pair of indices and traceless over this pair of indices. Note that the divergence of $\tilde{\gamma}_{ijs}$ of (33) on x'^s gives

$$\tilde{\gamma}_{ijs,s'}(\vec{x}, \vec{x}') = \frac{3}{2} D_{,ij}(\vec{x}-\vec{x}') - \frac{1}{2} \delta_{ij} \delta(\vec{x}-\vec{x}') \quad (35)$$

which is different from its divergence on x^j which is given by (22).

With the choice $\tilde{\alpha}_{ij} = 0$ and $\tilde{\gamma}_{ijs}$ given by (33), the h_{ij}^* is identical to the h_{ij}^{TT} of the A-D-M theory for spin 2, in the so called N-decomposition for a second rank symmetric tensor⁴. From (33), (35) and $\tilde{\alpha}_{ij} = 0$, we have for h_{ij}^* ,

$$\begin{aligned} h_{ij}^* = & h_{ij} - \frac{1}{2} \delta_{ij} h_{ss} + \frac{1}{2} \partial_{ij}^2 \frac{1}{\nabla^2} h_{ss} - \partial_j \frac{1}{\nabla^2} \partial_k h_{ki} \\ & - \partial_i \frac{1}{\nabla^2} \partial_k h_{kj} + \frac{1}{2} \delta_{ij} \partial_s \frac{1}{\nabla^2} \partial_m h_{ms} + \frac{1}{2} \partial_s \frac{1}{\nabla^2} \partial_{ij}^2 \frac{1}{\nabla^2} \partial_k h_{ks} \end{aligned}$$

An inspection on this formula shows that indeed the trace and divergence of h_{ij}^* are zero. This relation coincides with the usual form for presenting a TT part of a given tensor h_{ij} in the N-decomposition.

The gauge invariance of h_{ij}^* is made clear, even before the identification with h_{ij}^{TT} , since it commutes with the gauge generators (20) and (21) even for

q-number gauge transformations (when the gauge functions depend on the dynamical variables).

The A-D-M method may be looked as a process for producing field functionals such that from given initial arbitrary canonical fields we obtain new fields which satisfy the gauge conditions $A = B_s = 0$. The method for obtaining the starred field variables is similar in this point, and this is made clear from the fact that we used only the left hand side of the radiation gauge conditions for the definition of h_{ij}^* , and did not take directly $A = B_s = 0$, rather we showed that the final h_{ij}^* may be chosen so as to satisfy these requirements.

For the momentum p_{ij}^* of (24) we cannot use the gauge conditions under the form (22-2) since this conducts to a contradiction. Indeed, taking the commutator of p_{ij}^* of (24) with \mathcal{H}_L , we get

$$[P_{ij}, \mathcal{H}'_L] = 0$$

which cannot be true. For avoiding this difficulty, we rewrite the radiation gauge conditions for spin 2 in a form slightly different, but mathematically equivalent,⁵ the quantity A being replaced by

$$Q \equiv \nabla^2 P_{ss} - P_{rs,rs} \approx 0 \quad (36)$$

but the remaining conditions $B_s = 0$ are retained. Then,

$$\begin{aligned} p_{ij}^* = & P_{ij} + \int \lambda_{ijr}(\vec{x}, \vec{x}') \mathcal{H}'_r d_3 x' + \int \phi_{ij}(\vec{x}, \vec{x}') \mathcal{H}'_L d_3 x' \\ & + \int \psi_{ij}(\vec{x}, \vec{x}') Q' d_3 x' + \int \tau_{ijs}(\vec{x}, \vec{x}') B'_s d_3 x' \end{aligned} \quad (37)$$

The imposition that p_{ij}^* commutes with all constraints leads to the equations.

$$\nabla'^2 \tau_{ijr}(\vec{x}, \vec{x}') + \partial'^2_{rs} \tau_{ijs}(\vec{x}, \vec{x}') = 0 \quad (38)$$

$$\nabla'^2 \nabla'^2 \psi_{ij}(\vec{x}, \vec{x}') + \frac{1}{2} (\delta_{ij} \nabla'^2 - \partial'^2_{ij}) \delta(\vec{x} - \vec{x}') = 0 \quad (39)$$

$$\nabla'^2 \nabla'^2 \phi_{ij}(\vec{x}, \vec{x}') = 0 \quad (40)$$

$$\nabla'^2 \lambda_{ijs}(\vec{x}, \vec{x}') + \partial'^2_{rs} \lambda_{ijr}(\vec{x}, \vec{x}') = -\frac{1}{2} (\delta_{js} \partial'_i + \delta_{is} \partial'_j) \delta(\vec{x} - \vec{x}') \quad (41)$$

From (40) and (39) we conclude that

$$\phi_{ij} = 0 \quad (42)$$

$$\psi_{ij} = \frac{1}{\nabla'^4} \frac{1}{2} (\phi_{,ij}(\vec{x} - \vec{x}') - \delta_{ij} \nabla'^2 \delta(\vec{x} - \vec{x}')) \quad (43)$$

Now we note that all available two-point functions have to depend on $\delta(\vec{x} - \vec{x}')$ or on the Green function $D(\vec{x} - \vec{x}')$, since they have to be c-numbers. Thus, any double differentiation on x'^r is equivalent to differentiate on x^r , and we may rewrite (41) as

$$\nabla^2 \lambda_{ijs} + \partial^2_{rs} \lambda_{ijr} = -\frac{1}{2} (\delta_{js} \partial_i + \delta_{is} \partial_j) \delta(\vec{x} - \vec{x}')$$

Differentiation on x^s gives

$$\nabla^2 \lambda_{ijs,s}(\vec{x}, \vec{x}') = -\frac{1}{2} \delta_{,ij}(\vec{x} - \vec{x}')$$

which has as solution

$$\lambda_{ijs,s}(\vec{x}, \vec{x}') = -\frac{1}{2} \int D(\vec{x} - \vec{x}'') \delta_{,i''j''}(\vec{x}'' - \vec{x}') d_3 x'' \quad (44)$$

Differentiation on x'^r in Eq. (38) gives,

$$\nabla'^2 \tau_{ijs,s'}(\vec{x}, \vec{x}') = 0$$

this implies that

$$\tau_{ijs,s'}(\vec{x}, \vec{x}') = 0$$

and thus, τ_{ijs} is constant. We take this constant as zero since we know that P_{ij}^* cannot depend on the h_{ij} , as it would depend if τ_{ijs} do not vanish. The Eq.(44) is integrated over the delta function giving as result,

$$\lambda_{ijs,s}(\vec{x}, \vec{x}') = -\frac{1}{2} D_{,ij}(\vec{x}-\vec{x}') \quad (45)$$

The solution of this equation is,

$$\lambda_{ijs}(\vec{x}, \vec{x}') = -\frac{1}{2} \partial_{ij}^2 \partial_s \frac{1}{\nabla^2} D(\vec{x}-\vec{x}') + \theta_{ijs}(\vec{x}-\vec{x}') \quad (46)$$

for $\theta_{ijs}(\vec{x}-\vec{x}')$ a c-number two-point function symmetric over i,j and divergenceless over the last index.

$$\theta_{ijs,s}(\vec{x}-\vec{x}') = 0 \quad (47)$$

The value for θ_{ijs} is obtained by imposing consistency of the solution (46) with the original equation (41). With this end we compute the Laplacian of (46)

$$\nabla^2 \lambda_{ijs} = -\frac{1}{2} \partial_{ij}^2 \partial_s D(\vec{x}-\vec{x}') + \nabla^2 \theta_{ijs}(\vec{x}-\vec{x}')$$

and, also from (46),

$$\partial_{rs}^2 \lambda_{ijr}(\vec{x}, \vec{x}') = -\frac{1}{2} \partial_{ij}^2 \partial_r \partial_{rs}^2 \frac{1}{\nabla^2} D(\vec{x}-\vec{x}')$$

(condition (47) was used). Therefore,

$$\partial_{rs}^2 \lambda_{ijr}(\vec{x}, \vec{x}') = -\frac{1}{2} \partial_{ij}^2 \partial_s D(\vec{x}-\vec{x}')$$

and the left hand side of the original Eq. (41) is

$$\nabla^2 \lambda_{ijs} + \partial_{rs}^2 \lambda_{ijr} = -\partial_{ij}^2 \partial_s D(\vec{x}-\vec{x}') + \nabla^2 \theta_{ijs}(\vec{x}-\vec{x}')$$

By consistency we should have

$$-\partial_{ij}^2 \partial_s D(\vec{x}-\vec{x}') + \nabla^2 \theta_{ijs}(\vec{x}-\vec{x}') = -\frac{1}{2} (\delta_{js} \delta_{,i}(\vec{x}-\vec{x}') + \delta_{is} \delta_{,j}(\vec{x}-\vec{x}'))$$

This is a differential equation in Θ_{ijs} . Its solution being,

$$\Theta_{ijs} = -\frac{1}{2} \left(\delta_{js} \frac{1}{\nabla^2} \delta_{,i}(\vec{x}-\vec{x}') + \delta_{is} \frac{1}{\nabla^2} \delta_{,j}(\vec{x}-\vec{x}') \right) + \frac{1}{\nabla^2} \partial_{ij}^2 \partial_s D(\vec{x}-\vec{x}')$$

which may be written as,

$$\Theta_{ijs} = -\frac{1}{2} \left(\delta_{js} D_{,i}(\vec{x}-\vec{x}') + \delta_{is} D_{,j}(\vec{x}-\vec{x}') \right) + \frac{1}{\nabla^2} \partial_{ij}^2 \partial_s D(\vec{x}-\vec{x}') \quad (48)$$

The Θ_{ijs} of (48) satisfies the condition (47). Therefore, from (46) we have

$$\begin{aligned} \lambda_{ijs} = & -\frac{1}{2} \partial_{ij}^2 \partial_s \frac{1}{\nabla^2} D(\vec{x}-\vec{x}') + \frac{1}{\nabla^2} \partial_{ij}^2 \partial_s D(\vec{x}-\vec{x}') \\ & - \frac{1}{2} (\delta_{js} D_{,i}(\vec{x}-\vec{x}') + \delta_{is} D_{,j}(\vec{x}-\vec{x}')) \end{aligned} \quad (49)$$

We note that by an argument similar to that used for the λ -equation, we can write (43) in the form

$$\Psi_{ij} = \frac{1}{\nabla^4} \frac{1}{2} (\delta_{,ij}(\vec{x}-\vec{x}') - \delta_{ij} \nabla^2 \delta(\vec{x}-\vec{x}')) \quad (50)$$

This formula may be simplified to

$$\Psi_{ij} = \frac{1}{2} \frac{1}{\nabla^2} D_{,ij}(\vec{x}-\vec{x}') - \frac{1}{2} \delta_{ij} D(\vec{x}-\vec{x}') \quad (51)$$

Using the value for the several coefficients, we can finally write down the formula for the P_{ij}^* ,

$$\begin{aligned} P_{ij}^* = & P_{ij} + \int \left[\frac{1}{\nabla^2} \partial_{ij}^2 D_{,s}(\vec{x}-\vec{x}') - \frac{1}{2} \partial_{ij}^2 \partial_s \frac{1}{\nabla^2} D(\vec{x}-\vec{x}') - \right. \\ & \left. - \frac{1}{2} (\delta_{js} D_{,i}(\vec{x}-\vec{x}') + \delta_{is} D_{,j}(\vec{x}-\vec{x}')) \right] \mathcal{H}_s(x') d_3 x' + \\ & + \frac{1}{2} \int \left[\frac{1}{\nabla^2} D_{,ij}(\vec{x}-\vec{x}') - \delta_{ij} D(\vec{x}-\vec{x}') \right] Q(x') d_3 x' \end{aligned} \quad (52)$$

which is a functional of the original momentum p_{ij} . As for the configuration variables, we also have here that $P_{ij}^* = P_{ij}^{TT}$.

3. CONCLUSION

In the usual formalism, involving the gauge variant canonical variables h_{ij} and p_{ij} for the linearized gravitational spin 2 field, the gauge functions cannot be arbitrarily chosen in the radiation gauge. Under a gauge transformation the h_{ij} change as

$$h'_{ij} = h_{ij} - \Lambda_{i,j} - \Lambda_{j,i}$$

Then, the gauge conditions

$$A = h_{ss} = 0$$

$$B_s = h_{sr,r} = 0$$

are valid in a new gauge frame only if

$$\Lambda_{s,s} = 0, \quad \nabla^2 \Lambda_s = 0$$

These last relations are equivalent to impose that the left hand side of the above gauge conditions commute with the generators of the gauge transformations.

$$[A, J] = 0, \quad [B_s, J] = 0$$

(if we use in place of the gauge condition $A = h_{ss} = 0$, the condition

$Q = \nabla^2 p_{ss} - p_{rs,rs} = 0$ a similar situation holds). However, in the case where we work with the functionals h_{ij}^* and p_{ij}^* no condition need to be imposed on the gauge functions Λ_0 and Λ_s , since under gauge transformations

$$h_{ij}^{*' } = h_{ij}^*$$

$$p_{ij}^{*' } = p_{ij}^*$$

and thus h_{ij}^* and p_{ij}^* automatically commute with the generators of the gauge transformations.

$$[A^*, J] = 0, [B_s^*, J] = 0$$

and also trivially we have,

$$[A^*, G] = 0, [B_s^*, G] = 0 .$$

Since $[A^*, J] = [A, J]^*$, with the same property for the other commutators, we have that in the Dirac bracket algebra all eight constraints $\mathcal{H}_r, \mathcal{H}_L, A$ and B_s become of the first class. This indeed was the basic idea underlining this new commutation algebra. What was proven is that this new commutation algebra is just the commutation algebra of the transverse-transverse field functionals of the type used in the so called N-decomposition of the A-D-M formalism.

Since in the A-D-M theory a process is suggested for generalizing this for the full non-linear gravitational field equations of general relativity, in the so called C-decomposition, we may hope that similarly it may also apply for the starred field functionals underlined in the Dirac commutation algebra. Since no closed and simple form is known for g_{ij}^* and p^{*ij} in general relativity, it may happen that this analogy turns out useful in this case.

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