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D-DIMENSIONAL IDEAL GAS IN PARASTATISTICS: THERMODYNAMIC PROPERTIES

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#### ABSTRACT

We consider a parastatistics ideal gas with energy spectrum  $\varepsilon \propto |\vec{k}|^{\alpha}$  ( $\alpha > 0$ ) or even more general in a d-dimensional box with volume V (periodic boundary conditions), the number N of the gas particles being well determined (real particles) or not (quasi particles). We calculate the main thermodynamic quantities (chemical potential, internal energy, specific heat C, equation of states, latent heat, average numbers of particles) for arbitrary d, a, T (temperature) and p (maximal number of particles per state allowed in the parastatistics): The main asymptotic regimes are worked out explicitely. In particular, the Bose-Einstein condensation for fixed idensity. N/V appears as a non uniform convergence in the p \* 11mit, in complete analogy with the standard critical phenomena which appear in interac ting systems in the N →∞ limit. The system behaves ressentially like a Fermi-Dirac one for all finite values of p, and reveals a Bose-Einstein behavior only in the  $p \rightarrow \infty$  limit. For a instance, at low temperatures  $C \notin T$  if  $p < \infty$  and  $C \notin T^{d/\alpha}$  if  $p \to \infty$ . the Sommerfeld integral and its expansion are generalized to an arbitrary finite p.

Key-words: Parastatistics; Ideal gas; Bose-Einstein condensation; Sommerfeld integral.

### 1 INTRODUCTION

Since the pioneer work by Gentile [1] many different have been done which interpolate between Bose-Einstein statistics (see for instance Ref. [2]). Possible applica-Fermi-Dirac tions have been searched in Field Theory and Elementary Particles Physics [3-6] as well as in Condensed Matter Physics (Molecular excitors and magnons [7], quantum Hall effect [3]). The ideal gas in parastatistics naturally constitutes a privileged creference system which can be used as unperturbed starting to study more complex systems. Also it will be shown that many exact (or asymptotically exact) analytical expressions can be obtained which could serve as testing material for various approximation methods. The purpose of the present paper is establishment of the main thermodynamic properties gas at fixed volume in d-dimensions (d>0), and for a quite general energy spectrum  $\varepsilon = \left(\sum_{j=1}^{d} a_j k_j^2\right)^{\alpha/2}$  ( $a_j > 0$  and  $\alpha > 0$ ;  $\alpha = 1$  corresponds to photons and to short wave-vector acoustic phonons, among others; a = 2 corresponds to non relativistic free particles to short wave-vector acoustic magnons in Heisenberg ferromagnets, among others).

In Section 2 we introduce the specific ideal gas model and obtain the associated density of states; in Sections 3 and 4 we calculate the main thermodynamic quantities respectively corresponding to fixed (real particles) and unfixed (quasi-particles) total number of gas particles; finally we conclude in Section 5. In Appendix we generalize to parastatistics the sommerfeld integral and its standard expansion.

# PARASTATISTICS IDEAL GAS - DENSITY OF STATES

We consider an ideal gas of N particles (or quasi-particles) in a d-dimensional box. Each particle behaves as a planar wave with energy spectrum given by

$$\varepsilon = \left(\sum_{j=1}^{d} a_{j} k_{j}^{2}\right)^{\alpha/2} \tag{1}$$

where  $a_j > 0$   $\forall_j$ ,  $\alpha > 0$  and  $k_j$  is the j-th component of the wave vector  $\vec{k}$ . The particular case  $a_j = a$ ,  $\forall_j$ , yields  $\epsilon \propto k^{\alpha} (k \equiv |\vec{k}|)$ . Periodic boundary conditions are considered on the box which is assumed to be an orthogonal hyperparallelepiped with side lengths  $\{L_j\}$  and volume  $V \equiv \int_{j=1}^{d} L_j$ . The possible wave vectors are given by

$$k_{j} = \frac{2\pi}{L_{j}} n_{j}$$
  $(n_{j} = 0, \pm 1, \pm 2, ..., y_{j})$  (2)

which inserted into Eq. (1) yield

$$\frac{d}{j=1} \frac{(2\pi)^2 a_j}{\epsilon^{2/\alpha} L_j^2} n_j^2 = 1$$
(3)

This is the equation of an hyperellipsoid whose volume provides the number of states  $\phi(\epsilon)$  whose energy is equal to or lower than  $\epsilon$ . Consequently

$$\psi(\varepsilon) = \frac{\alpha}{d} \left(\frac{\varepsilon}{\varepsilon_0}\right)^{d/\alpha} \tag{4}$$

$$\varepsilon_{0} = \left[\frac{2^{d} \pi^{d/2} \Gamma(\frac{d}{2} + 1) \left(\frac{d}{2} + 1\right)^{1/2}}{Vd/\alpha}\right]^{\alpha/d}$$
(4')

The density of states  $\rho(\epsilon)$  is therefore given by

$$\rho(\varepsilon) = \frac{d\phi(\varepsilon)}{d\varepsilon} = \frac{1}{\varepsilon_0} \left(\frac{\varepsilon}{\varepsilon_0}\right)^{d/\alpha - 1}$$
 (5)

Notice that we treat  $\phi(\varepsilon)$  as if it was a "soft" function of  $\varepsilon$ : this is correct in the thermodynamic limit we are interested in  $(N \to \infty, V \to \infty, N/V \to constant, \varepsilon_0 \to 0)$ . The parastatistics thermal equilibrium average number  $f(\varepsilon)$  of particles per state is given by  $\square$ 

$$f(\varepsilon) = \frac{1}{e^{\beta(\varepsilon-\mu)} - \frac{p+1}{e^{(p+1)\beta(\varepsilon-\mu)} - 1}}$$
(6)

or alternatively by

$$f(\varepsilon) = \frac{\sum_{j=1}^{p} j^{\beta}(e^{-j}) \beta(\varepsilon - \mu)}{\sum_{j=0}^{p} e^{j\beta(\varepsilon - \mu)}}$$
(6')

where  $\beta \equiv 1/k_B T$  is the inverse temperature,  $\mu$  is the chemical potential, and p is the maximal number of particles allowed per state (p = 1 reproduces the Fermi-Dirac statistics and  $p \rightarrow \infty$  the Bose-Einstein one); see Fig. 1.

## 3 FIXED TOTAL NUMBER OF PARTICLES (REAL PARTICLES)

# 3.1 - Chemical potential and ground state population

We consider here the total number N of gas particles as well determined. If we respectively note  $N_{_{\rm O}}$  and  $N_{_{\rm E}}$  the populations of the ground state and the excited states, then

$$N = N_O(T) + N_e(T)$$
 (7)

We shall now calculate the chemical potential  $\mu(\textbf{T},p)$ , which is completely determined by

$$N = \int_{0}^{\infty} d\varepsilon \, \rho(\varepsilon) \, f(\varepsilon)$$
 (8)

Let us first consider the  $T \to 0$  limit, in which case  $f(\epsilon)$  is the step function indicated in Fig. 1. We then have  $N_0 = N$  and

$$N = \int_{0}^{\mu_{0}} d\varepsilon \, \rho(\varepsilon) \, p = \int_{0}^{\mu_{0}} \frac{p}{\varepsilon_{0}^{d/\alpha}} \frac{e^{d/\alpha - 1}}{d\varepsilon} \, d\varepsilon$$

$$= \frac{p}{\epsilon_0^{d/\alpha}} \frac{\alpha}{d} \mu_0^{d/\alpha}$$
 (9)

where  $\mu_0(p) \equiv \mu(0,p)$  and we have used Eq. (5) (with  $\rho(\epsilon) \equiv 0$  for  $\epsilon < 0$ ). Consequently we have that

$$\mu_{0}(p) = \frac{\mu_{0}(1)}{p^{\alpha/d}}$$
 (10)

where

$$\mu_{0}(1) = \left(\frac{Nd}{\alpha}\right)^{\alpha/d} \epsilon_{0} \tag{11}$$

or even

$$\mu_{0}(1) = \left[2^{d} \pi^{d/2} r \left(\frac{d}{2} + 1\right) \left(\frac{d}{j - 1} a_{j}\right)^{1/2} N/V\right]^{\alpha/d}$$
(12)

where we used Eq.(4\*) to make explicit the dependence of  $\mu_0$  on the concentration N/V. It is convenient to introduce the following reduced quantities:

$$t = k_B T / \mu_O(1), \qquad (13)$$

$$\overline{\mu} \equiv \mu/\mu_{0}(1), \qquad (14)$$

and

$$x \equiv \varepsilon/\mu_0(I) \tag{15}$$

Equation (8) can therefore be rewritten as follows:

$$\frac{\alpha}{d} = \int_{0}^{\infty} dx \, x^{d/\alpha - 1} \left[ \frac{1}{e^{(x - \bar{\mu})/t} - 1} - \frac{p + 1}{e^{(p+1)(x - \bar{\mu})/t} - 1} \right]$$
(16)

which makes obvious that, in energy units of  $\mu_0$  (1), the chemical potential (and in fact all the thermodynamic quantities—we shall be interested in depends on d and  $\alpha$  only through the ratio  $d/\alpha$  (see also Ref. [9]). Through a further transformation  $y \equiv x/t$  we obtain

$$\frac{\alpha}{d} t^{-d/\alpha} = \int_{0}^{\infty} dy y^{d/\alpha - 1} \left[ \frac{1}{e^{iy = i\mu/t} - 1} - \frac{p+1}{e^{(p+1)(y - i\mu/t)} - 1} \right]$$
 (17)

which provides t as an explicit function of  $\bar{\mu}/t$ , p and  $d/\alpha$ . The results are indicated in Fig. 2. As can be observed therein, the thermal dependence of  $\bar{\mu}$ , for fixed  $d/\alpha$ , flattens in the low temperature region for increasing p, and eventually exhibits, in the  $p+\infty$  limit and  $d/\alpha>1$ , a "plateau" at  $\bar{\mu}=0$  (the width of the plateau increases with increasing and not too high  $d/\alpha-1$ ). This is of course the Bose-Einstein condensation, and we see in the present framework that it appears in the  $p+\infty$  limit through a non uniform convergence completely similar to those which are observed, in the  $N+\infty$  limit, for interacting systems presenting phase transitions.

No Bose-Einstein condensation at finite temperatures exists for  $d/\alpha \le 1$ . In particulier for  $d/\alpha = 1$ , this is easy to verify as the integral of Eq. (17) is straightforwardly solved, and we obtain the following dependence of t on  $\bar{\mu}$  /t:

$$t = \frac{1}{\ln \frac{1 - e^{(p+1)\bar{\mu}/t}}{1 - e^{\bar{\mu}/t}}} = \frac{1}{\ln \sum_{j=0}^{p} e^{j\bar{\mu}/t}}$$
(18)

This expression leads, in the p  $+\infty$  limit (and taking into account that  $\overline{\mu}$  < 0 for t  $\times$ 0), to the standard d =  $\alpha$  = 2 result, namely

$$\tilde{\mu} = t \ln (1 - e^{-1/t})$$
 (19)

where we verify that  $\overline{\mu}$  vanishes at no other place than at t = 0. Let us now go back to the general expression indicated in

Eq. (17), and denote by t\* the (finite) temperature at which  $\bar{\mu}$  vanishes for fixed p and d/ $\alpha$ . It follows that

$$\frac{\alpha}{d}(t^*)^{-d/\alpha} = \int_0^\infty dy \, y^{d/\alpha - 1} \left[ \frac{1}{e^{y-1}} - \frac{p+1}{e^{(p+1)y} - 1} \right]$$
 (20)

The integral can be expressed (see page 325 of [10]) in terms of the Riemann zeta and gamma funtions  $\zeta$  and  $\Gamma$ , and we obtain

$$t* = \left\{ \left[ 1 - \frac{1}{(p+1)^{d/\alpha-1}} \right] \zeta(d/\alpha) \Gamma(d/\alpha+1) \right\}^{-\alpha/d} (d/\alpha > 1) \qquad (21)$$

These results are illustrated in Fig. 3; note that  $t_c(d/\alpha) = \lim_{p \to \infty} t^*(p,d/\alpha)$  precisely is the Bose-Einstein condensation critical (reduced) temperature (the phase transition only exists for  $p + \infty$ ), and it is given by

$$t_{c} = \left[\frac{d}{\alpha} \int_{0}^{\infty} dy \frac{y^{d/\alpha - 1}}{e^{y} - 1}\right]^{-\alpha/d}$$
 (22)

hence

$$= \left[\zeta \left(\frac{d}{\alpha}\right)\Gamma \left(\frac{d}{\alpha}+1\right)\right]^{-\alpha/d} \tag{22'}$$

$$\begin{array}{c}
\begin{pmatrix} (\bar{d}/\alpha - 1) \text{ if } d/\alpha - 1 \rightarrow +0 \\
5 \frac{\alpha}{d} \text{ if } d/\alpha \rightarrow \infty
\end{pmatrix} (22")$$

A few more steps will lead us to the thermal dependence of the (reduced) population  $N_0/N$  of the ground state below the critical temperature (above the critical temperature,  $N_0/N$  vanishes in

the thermodynamic limit  $N \rightarrow \infty$ ). We have that

$$\frac{N_{o}}{N} = 1 - \frac{N_{e}}{N} = 1 - \frac{1}{N} \int_{0}^{\infty} d\epsilon \rho(\epsilon) f(\epsilon; \mu = 0)$$

$$= 1 - \frac{d}{\alpha} t^{d/\alpha} \int_{0}^{\infty} dy \frac{y^{d/\alpha - 1}}{e^{y} = 1}$$

hence

$$\frac{N_{O}}{N} = 1 - \left(\frac{t}{t_{c}}\right)^{d/\alpha} \qquad (t \le t_{c})$$
 (23)

where we have used Eq. (22).

We now dedicate the rest of this subsection to determine the low- and high-temperature expansions of the chemical potential.

For the low-temperature asymptotic behavior we can use, into Eq. (17), expansion (A.15) of the Appendix (with A=1,  $\gamma = d/\alpha - 1$  and  $z = \bar{\mu}/t$ ) thus obtaining

$$\overline{\mu} = \frac{1}{p^{\alpha/d} \left\{ 1 + \sum_{m=1}^{\infty} \left\{ \frac{2d}{p\alpha} \zeta(2m) \left[ 1 - \frac{1}{(p+1)^{2m-1}} \right] \int_{j=0}^{2(m-1)} \left( \frac{d}{\alpha} - 1 - j \right) \right\} \left( \frac{t}{\mu} \right)^{2m} \right\}^{\alpha/d}}$$
(24)

The solution of this equation clearly has the form

$$\overline{\mu} = \frac{\mathbf{L}}{\mathbf{p}^{\alpha/d}} \left\{ \mathbf{1} + \sum_{r=1}^{\infty} \alpha_r \mathbf{t}^{2r} \right\}$$
 (25)

The substitution of this expression into both sides of Eq. (24) enables, by simple identification, the knowledge of  $\{\alpha_r\}$  up to any desired order. To the lowest correction we obtain

$$\frac{1}{\mu} \sim \frac{1}{p^{\alpha/d}} \left\{ 1 - \frac{\pi^3}{3} \frac{p^{2\alpha/d}}{p+1} \left( \frac{d}{\alpha} - 1 \right) t^2 \right\} (t + 0, \forall d/\alpha) \quad (26)$$

Note the changement of curvature which occurs at  $d/\alpha = 1$ . As a matter of fact, for  $d/\alpha = 1$ , the departure of  $\bar{\mu}$  from 1/p is slower than any power of t (and not only  $t^2$ ). For this case we obtain, from Eq. (18),

$$\frac{1}{\mu} \sim \begin{cases} \frac{1}{p} (1 - t e^{-1/pt}) & (t + 0, d/\alpha = 1, p < \infty) \\ -te^{-1/t} & (t + 0, d/\alpha = 1, p + \infty) \end{cases}$$
 (27)

For the high-temperature asymptotic behavior we can present Eq. (17) as follows:

$$\frac{\alpha}{d} t^{-d/\alpha} = \int_{0}^{\infty} dy \frac{\frac{d}{\alpha} - 1}{e^{y-z}} \frac{1}{1 - e^{z-y}} - (p+1) \int_{0}^{\infty} dy \frac{\frac{d}{\alpha} - 1}{e^{(p+1)(y-z)}} \frac{1}{1 - e^{(p+1)(z-y)}}$$

$$= e^{z} \int_{0}^{\infty} dy y^{\frac{d}{\alpha} - 1} e^{-y} (1 + e^{z-y} + \dots)$$

$$= (p+1) e^{(p+1)z} \int_{0}^{\infty} dy y^{\frac{d}{\alpha} - 1} e^{-y} + e^{2z} \int_{0}^{\infty} dy y^{\frac{d}{\alpha} - 1} e^{-2y} + \dots$$

$$= e^{z} \int_{0}^{\infty} dy y^{\frac{d}{\alpha} - 1} e^{-y} + e^{2z} \int_{0}^{\infty} dy y^{\frac{d}{\alpha} - 1} e^{-2y} + \dots$$

$$= (p+1) e^{(p+1)z} \int_{0}^{\infty} dy y^{\frac{d}{\alpha} - 1} e^{-(p+1)y} - (p+1) e^{2(p+1)z} \int_{0}^{\infty} dy y^{\frac{d}{\alpha} - 1} e^{-2(p+1)y} - \dots$$

$$= e^{z} \Gamma(\frac{d}{\alpha}) + \frac{e^{2z} \Gamma(\frac{d}{\alpha})}{2^{d/\alpha}} + \dots - \frac{e^{(p+1)z} \Gamma(\frac{d}{\alpha})}{(p+1)^{d/\alpha - 1}}$$

$$- \frac{e^{2(p+1)z} \Gamma(\frac{d}{\alpha})}{2^{d/\alpha}(p+1)^{d/\alpha - 1}} - \dots$$

hence

$$t^{-d/\alpha} \sim e^{z} \Gamma\left(\frac{d}{\alpha} + 1\right) \left\{1 + \frac{e^{z}}{2^{d/\alpha}} - \frac{e^{pz}}{(p+1)^{d/\alpha - 1}} - \frac{e^{(2p+1)z}}{2^{d/\alpha}(p+1)^{d/\alpha - 1}}\right\}$$

hence

$$z \sim -\frac{d}{\alpha} \ln t - \ln \Gamma \left( \frac{d}{\alpha} + 1 \right) - \frac{e^{z}}{2^{d/\alpha}} + \frac{e^{pz}}{(p+1)^{d/\alpha - 1}}$$

$$\sim -\frac{d}{\alpha} \ln t - \ln \Gamma \left( \frac{d}{\alpha} + 1 \right) - \frac{t^{-d/\alpha}}{2^{d/\alpha} \Gamma \left( \frac{d}{\alpha} + 1 \right)} + \frac{t^{-pd/\alpha}}{(p+1)^{\frac{d}{\alpha} - 1} \left[ \Gamma \left( \frac{d}{\alpha} + 1 \right) \right]^{p}}$$

Finally, by using  $z = \overline{\mu}/t$ , we obtain

$$\overline{\mu} \sim -\frac{d}{\alpha} \operatorname{ten} t - \operatorname{ten} \Gamma\left(\frac{d}{\alpha} + 1\right) - \frac{t^{-d/\alpha + 1}}{2^{d/\alpha} \Gamma\left(\frac{d}{\alpha} + 1\right)} + \frac{t^{-pd/\alpha + 1}}{(p+1)^{\frac{d}{\alpha} - 1} \left[\Gamma\left(\frac{d}{\alpha} + 1\right)\right]^{p}}$$

$$(t + \infty, Vd/\alpha) \tag{28}$$

Note that, as expected, the high-temperature leading terms do not depend on p (Maxwell-Boltzmann limit).

# 3.2 - Internal energy, specific heat, equation of states and latent heat

The internal energy is given by

$$U = \int_{0}^{\infty} d\varepsilon \, \varepsilon \, \rho(\varepsilon) \, f(\varepsilon)$$
 (29)

hence

$$\frac{\mathbf{U}}{\mathbf{N}} = \mu_{\mathbf{O}}(\mathbf{1})\mathbf{u} \tag{30}$$

with

$$u(t) = \frac{d}{\alpha} t^{\frac{d}{\alpha} + 1} \int_{0}^{\infty} dy \ y^{\frac{d}{\alpha}} \left[ \frac{1}{e^{y - \overline{\mu}/t} - 1} - \frac{p+1}{e^{(p+1)(y - \overline{\mu}/t)} - 1} \right]$$
(31)

where we have used Eqs. (5), (9) and (11). This expression, to gether with  $\bar{\mu}(t)$  determined in Section 3.1, completely determines—the thermal dependence of u and consequently U/N. The specific heat C is given by

$$C = \frac{dU}{dT} \tag{32}$$

Therefore the specific heat c per particle is given by

$$c \equiv C/Nk_B = \frac{d(U/N)}{k_B^2 dT} = \frac{du}{dt}$$
 (33)

where we have used Eq. (30). The results are presented in Fig. 4.

Let us now discuss the low-temperature behavior of c. For finite p we can expand the integral of Eq. (31) as indicated in the Appendix, and then use Eqs. (26) and (27); we obtain

$$c \sim \frac{2\pi^2}{3} \frac{d}{\alpha} \frac{p^{d/\alpha}}{p+1} t \qquad (t \rightarrow 0)$$
 (34)

This is an interesting result, as it shows that asymptotically C = T for all finite values of p and  $d/\alpha$ , thus generalizing the well known result for quasi-free electrons in a metal  $(d=3,\alpha=2 \text{ and } p=1)$ .

In the  $p \rightarrow \infty$  limit , Eq. (31) becomes

$$u(t) = \frac{d}{\alpha} t^{\frac{d}{\alpha} + 1} \int_{0}^{\infty} dy \frac{y^{d/\alpha}}{e^{y} - 1} \qquad (t < t_c)$$
 (35)

where we have used that  $\tilde{\mu} \equiv 0$  if  $t < t_c$ . By using  $[\tilde{1}_0]$  (page 325) we can rewrite this equation as follows

$$u(t) = \frac{d}{\alpha} \Gamma(\frac{d}{\alpha} + 1) \zeta(\frac{d}{\alpha} + 1) t^{\frac{d}{\alpha} + 1}$$
(36)

hence

$$c = \frac{d}{\alpha} \Gamma\left(\frac{d}{\alpha} + 2\right) \xi\left(\frac{d}{\alpha} + 1\right) t^{d/\alpha}$$
 (37)

for the case  $d/\alpha > 1$ 

$$c = \frac{d}{\alpha} \left( \frac{d}{\alpha} + 1 \right) \frac{\zeta \left( \frac{d}{\alpha} + 1 \right)}{\zeta \left( \frac{d}{\alpha} \right)} \left( \frac{t}{t_c} \right)^{d/\alpha}$$
(37')

where we have used Eq. (221). The C  $\propto$  T<sup>d/ $\alpha$ </sup> law we have obtained generalizes the Debye law for acoustic phonons (d = 3,  $\alpha$  = 1) as well as the T<sup>3/2</sup> law for magnons in the standard Heisenberg ferromagnet (d = 3,  $\alpha$  = 2).

Let us now focus the high-temperature behavior of c. The integral of Eq. (31) can be treated in the same way we did for Eq. (17); we obtain

$$c \sim \frac{d}{\alpha} \left\{ 1 + \frac{d/\alpha - 1}{2^{d/\alpha + 1} \Gamma(d/\alpha + 1) t^{d/\alpha}} - \frac{p(pd/\alpha - 1)}{(p+1)^{d/\alpha} \left[ \frac{\pi}{L} (d/\alpha + 1) \right]^p t^{pd/\alpha}} \right\}$$

$$(t \to \infty) \tag{38}$$

Note that in the Fermi-Dirac (Bose-Einstein) statistics, the specific heat approaches the classical  $d/\alpha$  value from below (above) when  $d/\alpha \gg 1$ , and the opposite happens when  $d/\alpha \ll 1$ ; for  $d/\alpha = 1$ , the approach occurs from below for all values of p. Consequently C vs. T presents, for  $d/\alpha \gg 1$ , a maximum for p high enough (the maximum becomes a cusp in the p  $+\infty$  limit); for  $d/\alpha \ll 1$ , it presents a maximum for p low enough.

Let us now deduce the equation of states. The grand-canonical partition function  $\frac{1}{2k}$  associated with the wavevector k is given by

$$= 1 + e^{-\beta(\varepsilon_{k}^{2} - \mu)} + \dots + e^{-p\beta(\varepsilon_{k}^{2} - \mu)} = \frac{1 - e^{-(p+1)\beta(\varepsilon_{k}^{2} - \mu)}}{1 - e^{-\beta(\varepsilon_{k}^{2} - \mu)}}$$
 (39)

The total partition function  $\frac{\pi}{2}$  equals  $\frac{\pi}{k}$ , consequently

$$\ln \Xi = \int_{0}^{\infty} d\varepsilon \ \rho(\varepsilon) \ln \frac{1 - e^{-(p+1)\beta(\hat{\varepsilon} - \mu)}}{1 - e^{-\beta(\varepsilon - \mu)}}$$
(40)

We know from Thermodynamics that the pressure P satisfies

$$\beta PV = \ell n \equiv$$

$$\beta PV = \int_{0}^{\infty} d\varepsilon \ \rho(\varepsilon) \ln \frac{1 - e^{-(p+1)\beta(\varepsilon - \mu)}}{1 - e^{-\beta(\varepsilon - \mu)}}$$

$$= \frac{\beta^{-d/\alpha}}{\varepsilon^{d/\alpha}} \int_{0}^{\infty} dy \ y^{d/\alpha - 1} \ln \frac{1 - e^{-(p+1)(y - \beta \mu)}}{1 - e^{-(y - \beta \mu)}}$$
(41)

where we have used Eq. (5). On the other hand, also from Thermo dynamics, we have that

$$U = -\frac{\partial \ln^{2}}{\partial \beta} \Big|_{\text{fixed } \beta \mu} = -\frac{\partial (\beta PV)}{\partial \beta} \Big|_{\text{fixed } \beta \mu}$$
$$= \frac{d}{\alpha} \frac{\beta^{-d/\alpha - 1}}{\epsilon_{0}^{d/\alpha}} \int_{0}^{\infty} dy \, y^{d/\alpha - 1} \ln \frac{1 - e^{-(p+1)(y - \beta \mu)}}{1 - e^{-(y - \beta \mu)}}.$$

And by using again Eq. (43) we obtain

$$P = \frac{\alpha}{d} \frac{U}{V}$$
 (42)

which is exact for all values of p. Eq. (42) transforms the analysis of the pressure into that of U, which we have already done.

Let us finally calculate the latent-heat L (per particle) associated with the Bose-Einstein condensation (first order) phase transition. The Clapeyron equation states that

$$L = T \left(\frac{dP}{dT}\right)_{T_c} (v_n - v_c) \tag{43}$$

where  $v_n(v_c)$  is the volume per particle in the normal(condensed) phase. But in the N >> 1 limit  $v_c$  vanishes 1 and  $v_n$  equals V/N, therefore

$$L = T \left(\frac{dP}{dT}\right)_{T_{c}} \frac{V}{N}$$

$$= T \frac{\alpha}{d} \left[\frac{d(U/N)}{dT}\right]_{T_{c}} = k_{B}T \frac{\alpha}{d} \left(\frac{du}{dt}\right)_{t_{c}}$$

where we have used Eq. (42) and the definitions of u and t. By replacing Eq. (36) in this expression we obtain

$$L = k_B T \Gamma \left(\frac{d}{\alpha} + 1\right) \zeta \left(\frac{d}{\alpha} + 1\right) \left(\frac{d}{\alpha} + 1\right) t_c^{d/\alpha}$$

and, by using Eq. (22'), we finally obtain

$$\frac{L}{k_B^T} = \frac{\zeta(\frac{d}{\alpha} + 1)}{\zeta(\frac{d}{\alpha})} \left(\frac{d}{\alpha} + 1\right)$$
(44)

$$\begin{array}{c}
\frac{1}{2}(d/\alpha-1) & \text{if } d/\alpha-1 \rightarrow +0 \\
d/\alpha+1 & \text{if } d/\alpha \rightarrow \infty
\end{array}$$
(44<sup>+</sup>):

The dependence of L on  $d/\alpha$  is represented in Fig. 5.

# 4 UNFIXED TOTAL NUMBER OF PARTICLES (QUASI-PARTICLES)

In the present situation N is unfixed, therefore  $\mu$  vanishes for all temperatures. The average number <N> of particles is given by

$$\langle N \rangle = \int_{0}^{\infty} d \varepsilon \rho(\varepsilon) \left[ \frac{1}{e^{\beta \varepsilon} - 1} - \frac{p+1}{e^{(p+1)\beta \varepsilon} - 1} \right]$$
 (45)

hence

$$\frac{\langle N \rangle}{V} = \frac{1}{V} \left( \frac{k_B T}{\epsilon_0} \right)^{d/\alpha} \int_0^\infty dy \, y^{d/\alpha - 1} \left[ \frac{1}{e^{y} - 1} - \frac{p + 1}{e^{(p+1)y} - 1} \right]$$

where we have used Eq. (5). It follows that for  $d/\alpha > 1$ 

$$\frac{\langle N \rangle}{V} = \frac{\zeta(d/\alpha) \Gamma(d/\alpha+1)}{2^{d} \pi^{d/2} \Gamma(d/2+1) \binom{d}{\pi} a_{j}^{1/2}} \left[ 1 - \frac{1}{(p+1)^{d/\alpha-1}} \right] (k_{B}T)^{d/\alpha}$$
(46)

where we have used Eq. (41).

Let us now focus the internal energy. It is given by

$$U = \int_{0}^{\infty} d\varepsilon \, \varepsilon \, \rho \, (\varepsilon) \left[ \frac{1}{e^{\beta \varepsilon} - 1} - \frac{p+1}{e^{(p+1)\beta \varepsilon} - 1} \right] \tag{47}$$

hence

$$\frac{\mathbf{U}}{\mathbf{V}} = \frac{1}{\mathbf{V}} \frac{\left(\mathbf{k_B T}\right)^{d/\alpha + 1}}{\varepsilon_{\alpha}^{d/\alpha}} \int_{0}^{\infty} d\mathbf{y} \, \mathbf{y}^{d/\alpha} \left[ \frac{1}{e^{\mathbf{y}} - 1} - \frac{\mathbf{p} + 1}{e^{(\mathbf{p} + 1)\mathbf{y}} - 1} \right]$$

hence

$$\frac{U}{V} = \frac{\zeta(d/\alpha+1) \Gamma(d/\alpha+1) d/\alpha}{2^{d} \pi^{d/2} \Gamma(d/2+1) \binom{d}{j} \frac{1}{2} a_{j}^{1/2}} \left[1 - \frac{1}{(p+1)^{d/\alpha-1}}\right] (k_{B}T)^{d/\alpha+1}$$
(48)

where we have used again Eq. (4'). The present  $U/V \propto T$  law generalizes the well known black-body  $T^4$  Stefan-Boltzmann law (d=: 3,  $\alpha=1$ ,  $p\to\infty$ ). The generalized radiation pressure can be obtained by replacing Eq. (48) into Eq. (42) which still holds.

#### CONCLUSION

In an unified framework, we treat the parastatistics of an ideal confined (fixed volume) d-dimensional gas of N particles (or quasi-particles) whose energy spectrum is given by Eq. (1) (which contains the important isotropic case  $\varepsilon \propto |\vec{k}|^{\alpha}(\alpha > 0)$ ); each state can be occupied at most by p particles. We have studied, for arbitrary d,  $\alpha$ , p, T, and for both fixed and unfixed N cases, various thermodynamic quantities (chemical potential, average populations, internal energy, specific heat, pressure and latent heat). In what follows we summarize the main results.

- (i) The density of states satisfies  $\rho(\epsilon) \propto \epsilon^{d/\alpha+1}$ . This form will imply that the thermal dependence of all equilibrium thermodynamic quantities, and for all p, will depend on d and  $\alpha$  on through the ratio  $d/\alpha$ .
- (ii) The Bose-Einstein condensation appears, for  $d/\alpha > 1$ , as a non uniform convergence in the  $p \to \infty$  limit, in complete analogy with phase transitions in interacting systems which appear as non uniform convergences in the thermodynamic limit  $N \to \infty$ . We obtain the explicit dependence of the critical temperature  $T_c$  on  $d/\alpha$  (Eq. (22')): it is  $T_c \propto (d/\alpha 1)$  for  $d/\alpha \gtrsim 1$ , and  $T_c \propto 5 \alpha/d$  for  $d/\alpha >> 1$ . The macroscopic population of the ground state satisfies  $N_0/N = 1 (T/T_c)^{d/\alpha}$  for  $T \le T_c$  and  $d/\alpha \times 1$ . The latent heat L is obtained as an explicit function of  $d/\alpha$  (Eq. 44)); it satisfies  $L/T \propto 1/2(d/\alpha 1)$  for  $d/\alpha \gtrsim 1$ , and  $L/T \sim d/\alpha + 1$  for  $d/\alpha \gg 1$ . The picture which emerges (here and in the paragraphs (iv) and (v) which follow) is that the  $p \to \infty$  limit is deeply different from any

other case (i.e.  $0 ). Moroever, all finite p paragases are qualitatively similar, and are well-characterized by the Fermi-Dirac case (p = 1). Within this context, the view developed in Ref. [2], where the d = 3, <math>\alpha$  = 2 Bose-Einstein condensation is referred for all values of p, can be considered as deeply misleading.

- (iii) For fixed N, finite p and low temperature, the chemical potential presents a quadratic departure from the T=0 value (Eq. (26)); the curvature changes its sign at  $d/\alpha=1$ . In the high temperature regime the first two (or three for p>1) dominant terms (of the chemical potential) do not depend on p.
  - (iv) For fixed N, finite p, arbitrary  $d/\alpha$  and low temperature, the specific heat satisfies  $C \propto T$  thus generalizing the standard result for quasi-free electrons in a conductor  $(d=3,\alpha=2,p=1)$ . In the  $p+\infty$  limit we obtain, for both fixed and unfixed N cases,  $C \propto T^{d/\alpha}$  thus generalizing the Debye-law for acoustic phonons in a crystal  $(d=3,\alpha=1)$  and the  $T^{3/2}$ -law for magnons in a Heisenberg ferromagnet  $(d=3,\alpha=2)$ .
- (v) For fixed N, arbitrary p and  $d/\alpha$ , and high temperature, C approaches the classical value  $(d/\alpha)Nk_B$ . For p low (high) enough, C approaches this value from below (above) when  $d/\alpha > 1$ , and the opposite happens when  $d/\alpha < 1$ ; consequently C presents, for  $d/\alpha > 1$ , a maximum for p high enough (cusp in the p+ $\infty$  limit) and presents, for  $d/\alpha < 1$ , a maximum for p low enough.

- (vi) For arbitrary p, d/ $\alpha$ , T and for both fixed and unfixed N cases, the pressure P and the internal energy U are related through U = (d/ $\alpha$ )PV (V = volume), thus generalizing the standard result U = (3/2)PV for d = 3,  $\alpha$  = 2 and p = 1, $\infty$ .
- (vii) For unfixed N, arbitrary p,d/ $\alpha$  and T, the density of the internal energy (proportional to the power irradiated per unit area of the confining box) satisfies U/V  $\propto$  T<sup>d/ $\alpha$ +1</sup> (Eq. (48)), thus generalizing the T<sup>4</sup> Stefan-Boltzmann law (d = 3,  $\alpha$  = 1, p  $\rightarrow \infty$ ).

Also, since the 1968 paper by Gunton and Buckingham [9] (see also [12] and references therein), the Bose-Einstein condensation is known to be related to the criticality of the spherical model (n-vector model with  $n \to \infty$ ). Consistently, the fact that  $T_c$  should vanish in the  $d/\alpha \to 1$  limit can be inferred from this relationship.

Finally, as a closing remark let us recall that several studies  $\begin{bmatrix} -13-16 \end{bmatrix}$  are already available in which the bosonic limit  $(p \to \infty)$  is analysed under a variety of conditions (for both spinless and magnetized bosons). By following along the lines of the present paper, it could be interesting to extend those results to arbitrary values of p.

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## APPENDIX-GENERALIZED SOMMERFELD INTEGRAL

Here we generalize to an arbitrary finite p the standard Schmerfeld integral for the FD statistics and its low temperature expansion. We consider the integral

$$I(z) \equiv \int_{-\infty}^{\infty} dy H(y) f(y;z)$$
 (A.1)

where

$$f(y;z) \equiv \frac{1}{e^{y-z}-1} - \frac{p+1}{e^{(p+1)(y-z)}-1}$$
  $(p > 0)$ , (A.2)

To characterize the function H(y) we introduce

$$K(\lambda) \equiv \sum_{n=0}^{\infty} d\lambda_{n} H(\lambda_{n})$$
 (Y')

K(y) is assumed to satisfy the following requirements:

(i) 
$$\lim_{y \to -\infty} K(y) = 0 \qquad (A.4)$$

hence

$$H(y) = \frac{dK(y)}{dy} ; \qquad (A.5)$$

(ii) 
$$\lim_{y\to\infty} f(y;z)K(y) = 0 ; \qquad (A.6)$$

(iii) 
$$K(y)$$
 is analytic at  $y = z$ . (A.7)

These conditions are almost always satisfied in the physical

situations as typical functions H(y) identically vanish for y < 0, as symptotically behave as a power-law in the  $y \rightarrow \infty$  limit, and are soft functions at y = z (H(y) might present singularities, but they are normally integrable and located at places such as y = 0, and not at y = z).

Integrating (A.1) by parts we obtain

$$I(z) = K(y) f(y; z) \Big|_{-\infty}^{\infty} - \int_{-\infty}^{\infty} dy K(y) \frac{df(y; z)}{dy}$$
$$= - \int_{-\infty}^{\infty} dy K(y) \frac{df(y; z)}{dy}$$
(A.8)

where we have used conditions (A.4) and (A.6) and the fact that  $\lim_{y\to-\infty} f(y;z) = p$ . By expanding K(y) at y=z we obtain

$$I(z) = -K(z) \int_{-\infty}^{\infty} dy \frac{df}{dy} - \sum_{n=1}^{\infty} \left\{ \frac{1}{n!} \left[ \frac{d^n K(y)}{dy^n} \right]_{y=z} - \int_{-\infty}^{\infty} dy (y-z)^n \frac{df}{dy} \right\}$$

$$= pK(z) - \sum_{m=1}^{\infty} \left\{ \frac{1}{(2m)!} \left[ \frac{d^{2m} K(y)}{dy^{2m}} \right]_{y=z} - \sum_{-\infty}^{\infty} dy (y-z)^{2m} \frac{df}{dy} \right\}$$

$$(A.9)$$

where we have used the facts that

$$\int_{-\infty}^{\infty} dy \frac{df}{dy} = -p$$

and that df/dy is an even function of (y-z) and therefore all odd terms in the sum over in vanish. By using the connection between K(y) and H(y) we can rewrite Eq. (A.9) as follows:

$$I(z) = p \int_{-\infty}^{z} H(y) dy + \sum_{m=1}^{\infty} \left[ \frac{d^{2m-1}H(y)}{dy^{2m-1}} \right]_{y=z} a_{m}$$
(A.10)

with

$$a_{m} = \frac{1}{(2m)!} \int_{-\infty}^{\infty} dx \, x^{2m} \, \frac{d}{dx} \left[ \frac{p+1}{e^{(p+1)x} - 1} - \frac{1}{e^{x} - 1} \right]$$

$$= \frac{1}{(2m)!} \left[ 1 - \frac{1}{(p+1)^{2m-1}} \right] \int_{-\infty}^{\infty} dx \, x^{2m} \, \frac{d}{dx} \, \frac{1}{1 - e^{x}}$$

$$= \frac{1}{(2m)!} \left[ 1 - \frac{1}{(p+1)^{2m-1}} \right] 2^{2m} \int_{-\infty}^{\infty} dx \, \frac{x^{2m}}{\sinh^{2}x}$$

$$= \frac{1}{(2m)!} \left[ 1 - \frac{1}{(p+1)^{2m-1}} \right] (2\pi)^{2m} |B_{2m}| \qquad (A.11)$$

where in the last step we have used [16] (page 352),  $B_{2m}$  being the Bernoulli numbers. By using the fact that  $|B_{2m}| = (2m)! \chi(2m)/(2^{2m-1}\pi^{2m})$  we can finally express  $a_m$  as follows:

$$a_m = 2\zeta(2m)\left[1 - \frac{1}{(p+1)^{2m-1}}\right]$$
 (A.12)

and consequently

$$\int_{-\infty}^{\infty} dy \, H(y) \, f(y; z) = p \int_{-\infty}^{z} H(y) \, dy$$

$$+ 2 \int_{m=1}^{\infty} \left\{ \zeta(2m) \left[ 1 - \frac{1}{(p+1)^{2m-1}} \right] \left[ \frac{d^{2m-1}H(y)}{dy^{2m-1}} \right]_{y=z} \right\}$$

$$= p \int_{-\infty}^{z} H(y) \, dy + \frac{\pi^{2}}{3} \left[ 1 - \frac{1}{p+1} \right] H^{1}(z)$$

$$+ \frac{\pi^{4}}{45} \left[ 1 - \frac{1}{(p+1)^{3}} \right] H^{m_{1}}(z) + \dots \qquad (A.13)$$

which is the generalization we were looking for.

Quite frequently H(y) is defined as follows:

$$H(y) = \begin{cases} 0 & \text{if } y < 0 \\ \\ Ay^{\gamma} & \text{if } y \ge 0 \end{cases}$$
 (A.14)

In such case, Eq. (A.13) can be rewritten as follows:

$$\int_{0}^{\infty} dy \, H(y) \, f(y; z)$$

$$= A z^{\gamma+1} \left\{ \frac{p}{\gamma+1} + 2 \sum_{m=1}^{\infty} \frac{\zeta(2m)}{z^{2m}} \left[ 1 - \frac{1}{(p+1)^{2m-1}} \right]_{j=0}^{2(m-1)} (\gamma + j) \right\} (A.15)$$

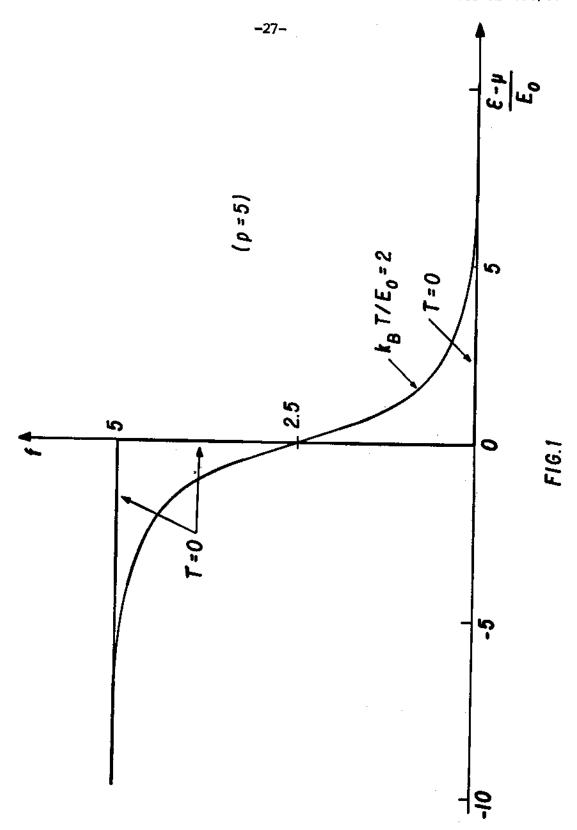
Finally if we identify  $y = \epsilon/k_BT$  and  $z = \mu/k_BT$ , Eq. (A.13) takes the form

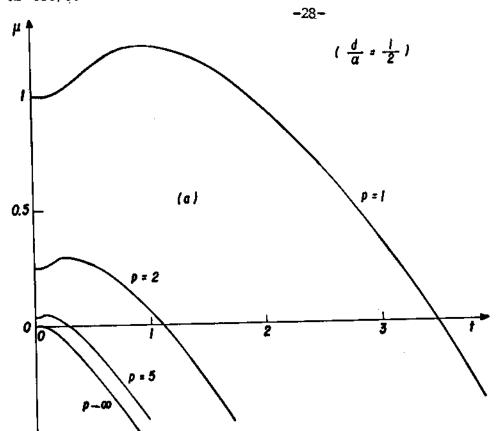
$$\int_{-\infty}^{\infty} d\varepsilon H(\varepsilon) f(\varepsilon) = p \int_{-\infty}^{\mu} H(\varepsilon) d\varepsilon 
+ 2 \sum_{m=1}^{\infty} \left\{ \zeta(2m) \left[ 1 - \frac{1}{(p+1)^{2m-1}} \right] \left[ \mu^{2m} H^{(2m-1)}(\mu) \right] \left( \frac{k_B T}{\mu} \right)^{2m} \right\} 
= p \int_{-\infty}^{\mu} H(\varepsilon) d\varepsilon + \frac{\pi^2}{3} \left[ 1 - \frac{1}{p+1} \right] \left[ \mu^2 H^{T}(\mu) \right] \left( \frac{k_B T}{\mu} \right)^2 
+ \frac{\pi^4}{45} \left[ 1 - \frac{1}{(p+1)^3} \right] \left[ \mu^4 H^{T}(\mu) \right] \left( \frac{k_B T}{\mu} \right)^4 + o \left[ \left( \frac{k_B T}{\mu} \right)^6 \right]$$
(A.16)

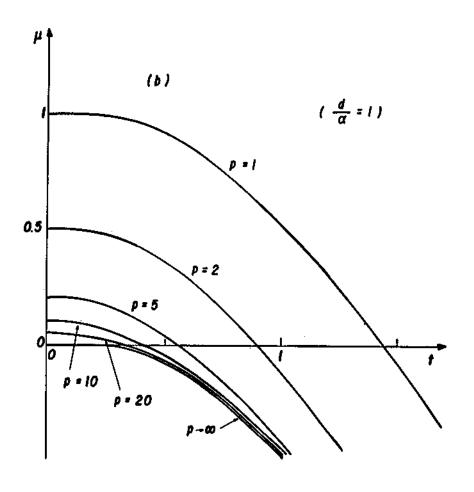
where in the estimation of the rest o  $\left[\left(k_BT/\mu\right)^6\right]$  we have assumed the quite frequent fact that  $H^{(n)}\left(\mu\right)$  is of the order of  $H(\mu)/\mu^n$ .

### CAPTION FOR FIGURES

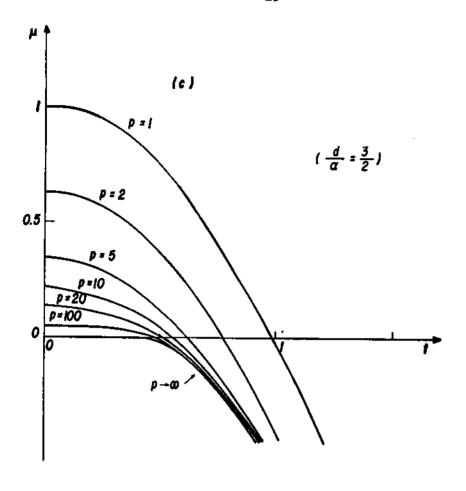
- Fig. 1 Parastatistics with maximal occupation p: average population per state as a function of the energy  $\epsilon$  measured from the chemical potential  $\mu$ , for finite and vanishing temperatures. The point  $(\epsilon-\mu,f)=(0,p/2)$  is a center of symmetry.
- Fig. 2 Thermal dependence of the (reduced) chemical potential for typical values of p and d/a. Notice that for all  $p < \infty$  the behavior is qualitatively the same (in the sense that  $\overline{\mu}$  is analytic for all finite values of t) no matter the value of d/a, and can be characterized by the Fermi-Dirac (p = 1) behavior, wheras the p +  $\infty$  limit (Bose Einstein) presents two different regimes, one occurring for d/a  $\leq$  1 (no Bose-Einstein condensation), the other one occurring for d/a  $\geq$ 1 (Bose-Einstein condensation due to non-analyticity of  $\overline{\mu}$  appearing at a finite temperature).
- Fig. 3 Reduced temperature at which  $\mu$  vanishes as a function of d/ $\alpha$  and p. The p +  $\infty$  limit yields the d/ $\alpha$ -dependence of the (reduced) critical temperature corresponding to the Bose-Einstein condensation.
- Fig. 4 Thermal dependence of the (reduced) specific heat for typical values of p\_and d/α (a-e). For the p →∞ limit we have represented in (f) the d/α-dependence of the height of the cusp which appears in c vs. t for d/α >1.
- Fig. 5 The d/ $\alpha$ -dependence of the latent heat L per particle (in units of  $k_B^{\rm T}$ ) associated with the Bose-Einstein condensation (first-order phase transition).

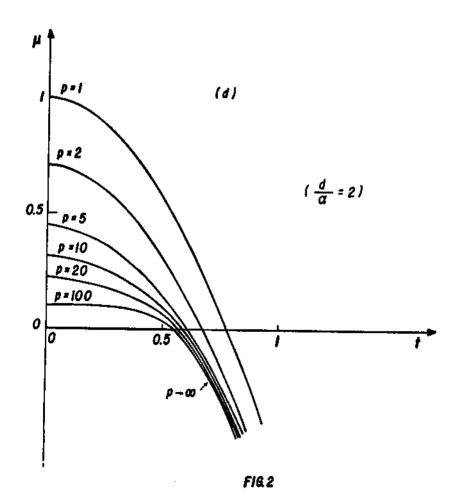


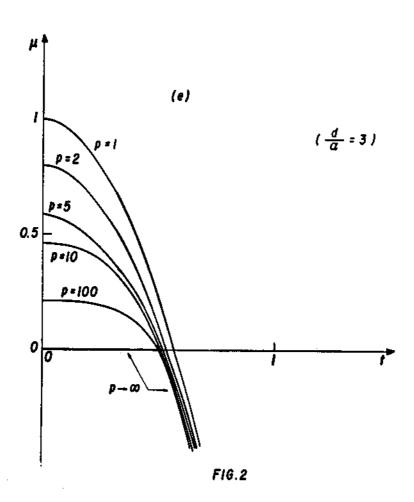


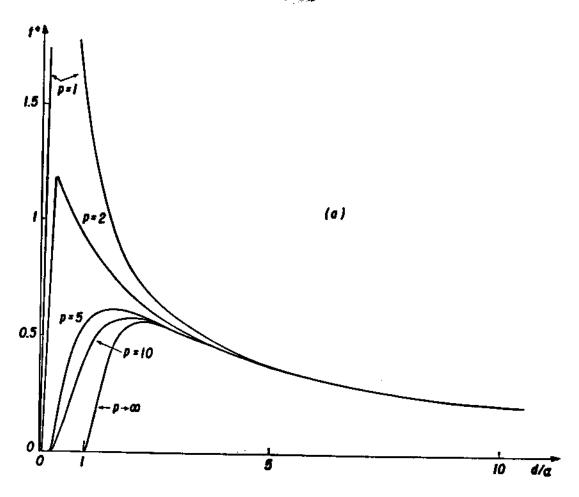


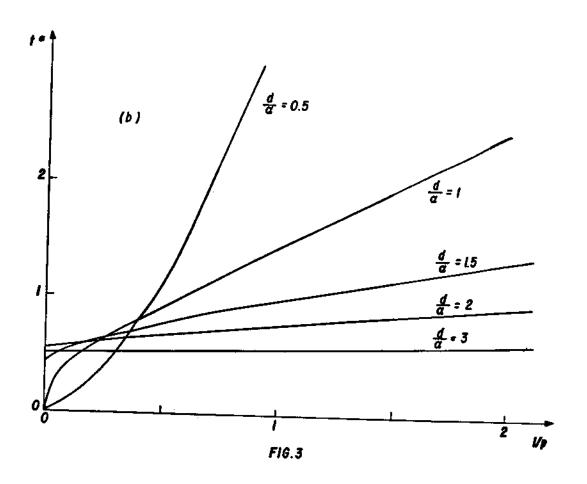
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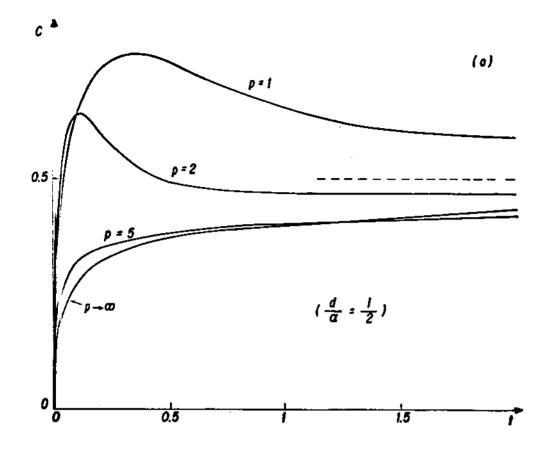


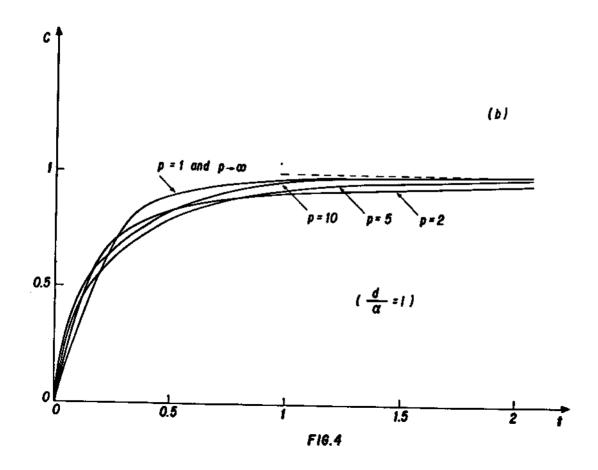


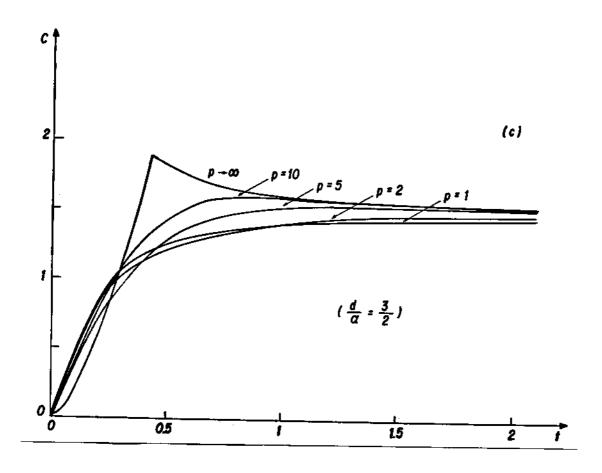


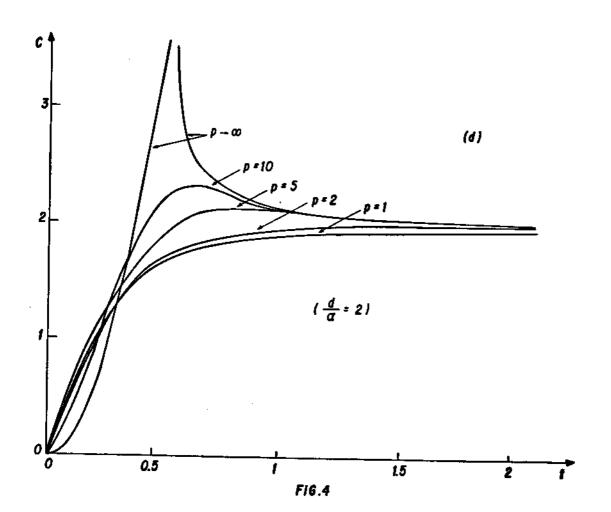


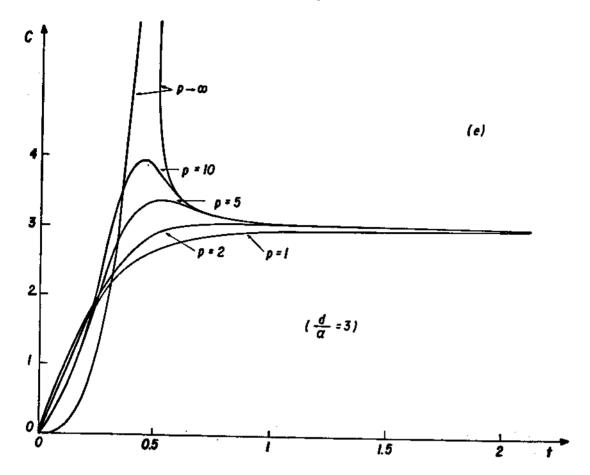


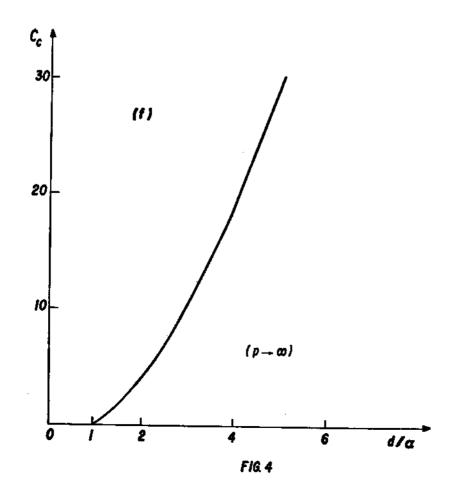


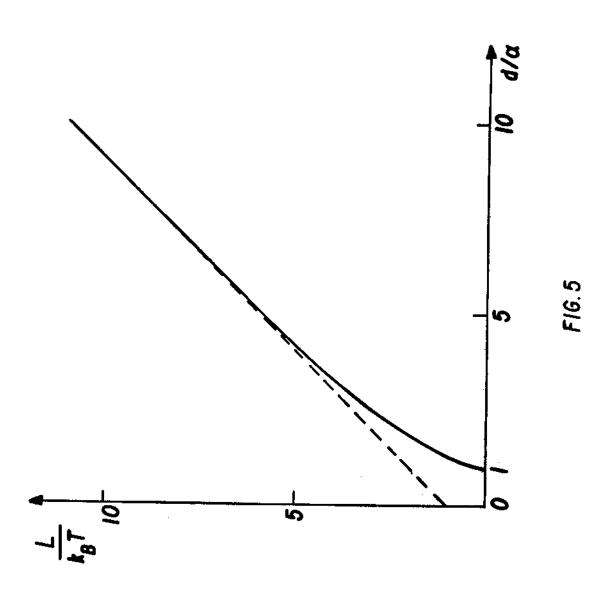












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